

Carrollian and Galilean limits of deformed symmetries in 3D gravity

Tomasz Trzeźniewski*

Institute of Theoretical Physics,
University of Wrocław, Poland

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***T. T.**, arXiv:2306.05409 [hep-th]
Kowalski-Glikman, Lukierski & **T. T.**, JHEP **09**, 096 (2020)

Outline:

- 1 Introduction
- 2 Kinematical algebras, r -matrices and quantum contractions
 - Classical Carrollian and Galilean symmetries
 - Deriving their coboundary deformations
- 3 Pictorial overview of (almost) all (coboundary) deformations
 - Comparing with the classifications of r -matrices
 - Carrollian and Galilean cases

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Context and motivation

Landscape of spacetime symmetries:

- Kinematical algebras, e.g. Poincaré, Carroll and Galilei; also their (central) extensions, e.g. Bargmann
- Asymptotic-symmetry algebras, e.g. BMS (and extensions); also their non-Lorentzian versions, e.g. **Carroll-BMS** and **Galilei-BMS**
- Quantum (Hopf-algebraic) deformations of both kinds of algebras, e.g. κ -Poincaré, as well as e.g. κ -**BMS_{ext}**
- Non-Lorentzian versions of the latter, e.g. κ -**Carroll** and κ -**Galilei**

In 2+1 dimensions, with the cosmological constant Λ :

- Recently completed classification of (quantum) deformations
- The cases of $\Lambda \neq 0$ and $\Lambda = 0$ related by quantum contractions^a
- Such deformations arise in the classical theory of (2+1)d gravity

^aKowalski-Glikman, Lukierski & T. T., JHEP **09**, 096 (2020)

Context and motivation – non-Lorentzian kinematics

Carrollian symmetries:

- Associated with the Carroll (or ultrarelativistic) limit $c \rightarrow 0$
 - Ultralocality – trivial dynamics of free particles
 - Two Carroll limits of GR: “electric” and “magnetic”
 - Strong-gravity expansion, BKL conjecture, asymptotic silence^a
- Symmetries of null hypersurfaces one dimension higher
 - Black-hole horizons, plane gravitational waves
 - BMS group \cong a conformal extension of Carroll group

^aMielczarek & T. T., PRD **96**, 024012 (2017)

Galilean symmetries:

- Associated with the Galilei (or “nonrelativistic”) limit $c \rightarrow \infty$
 - Weak-gravity expansion, gravitational waves research
- Algebraic/geometric structures “dual” to the Carrollian ones

Lorentz, Carroll and Galilei in (2+1)d

The brackets of **Poincaré** and **(anti)-de Sitter** algebras in (2+1)d can be written in a unified fashion (with $\Lambda = 0$, $\Lambda < 0$ or $\Lambda > 0$):

$$\begin{aligned}
 [J_0, K_a] &= \epsilon_a^b K_b, & [K_1, K_2] &= -J_0, & [J_0, P_a] &= \epsilon_a^b P_b, & [J_0, P_0] &= 0, \\
 [K_a, P_b] &= \delta_{ab} P_0, & [K_a, P_0] &= P_a, & [P_1, P_2] &= \Lambda J_0, & [P_0, P_a] &= -\Lambda K_a. \quad (1)
 \end{aligned}$$

Denoting $J := J_0$, $T_a := P_a$ and rescaling $Q_a := c K_a$, $T_0 := c P_0$, we take the limit $c \rightarrow 0$ to obtain **Carroll / (anti)-de Sitter-Carroll** algebra:

$$\begin{aligned}
 [J, Q_a] &= \epsilon_a^b Q_b, & [Q_1, Q_2] &= 0, & [J, T_a] &= \epsilon_a^b T_b, & [J, T_0] &= 0, \\
 [Q_a, T_b] &= \delta_{ab} T_0, & [Q_a, T_0] &= 0, & [T_1, T_2] &= \Lambda J, & [T_a, T_0] &= \Lambda Q_a. \quad (2)
 \end{aligned}$$

