Claire Rigouzzo¹

¹Theoretical Particle Physics and Cosmology, King's College London, WC2R 2LS, UK

Claire Rigouzzo,^a Sebastian Zell^a

^a Institue of Physics, Laboratory for Particle Physics and Cosmology, École Polytechnique Fédérale de Lausanne, CH-1015 Lausanne, Switzerland

Claire Rigouzzo,^a Sebastian Zell^a

^a Institue of Physics, Laboratory for Particle Physics and Cosmology, École Polytechnique Fédérale de Lausanne, CH-1015 Lausanne, Switzerland

1. Different Formulations of Gravity

Claire Rigouzzo,^{*a*} Sebastian Zell^{*a*}

^a Institue of Physics, Laboratory for Particle Physics and Cosmology, École Polytechnique Fédérale de Lausanne, CH-1015 Lausanne, Switzerland

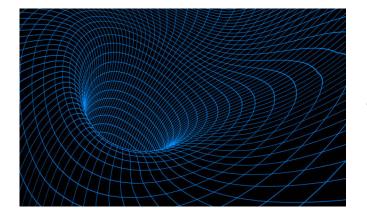
- 1. Different Formulations of Gravity
- 2. Breaking the Equivalence Between the Different Formulations

Claire Rigouzzo,^a Sebastian Zell^a

^a Institue of Physics, Laboratory for Particle Physics and Cosmology, École Polytechnique Fédérale de Lausanne, CH-1015 Lausanne, Switzerland

- 1. Different Formulations of Gravity
- 2. Breaking the Equivalence Between the Different Formulations
- 3. Phenomenology of Higgs Inflation

A. Metric Gravity



Degrees of freedom: $g_{\mu\nu}$

The connection is **uniquely** determined by the metric:

$$\mathring{\Gamma}^{\alpha}_{\ \beta\gamma} = \frac{1}{2} g^{\alpha\mu} \left(\partial_{\beta} g_{\mu\gamma} + \partial_{\gamma} g_{\mu\beta} - \partial_{\mu} g_{\beta\gamma} \right)$$

B. Palatini Gravity

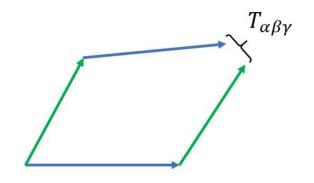


Degrees of freedom: $\{g_{\mu\nu}, \Gamma^{\alpha}_{(\beta\gamma)}\}$

The connection is no longer determined by the metric,

they are **a priori** independent.

C. Einstein-Cartan Gravity



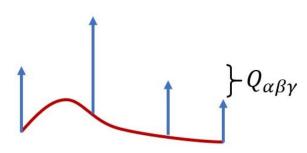
Degrees of freedom: $\{g_{\mu\nu}, \Gamma^{\alpha}_{\beta\gamma}\}$

 $\Gamma^{\alpha}_{\ \beta\gamma}$ need not be symmetric in the last indices

 \Rightarrow Torsion:

$$T^{\alpha}_{\ \beta\gamma} = \Gamma^{\alpha}_{\ [\beta\gamma]}$$

D. Metric-Affine Gravity



Degrees of freedom: $\{g_{\mu\nu}, \Gamma^{\alpha}_{\beta\gamma}\}$

Most general formulation of gravity

 \Rightarrow Non-metricity:

$$Q_{\alpha\beta\gamma} = \nabla_{\alpha}g_{\beta\gamma}$$

E. Summary

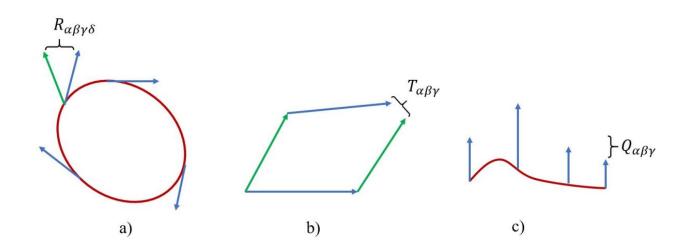


Figure 1: Schematic representation of the change of a vector under parallel transport due to the presence of: a) curvature b) torsion c) non-metricity.