If we denote $J := J_0$, $T_0 := P_0$ and rescale $Q_a := c^{-1} K_a$, $T_a := c^{-1} P_a$, the limit $c \rightarrow \infty$ leads to **Galilei / (anti)-de Sitter-Galilei** algebra:

$$\begin{aligned}
 [J, Q_a] &= \epsilon_a^b Q_b, & [Q_1, Q_2] &= 0, & [J, T_a] &= \epsilon_a^b T_b, & [J, T_0] &= 0, \\
 [Q_a, T_b] &= 0, & [Q_a, T_0] &= T_a, & [T_1, T_2] &= 0, & [T_a, T_0] &= \Lambda Q_a. \quad (3)
 \end{aligned}$$

These are examples of **contractions**, which relate various kinematical (Lie) algebras with each other.

Why para-Euclidean and para-Poincaré?

Let us also show the brackets of (inhomogeneous) Euclidean algebra:

$$\begin{aligned}
 [J_3, K_a] &= \epsilon_a^b K_b, & [K_1, K_2] &= J_3, & [J_3, P_a] &= \epsilon_a^b P_b, & [J_3, P_3] &= 0, \\
 [K_a, P_b] &= -\delta_{ab} P_3, & [K_a, P_3] &= P_a, & [P_1, P_2] &= 0, & [P_3, P_a] &= 0.
 \end{aligned} \quad (4)$$

It describes different kinematics but is related by the isomorphism

$$K_a \mapsto \Lambda^{-1/2} T_a, \quad P_a \mapsto \Lambda^{1/2} Q_a, \quad J_3 \mapsto J, \quad P_3 \mapsto T_0 \quad (5)$$

with de Sitter-Carroll algebra, hence called the “para-Euclidean”. Meanwhile, Poincaré algebra is mathematically related by the isomorphism

$$K_a \mapsto |\Lambda|^{-1/2} T_a, \quad P_a \mapsto -|\Lambda|^{1/2} Q_a, \quad J_0 \mapsto J, \quad P_0 \mapsto T_0 \quad (6)$$

with anti-de Sitter-Carroll algebra, hence called the “para-Poincaré”.

Meanwhile, the name “expanding/oscillating Newton-Hooke” is sometimes used for dS-Galilei/adS-Galilei algebra.

Coboundary (quantum) deformations

A **Lie bialgebra** is a Lie algebra $(\mathfrak{g}, [,])$ equipped with a cobracket:

$$\begin{aligned} \delta : \mathfrak{g} &\rightarrow \mathfrak{g} \otimes \mathfrak{g}, & (\mathfrak{g}, \delta) &\text{ is a Lie coalgebra,} \\ \forall x, y \in \mathfrak{g} : \delta([x, y]) &= [x \otimes 1 + 1 \otimes x, \delta(y)] + [\delta(x), y \otimes 1 + 1 \otimes y]. \end{aligned} \quad (7)$$

The structure of a **coboundary Lie bialgebra** is determined by

$$r \in \mathfrak{g} \wedge \mathfrak{g}, \quad \forall x \in \mathfrak{g} : \delta(x) = [x \otimes 1 + 1 \otimes x, r]. \quad (8)$$

Such r is called a (antisymmetric) **classical r -matrix** and is actually an **equivalence class** with respect to automorphisms of \mathfrak{g} . Moreover, it is a solution of the classical Yang-Baxter equation

$$[[r, r]] = t\Omega, \quad \Omega \in \mathfrak{g} \otimes \mathfrak{g} \otimes \mathfrak{g}, \quad t \in \mathbb{C}, \quad (9)$$

where Ω is \mathfrak{g} -invariant and $[[,]]$ denotes Schouten bracket. r -matrix is **quasitriangular** if $t \neq 0$, or **triangular** if $t = 0$.

Quantum contractions and deformation parameters

To perform a (quantum) contraction of a quantum-deformed algebra, one not only needs to rescale the appropriate generators but also each deformation parameter q is rescaled to:

$$\hat{q} := q/\omega^2 \quad \text{or} \quad \tilde{q} := q/\omega \quad \text{or} \quad q = q; \quad (10)$$

with $\omega = |\Lambda|$ for $\Lambda \rightarrow 0$, and $\omega = c$ for $c \rightarrow 0$, and $\omega = c^{-1}$ for $c \rightarrow \infty$.

Obtaining the most general contraction limit may require a redefinition of parameters before their rescaling, e.g. γ, ς are replaced by $\tilde{\gamma}, \hat{\gamma}$, where $\varsigma = (\varsigma - \gamma) + \gamma =: c^2 \hat{\gamma} + c \tilde{\gamma}$ (see the next slide).

On the other hand, after the contraction, some parameters can be eliminated by an automorphism (see the next slide).

Quantum contractions and automorphisms

If we transform a deformed algebra by a **suitable automorphism**, this may lead to a **separate contraction limit**, e.g. two representatives of the r -matrix class r_{IV} for anti-de Sitter algebra

$$\begin{aligned}
 r_{IV}(\gamma, \varsigma) &= \gamma (J_0 \wedge K_2 - P'_0 \wedge P'_1 - K_1 \wedge P'_2) - \frac{\varsigma}{2} (J_0 - P'_1) \wedge (K_2 + P'_0), \\
 r_{IV}^a(\gamma, \varsigma) &= -\gamma (J_0 \wedge P'_1 + K_2 \wedge P'_0 + K_1 \wedge P'_2) + \frac{\varsigma}{2} (J_0 - K_2) \wedge (P'_0 + P'_1) \quad (11)
 \end{aligned}$$

($P'_\mu \equiv |\Lambda|^{-1/2} P_\mu$) **are equivalent** but their Carrollian contraction limits

$$\begin{aligned}
 r_{CIV}(\tilde{\gamma}, \tilde{\varsigma}) &= \tilde{\gamma} (J \wedge Q_2 - T'_0 \wedge T'_1 - Q_1 \wedge T'_2) - \frac{\tilde{\varsigma}}{2} (J - T'_1) \wedge (Q_2 + T'_0), \\
 r_{CIVa}(\tilde{\gamma}, \hat{\gamma}) &= -\tilde{\gamma} (J \wedge T'_0 + Q_1 \wedge T'_2 - Q_2 \wedge T'_1) - \hat{\gamma} Q_2 \wedge T'_0 \\
 &\cong -\tilde{\gamma} (J \wedge T'_0 + Q_1 \wedge T'_2 - Q_2 \wedge T'_1) = r_{CIVa}(\tilde{\gamma}) \quad (12)
 \end{aligned}$$

($T'_\mu \equiv |\Lambda|^{-1/2} T_\mu$), describing deformations of adSC algebra, **are not**. The corresponding **automorphism** of adS is **not inherited** by adSC. Automorphisms yield additional contraction limits also for dSG.

Trivialized/reduced deformations – examples

A classical r -matrix is determined up to an **antisymmetric split-Casimir**, i.e. such $\mathcal{C}_s \in \mathfrak{g} \wedge \mathfrak{g}$ that $\forall_{x \in \mathfrak{g}} : [x \otimes 1 + 1 \otimes x, \mathcal{C}_s] = 0$. We find that **Galilei algebra** has an antisymmetric split-Casimir

$$\mathcal{C}_{s1} := Q_1 \wedge T_1 + Q_2 \wedge T_2, \quad (13)$$

while **(anti-)de Sitter-Galilei algebra** has both (13) and

$$\mathcal{C}_{s2} := Q_1 \wedge Q_2 - \Lambda^{-1} T_1 \wedge T_2. \quad (14)$$

The quantum contraction limits are simplified by **dropping such terms**. Incidentally, r -matrix of the form (13) describes **timelike κ -deformation**.