E. Summary

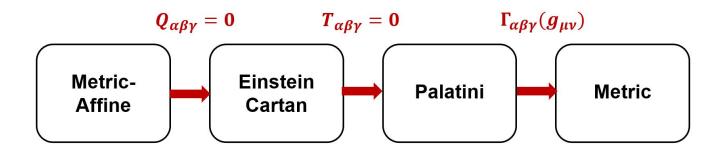


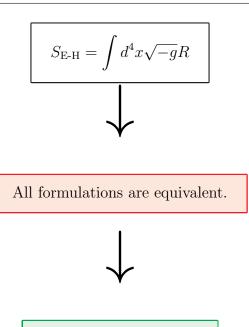
Figure 2: Relation between the different formulations of gravity.

Are they equivalent?

If not, what are the phenomenological consequences? Can we measure it?

$$S_{\text{E-H}} = \int d^4x \sqrt{-g}R$$

All formulations are equivalent.



So, Why do we care?...

A. Higher derivatives

$$S = \int d^4x \sqrt{-g} [R + a_1 R^2 + a_2 R^{\mu\nu} R_{\mu\nu} + \dots$$

Schematically, the equivalence is broken due to symmetry properties of the Riemann tensor :

A. Higher derivatives

	Metric	Einstein-Cartan	Metric-affine
$R_{ab[cd]}$	\checkmark	\checkmark	\checkmark
$R_{[ab]cd}$	\checkmark	\checkmark	Х
$R_{(ab)(cd)}$	\checkmark	Х	Х
$R_{a[bcd]} = 0$	\checkmark	Х	Х

Table I: Properties of Riemann tensor

e.g: In the metric-affine formalism, we can write a term like :

$$S = \int d^4x \sqrt{-g} R^{\alpha}_{\ \alpha\beta\gamma} R^{\ \mu\beta\gamma}_{\mu}$$

A. Higher derivatives

$$S = \int d^4x \sqrt{-g} [R + a_1 R^2 + a_2 R^{\mu\nu} R_{\mu\nu} + \dots$$

Few remarks:

- (a) New propagating d.o.fs \Rightarrow quite a big deviation from GR.
- (b) Some may be healthy, some unhealthy (ghosts or tachyons) [1]

B. Matter coupled to gravity

$$S = \int d^4x \sqrt{-g} (1 + \xi \phi^2) R + S_\phi$$

Few remarks:

- (a) Non-minimal coupling terms come naturally when considering renormalization properties of a scalar field in a curved spacetime background [2].
- (b) Gravity sector stays the same \Rightarrow no new *propagating* d.o.fs \Rightarrow minimal deformation to GR.

[2] N. D. Birrell and P. C. W. Davies, "Quantum Fields in Curved Space". Cambridge Univ. Press, (1984)

B. Matter coupled to gravity

Conclusion: different formulations are no longer equivalent when the action is more complicated.

A. Motivation

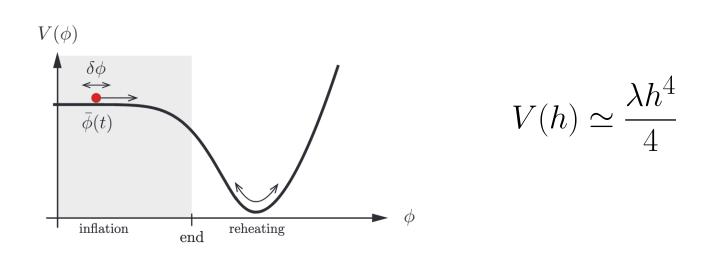
- Matter field \Rightarrow different formulations are no longer equivalent.
- Choose the most general formulation, i.e metric-affine.

Presence of curvature, torsion and non-metricity

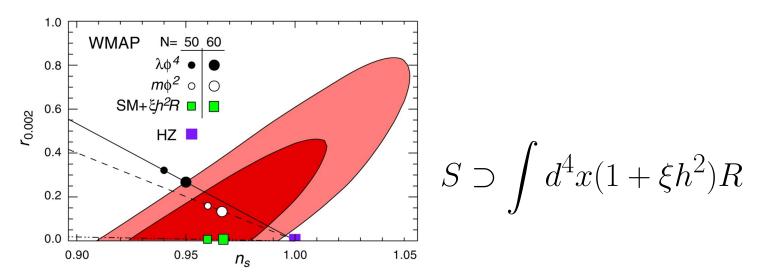
- Torsion and non-metricity are non-dynamical (no kinetic terms).
- They correspond to high energy effects.