Spacelike κ -deformation for **Carroll** and **(a)dS-Carroll** is **reduced** to:

$$r(\gamma) = \gamma (J_0 \wedge P_1 + K_2 \wedge P_0) \xrightarrow{c \rightarrow 0} r(\hat{\gamma}) = \hat{\gamma} Q_2 \wedge T_0. \quad (15)$$

Classifications of deformations vs their contractions

Semisimple or inhomogeneous-(pseudo)orthogonal algebras have only coboundary deformations, which can be **completely classified in terms of r -matrices**. This has been achieved for 2+1-dimensional algebras:

- Poincaré (as well as Euclidean)^a,
- (anti-)de Sitter^b,
- (anti-)de Sitter-Carroll (isomorphisms with Poincaré/Euclidean)^c.

Quantum contractions of (anti-)de Sitter r -matrices in the limit:

- $\Lambda \rightarrow 0$, leading to Poincaré^d,
- $c \rightarrow 0$, leading to (a)dS-Carroll^c,

recover all r -matrix classes for a given target algebra, up to a few missing terms in some classes.

^aStachura, JPA **31**, 4555 (1998)

^bBorowiec, Lukierski & Tolstoy, JHEP **11**, 187 (2017)

^cT. T., arXiv:2306.05409 [hep-th]

^dKowalski-Glikman, Lukierski & T. T., JHEP **09**, 096 (2020)

Deformations of Poincaré and (a)dS algebras

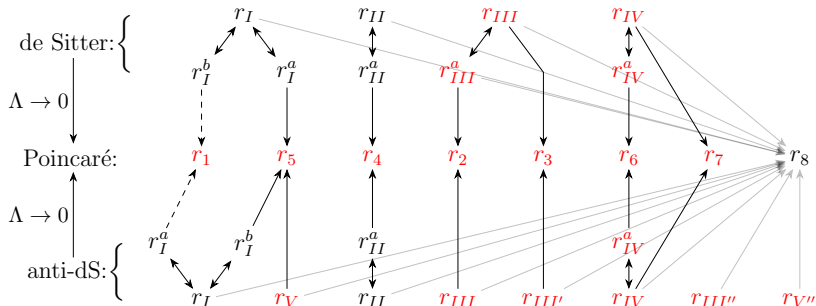


Figure: Quantum $\Lambda \rightarrow 0$ contractions relating all r -matrices for (anti-)de Sitter and Poincaré algebras; double-headed arrows denote automorphisms; arrows leading to r_8 are lightened for legibility; quasitriangular cases are red.

Deformations of Carroll and Galilei algebras

We derived the Carroll/Galilei r -matrices by quantum $c \rightarrow 0 / c \rightarrow \infty$ contractions of the Poincaré ones. Possibly, some deformations can not be obtained in this way. There may also exist **non-coboundary** deformations of these algebras^a.

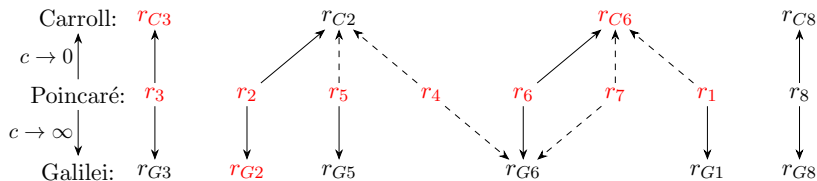


Figure: Quantum $c \rightarrow 0 / c \rightarrow \infty$ contractions relating all r -matrices for Poincaré with those obtained for Carroll/Galilei algebra; a dashed line means that a given contraction leads to a subclass; quasitriangular cases are red.

^aBallesteros et al., PLB **805**, 135461 (2020)

Deformations of (a)dS, (a)dSC and Carroll algebras

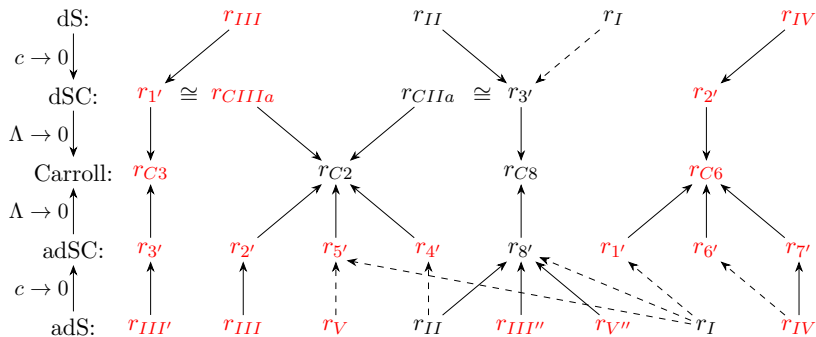


Figure: Quantum $c \rightarrow 0$ and $\Lambda \rightarrow 0$ contractions relating all r -matrices for (anti-)de Sitter and (a)dS-Carroll, and those obtained for Carroll algebra; a dashed line means that a given $c \rightarrow 0$ contraction leads to a subclass.

Deformations of (a)dS, (a)dSG and Galilei algebras

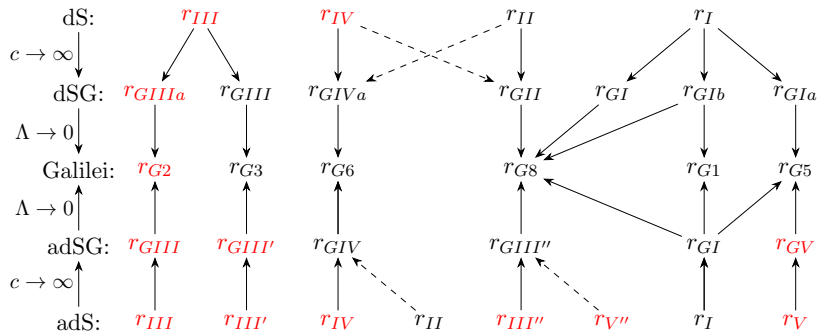


Figure: Quantum $c \rightarrow \infty$ and $\Lambda \rightarrow 0$ contractions relating all r -matrices for (anti-)de Sitter with those obtained for (a)dS-Galilei and Galilei algebras; a dashed line means that a given $c \rightarrow \infty$ contraction leads to a subclass.

Possibly, not all deformations of (a)dSG can be obtained by contractions and there may also exist non-coboundary ones.

Summary – special cases of deformations

The cases of particular interest are **time-** and **spacelike κ -deformations**, and the **Lorentz double**. They also survive under (almost) all quantum contractions ($\Lambda \rightarrow 0$, $c \rightarrow 0$, or $c \rightarrow \infty$) for both $\Lambda > 0$ and $\Lambda < 0$.

algebra	timelike κ -deformation	spacelike κ -deformation	Lorentz double
dSC	$r_{CIII}(\tilde{\gamma}_+) \cong r_{1'}(\gamma)$	$r_{CIIIa}(\hat{\gamma}_-) \cong r_{1'}(\theta_{12})$	$r_{CIV}(\tilde{\gamma}) \cong r_{2'}(\gamma)$
Carroll	$r_{C3}(\tilde{\gamma})$	$r_{C2}(\hat{\gamma})$	$r_{C6}(\tilde{\gamma})$
adSC	$r_{CIII'}(\tilde{\gamma}_+) \cong r_{3'}(\gamma)$	$r_{CIII}(\hat{\gamma}_+) \cong r_{2'}(\theta_{20})$	$r_{CIVa}(\tilde{\gamma}) \cong r_{7'}(\gamma)$
dS	$r_{III}(\gamma_+)$	$r_{III}^a(\gamma_-) \cong r_{III}(\gamma_-)$	$r_{IV}(2\gamma = \varsigma)$
Poincaré	$r_3(\gamma)$	$r_2(\gamma)$	$r_7(\gamma)$
adS	$r_{III'}(\gamma_+) \cong r_{III'}(\gamma_-)$	$r_{III}(\gamma_+) \cong r_{III}(\gamma_-)$	$r_{IV}(2\gamma = -\varsigma)$
dSG	0	$r_{GIIIa}(\tilde{\gamma}_-)$	$r_{GIVa}(2\hat{\gamma} = -\hat{\varsigma})$
Galilei	0	$r_{G2}(\tilde{\gamma})$	$r_{G6}(\hat{\gamma} = \hat{\varsigma})$
adSG	0	$r_{GIII}(\tilde{\gamma}_+) \cong r_{GIII}(\tilde{\gamma}_-)$	$r_{GIV}(2\hat{\gamma} = -\hat{\varsigma})$