Inflation

B. Recap of Higgs inflation



B. Recap of Higgs inflation



[3] F. Bezrukov, M. Shaposhnikov, The Standard Model Higgs boson as the inflaton, 0710.3755

C. The theory, and its consequences

Steps followed in the paper:

1. Write down the most general action including torsion and non-metricity.

C. The theory, and its consequences

- 1. Write down the most general action including torsion and non-metricity.
- 2. Find solution for torsion $T_{\alpha\beta\gamma}$ and non-metricity $Q_{\alpha\beta\gamma}$.

C. The theory, and its consequences

- 1. Write down the most general action including torsion and non-metricity.
- 2. Find solution for torsion $T_{\alpha\beta\gamma}$ and non-metricity $Q_{\alpha\beta\gamma}$.
- 3. Plug them back in the action

C. The theory, and its consequences

- 1. Write down the most general action including torsion and non-metricity.
- 2. Find solution for torsion $T_{\alpha\beta\gamma}$ and non-metricity $Q_{\alpha\beta\gamma}$.
- 3. Plug them back in the action
- 4. Perform a conformal transformation of the metric to get rid of the non-minimal coupling

C. The theory, and its consequences

- 1. Write down the most general action including torsion and non-metricity.
- 2. Find solution for torsion $T_{\alpha\beta\gamma}$ and non-metricity $Q_{\alpha\beta\gamma}$.
- 3. Plug them back in the action
- 4. Perform a conformal transformation of the metric to get rid of the non-minimal coupling

$$S = \int \mathrm{d}^4x \sqrt{-g} \left[\frac{1}{2}R - \frac{1}{2}K(h)\partial_\alpha h\partial^\alpha h - \frac{\lambda h^4}{(1+\xi h^2)^2}\right].$$

C. The theory, and its consequences

• modified kinetic term for the Higgs field:

$$K(h) = \frac{1}{(1+\xi h^2)} \left[1 + \frac{h^2}{(\sum_{m=0}^4 O_m h^{2m})^2} \sum_{n=0}^7 P_n h^{2n} + \frac{6\xi^2 h^2}{(1+\xi h^2)} \right] .$$

• modified potential:

$$V(h) = \frac{\lambda h^4}{(1+\xi h^2)^2}$$

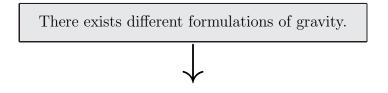
C. The theory, and its consequences

Steps followed in the paper:

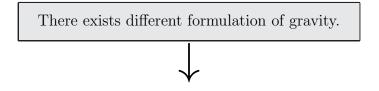
- 1. Write down the most general action including torsion and non-metricity.
- 2. Find solution for torsion $T_{\alpha\beta\gamma}$ and non-metricity $Q_{\alpha\beta\gamma}$.
- 3. Plug them back in the action
- 4. Perform a conformal transformation of the metric to get rid of the non-minimal coupling

Conclusion: Flattened potential and new higher mass dimension self-interaction terms for the Higgs

There exists different formulations of gravity.



Different formulations are no longer equivalent when the action is more complicated.



Different formulations are no longer equivalent when the action is more complicated.

Lead to different predictions at high energy.

New phenomenology at high energy:

- Flatter potential and new self-interactions for the Higgs field. [this paper]
- Production of Dark Matter through fermions coupled to gravity [4][to appear...].
- Different behaviour for singularities inside black holes [5]

[4] M. Shaposhnikov, A.Shrekin, I.Timiryasov and S.Zell, Einstein-Cartan Portal to Dark Matter, 2008.11686

[5] J.A.R.Cembranos, J. Gigante Valcarcel, and F.J. Maldonado Torralba, Singularities and n-dimensional black holes in torsion theories, 1609.07814

V. BIBLIOGRAPHY

[0] Template for the slides: D. Backhouse.