Table: r -matrices (only $\neq 0$ deformation parameters) that characterize **special cases** of deformations, depending on a kinematical algebra; quasitriangular ones are red.

Abridged definition of the Hopf algebra

A Hopf algebra A is the vector space over a field K , equipped with a product (e.g. a Lie bracket) $\nabla : A \otimes A \rightarrow A$, satisfying the associativity

$$\nabla \circ (\nabla \otimes \text{id}) = \nabla \circ (\text{id} \otimes \nabla); \quad (16)$$

a coproduct $\Delta : A \rightarrow A \otimes A$, satisfying the coassociativity

$$(\Delta \otimes \text{id}) \circ \Delta = (\text{id} \otimes \Delta) \circ \Delta; \quad (17)$$

and an antipode $S : A \rightarrow A$, satisfying the relation

$$\nabla \circ (S \otimes \text{id}) \circ \Delta = \nabla \circ (\text{id} \otimes S) \circ \Delta = \mathbb{1}. \quad (18)$$

The tensor product of a pair of algebra representations (ρ_1, V_1) , (ρ_2, V_2) (where $\rho_{1,2} : A \rightarrow \text{GL}(V_{1,2})$) is given by $(\rho, V_1 \otimes V_2)$, such that

$$\rho(\mathbf{a})(v_1 \otimes v_2) = (\rho_1 \otimes \rho_2)(\Delta(\mathbf{a}))(v_1 \otimes v_2), \quad (19)$$

where $\mathbf{a} \in A$, $v_{1,2} \in V_{1,2}$.

Example – the Hopf algebra corresponding to r_{III}

Denoting $H_0 \equiv H$, $H_1 \equiv \bar{H}$, $E_{0\pm} \equiv E_{\pm}$, $E_{1\pm} \equiv \bar{E}_{\pm}$ and $q_0 \equiv e^{\gamma/2}$, $q_1 \equiv e^{\bar{\gamma}/2}$, $\theta \equiv e^{\eta/4}$, we write down the deformed brackets

$$[H_k, E_{k\pm}] = E_{k\pm}, \quad [E_{k+}, E_{k-}] = \frac{q_k^{2H_k} - q_k^{-2H_k}}{q_k - q_k^{-1}}, \quad (20)$$

where $k = 0, 1$. In the limit $q_k \rightarrow 1$ it reduces to $[E_{k+}, E_{k-}] = 2H_k$. Meanwhile, the coproducts have the form

$$\begin{aligned} \Delta(H_k) &= H_k \otimes 1 + 1 \otimes H_k, \\ \Delta(E_{k\pm}) &= E_{k\pm} \otimes q_k^{H_k} \theta^{\mp(-1)^k H_{k+1}} + \theta^{\pm(-1)^k H_{k+1}} q_k^{-H_k} \otimes E_{k\pm} \end{aligned} \quad (21)$$

and antipodes

$$S(H_k) = -H_k, \quad S(E_{k\pm}) = -q_k^{\pm 1} E_{k\pm}. \quad (22)$$

The dual of the subalgebra of translations are spacetime coordinates

$$[X_0, X_a] = 2\gamma X_a, \quad [X_a, X_b] = 0, \quad a, b = 1, 2. \quad (23)$$

Poisson structure via the Fock-Rosly construction

\mathfrak{g} equipped with r becomes the Lie algebra of a **Poisson-Lie group** of spacetime symmetries, dual to the particle phase space. At the same time, r determines the **Hopf-algebraic deformation** of \mathfrak{g} , providing the quantization of the theory. The consistency with 3D gravity requires

$$r = r_A + r_S, \quad r_S = \alpha (P_\mu \otimes J^\mu + J^\mu \otimes P^\mu) + \beta (\wedge J^\mu \otimes J_\mu - P^\mu \otimes P_\mu), \quad \alpha, \beta \in \mathbb{R}, \quad (24)$$

where r_S corresponds to the generalized form of the inner product in Chern-Simons action ($\beta = 0$ in the standard case), while r satisfies the homogeneous **Yang-Baxter equation**, hence r_A :

$$\begin{aligned} [[r_A, r_A]] &= -[[r_S, r_S]] \\ &= -(\alpha^2 - \Lambda\beta^2) (\wedge J_0 \wedge J_1 \wedge J_2 + \frac{1}{2} \epsilon^{\mu\nu\sigma} J_\mu \wedge P_\nu \wedge P_\sigma) \\ &\quad - 2\alpha\beta \left(\frac{1}{2} \wedge \epsilon^{\mu\nu\sigma} J_\mu \wedge J_\nu \wedge P_\sigma + P_0 \wedge P_1 \wedge P_2 \right). \end{aligned} \quad (25)$$

We call such a r_A to be **FR-compatible** and **classify all of them** in J. Kowalski-Glikman, J. Lukierski & T. T., JHEP **09**, 096 (2020).

r -matrices of 3D (A)dS algebra relevant for gravity

Calculating the Schouten bracket $[[r_A, r_A]]$, we find that r -matrices are:

	FR-compatible $\forall \alpha, \beta$	FR-compatible for $\beta = 0$	FR-compatible for $\alpha, \beta \neq 0$
$\mathfrak{o}(3, 1)$	r_{III}, r_{III}^a	r_{IV}, r_{IV}^a	
$\mathfrak{so}(2, 2)$	r_{III}	r_{IV}, r_{IV}^a	r_V
$\mathfrak{so}'(2, 2)$			r_{III}

Example – FR-compatible r -matrices of dS algebra:

$$\begin{aligned}
 r_{III}(\gamma - \bar{\gamma}, \gamma + \bar{\gamma}, \eta; \Lambda) &= \frac{1}{2}(\gamma - \bar{\gamma}) \left(J_1 \wedge J_2 - \Lambda^{-1} P_1 \wedge P_2 \right) \\
 &\quad + \Lambda^{-1/2} \frac{1}{2}(\gamma + \bar{\gamma}) \left(J_1 \wedge P_2 - J_2 \wedge P_1 \right) + \Lambda^{-1/2} \frac{\eta}{2} J_0 \wedge P_0, \\
 r_{III}^a(\gamma - \bar{\gamma}, \gamma + \bar{\gamma}, \eta; \Lambda) &= \Lambda^{-1/2} \frac{1}{2}(\gamma - \bar{\gamma}) \left(J_0 \wedge P_2 - J_2 \wedge P_0 \right) \\
 &\quad + \frac{1}{2}(\gamma + \bar{\gamma}) \left(J_0 \wedge J_2 - \Lambda^{-1} P_0 \wedge P_2 \right) + \Lambda^{-1/2} \frac{\eta}{2} J_1 \wedge P_1, \\
 r_{IV}(\gamma, \varsigma; \Lambda) &= \gamma \left(J_1 \wedge J_2 - \Lambda^{-1/2} J_0 \wedge P_0 - \Lambda^{-1} P_1 \wedge P_2 \right) \\
 &\quad + \frac{\varsigma}{2} \left(J_1 - \Lambda^{-1/2} P_2 \right) \wedge \left(J_2 + \Lambda^{-1/2} P_1 \right), \\
 r_{IV}^a(\gamma, \varsigma; \Lambda) &= \Lambda^{-1/2} \gamma \left(J_0 \wedge P_1 - J_1 \wedge P_0 - J_2 \wedge P_2 \right) \\
 &\quad + \Lambda^{-1/2} \frac{\varsigma}{2} \left(J_0 - J_1 \right) \wedge \left(P_0 - P_1 \right). \tag{26}
 \end{aligned}$$