 [1] Lin, Y., Hobson, M.P., Lasenby, A.N. (2020). Power-counting renormalizable, ghost-and-tachyon-free Poincare gauge theories. *Physical Review D*

[2] N. D. Birrell and P. C. W. Davies, "Quantum Fields in Curved Space". Cambridge Univ. Press, (1984)

[3] F. Bezrukov, M. Shaposhnikov, The Standard Model Higgs boson as the inflaton, 0710.3755

[4] M. Shaposhnikov, A.Shrekin, I.Timiryasov and S.Zell, Einstein-Cartan Portal to Dark Matter, 2008.11686

[5] J.A.R.Cembranos, J. Gigante Valcarcel, and F.J. Maldonado Torralba, Singularities and n-dimensional black holes in torsion theories, 1609.07814

Further Reading:

[6] Beltran Jimenez, J., Heisenberg, L., Koivisto, T. S. (2019). The geometrical trinity of gravity. Universe.
[7] Karananas, G. K., Shaposhnikov, M., Shkerin, A., Zell, S. (2021). Matter matters in Einstein-Cartan gravity. Physical Review D, 104(6), 064036.

A. The action

$$S = \int d^4x \sqrt{-g} \Big[\frac{1}{2} (1 + \xi h^2) \mathring{R} - \frac{1}{2} \tilde{K}(h) g^{\alpha\beta} \partial_\alpha h \partial_\beta h - V(h) + A_1(h) \mathring{\nabla}_\alpha \hat{T}^\alpha + A_2(h) \mathring{\nabla}_\alpha T^\alpha + A_3(h) \mathring{\nabla}_\alpha \hat{Q}^\alpha + A_4(h) \mathring{\nabla}_\alpha Q^\alpha + B_1(h) Q_\alpha Q^\alpha + B_2(h) \hat{Q}_\alpha \hat{Q}^\alpha + B_3(h) Q_\alpha \hat{Q}^\alpha + B_4(h) q_{\alpha\beta\gamma} q^{\alpha\beta\gamma} + B_5(h) q_{\alpha\beta\gamma} q^{\beta\alpha\gamma} + C_1(h) T_\alpha T^\alpha + C_2(h) \hat{T}_\alpha \hat{T}^\alpha + C_3(h) T_\alpha \hat{T}^\alpha + C_4(h) t_{\alpha\beta\gamma} t^{\alpha\beta\gamma} + D_1(h) \epsilon_{\alpha\beta\gamma\delta} t^{\alpha\beta\lambda} t^{\gamma\delta}_{\ \lambda} + D_2(h) \epsilon_{\alpha\beta\gamma\delta} q^{\alpha\beta\lambda} q^{\gamma\delta}_{\ \lambda} + D_3(h) \epsilon_{\alpha\beta\gamma\delta} q^{\alpha\beta\lambda} t^{\gamma\delta}_{\ \lambda} + E_1(h) T_\alpha Q^\alpha + E_2(h) \hat{T}_\alpha Q^\alpha + E_3(h) T_\alpha \hat{Q}^\alpha + E_4(h) \hat{T}_\alpha \hat{Q}^\alpha + E_5(h) t^{\alpha\beta\gamma} q_{\beta\alpha\gamma} \Big].$$

A. The action

Since torsion and non-metricity each carry three tensor indices, it is convenient to split them further into vector- and pure tensor-parts. This is done by contracting all possible indices following the symmetry properties. For torsion, this gives:

the trace vector:
$$T^{\alpha} = g_{\mu\nu}T^{\mu\alpha\nu}$$
, (30)

the pseudo trace axial vector:
$$\hat{T}^{\alpha} = \epsilon^{\alpha\beta\mu\nu}T_{\beta\mu\nu}$$
, (31)

the pure tensor part:
$$t^{\alpha\beta\gamma}$$
 that satisfies $g_{\mu\nu}t^{\mu\alpha\nu} = 0 = \epsilon^{\alpha\beta\mu\nu}t_{\beta\mu\nu}$. (32)

Torsion can be be reconstructed in terms of these irreducible pieces as:

$$T_{\alpha\beta\gamma} = -\frac{2}{3}g_{\alpha[\beta}T_{\gamma]} + \frac{1}{6}\epsilon_{\alpha\beta\gamma\nu}\hat{T}^{\nu} + t_{\alpha\beta\gamma} .$$
(33)

Similarly, we can split further non-metricity into three contributions:

a first vector:
$$Q^{\gamma} = g_{\alpha\beta} Q^{\gamma\alpha\beta}$$
, (34)

a second vector:
$$\hat{Q}^{\gamma} = g_{\alpha\beta} Q^{\alpha\gamma\beta}$$
, (35)

the pure tensor part:
$$q^{\alpha\beta\gamma}$$
 that satisfies $g_{\alpha\beta}q^{\gamma\alpha\beta} = 0 = g_{\alpha\beta}q^{\alpha\gamma\beta}$. (36)

In terms of the components of (34) to (36), non-metricity can expressed as:

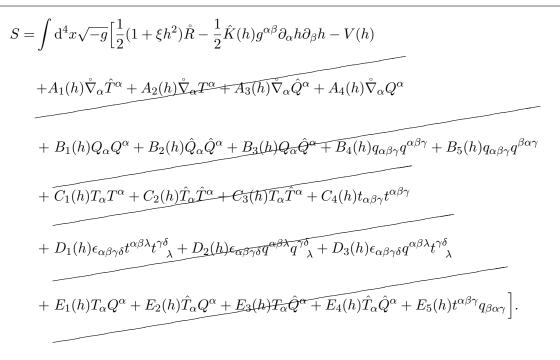
$$Q_{\alpha\beta\gamma} = \frac{1}{18} [g_{\beta\gamma} (5Q_{\alpha} - 2\hat{Q}_{\alpha}) + 2g_{\alpha(\beta} (4\hat{Q}_{\gamma)} - Q_{\gamma)})] + q_{\alpha\beta\gamma} .$$

$$(37)$$

B. Finding Solutions

$$Q^{\alpha} = \frac{V}{Z} \partial^{\alpha} h , \qquad \hat{Q}^{\alpha} = \frac{W}{Z} \partial^{\alpha} h , \qquad T^{\alpha} = \frac{X}{Z} \partial^{\alpha} h , \qquad \hat{T}^{\alpha} = \frac{Y}{Z} \partial^{\alpha} h , \qquad t_{\alpha\beta\gamma} = q_{\alpha\beta\gamma} = 0 . \tag{38}$$
And the common denominator reads
$$Z = B_3^2 (4C_1C_2 - C_3^2) + 4B_2C_2E_1^2 - 4B_2C_3E_1E_2 + 4B_2C_1E_2^2 - E_2^2E_3^2 + 2E_1E_2E_3E_4 \\ - E_1^2E_4^2 + B_3(-4C_2E_1E_3 + 2C_3E_2E_3 + 2C_3E_1E_4 - 4C_1E_2E_4) + 4B_1(B_2(-4C_1C_2 + C_3^2) \\ + C_2E_3^2 - C_3E_3E_4 + C_1E_4^2) , \qquad (39)$$

C. Equivalent Metric Theory



D. Decomposition of the scalar curvature

$$R = \mathring{R} + \mathring{\nabla}_{\alpha}(Q^{\alpha} - \hat{Q}^{\alpha} + 2T^{\alpha}) - \frac{2}{3}T_{\alpha}(T^{\alpha} + Q^{\alpha} - \hat{Q}^{\alpha}) + \frac{1}{24}\hat{T}^{\alpha}\hat{T}_{\alpha} + \frac{1}{2}t^{\alpha\beta\gamma}t_{\alpha\beta\gamma} - \frac{11}{72}Q_{\alpha}Q^{\alpha} + \frac{1}{18}\hat{Q}_{\alpha}\hat{Q}^{\alpha} + \frac{2}{9}Q_{\alpha}\hat{Q}^{\alpha} + \frac{1}{4}q_{\alpha\beta\gamma}(q^{\alpha\beta\gamma} - 2q^{\gamma\alpha\beta}) + t_{\alpha\beta\gamma}q^{\beta\alpha\gamma},$$

E. Decomposition of the Holst term

$$\epsilon^{\mu\nu\rho\sigma}R_{\mu\nu\rho\sigma} = \frac{1}{3}\hat{Q}^{\alpha}\hat{T}_{\alpha} - \frac{1}{3}Q^{\alpha}\hat{T}_{\alpha} - \frac{2}{3}\hat{T}^{\alpha}\hat{T}_{\alpha} + \mathring{\nabla}_{\alpha}T^{\alpha} - \frac{1}{2}\epsilon_{\beta\gamma\delta\mu}t_{\alpha}^{\ \delta\mu}t^{\alpha\beta\gamma} - \epsilon_{\alpha\gamma\delta\mu}q^{\alpha\beta\gamma}t_{\beta}^{\ \delta\mu}$$