To be compared with P. K. Osei & B. J. Schroers, CQG **35**, 075006 (2018).

r -matrices of (A)dS algebra in the $\Lambda \rightarrow 0$ limit

Quantum IW contractions of r -matrices of (A)dS algebra lead to the following r -matrices of 3D Poincaré algebra:

r -matrix automorphism class ^a	$\mathfrak{o}(3, 1) \downarrow$	$\mathfrak{d}(2, 2) \downarrow$	$\mathfrak{d}'(2, 2) \downarrow$
$r_1 = \chi (J_0 + J_1) \wedge J_2$	r_I^b	r_I^a	
$\hat{r}_2 = \hat{\gamma} (J_0 \wedge P_2 - J_2 \wedge P_0) + \frac{1}{2} \hat{\eta} J_1 \wedge P_1$	\hat{r}_{III}^a	\hat{r}_{III}	
$\hat{r}_3 = \hat{\gamma} (J_1 \wedge P_2 - J_2 \wedge P_1) + \frac{1}{2} \hat{\eta} J_0 \wedge P_0$	\hat{r}_{III}		\hat{r}_{III}
$\hat{r}_4 = \frac{1}{\sqrt{2}} \hat{\chi} (J_+ \wedge P_1 - J_1 \wedge P_+) - \hat{\zeta} J_+ \wedge P_+$	\hat{r}_{II}^a	\hat{r}_{II}^a	
$\hat{r}_5 = \frac{1}{2} \hat{\chi} J_1 \wedge (P_0 + P_2)$	\hat{r}_I^a	\hat{r}_V	
$\hat{r}_6 = \hat{\gamma} (J_0 \wedge P_2 - J_2 \wedge P_0 - J_1 \wedge P_1) - \hat{\zeta} J_+ \wedge P_+$	\hat{r}_{IV}^a	\hat{r}_{IV}^a	
$\hat{r}_7 = \hat{\gamma} (J_0 \wedge P_0 - J_1 \wedge P_1 - J_2 \wedge P_2)$	\hat{r}_{IV}	\hat{r}_{IV}	

(as well as the irrelevant cases $\sim P_\mu \wedge P_\nu$), where $J_+ \equiv \frac{1}{\sqrt{2}}(J_0 + J_2)$, $P_+ \equiv \frac{1}{\sqrt{2}}(P_0 + P_2)$. Only \hat{r}_2 , \hat{r}_6 and \hat{r}_7 are relevant for 3D gravity, i.e.

$$\begin{aligned}
 [[r_1, r_1]] &= [[\hat{r}_4, \hat{r}_4]] = [[\hat{r}_5, \hat{r}_5]] = 0, \\
 [[\hat{r}_3, \hat{r}_3]] &= \hat{\gamma}^2 \epsilon^{\mu\nu\sigma} J_\mu \wedge P_\nu \wedge P_\sigma, \\
 [[\hat{r}_2, \hat{r}_2]] &= [[\hat{r}_6, \hat{r}_6]] = [[\hat{r}_7, \hat{r}_7]] = -\hat{\gamma}^2 \epsilon^{\mu\nu\sigma} J_\mu \wedge P_\nu \wedge P_\sigma. \quad (27)
 \end{aligned}$$

^aP. Stachura, J. Phys. A: Math. Gen. **31**, 4555 (1998)

Chern-Simons action of 3D gravity (with Λ)

Instead of the metric $g_{\alpha\beta}$, gravity can be described in terms of the vielbein e_α^μ and spin connection $\omega_\alpha^{\mu\nu}$, defined as

$$e_\alpha^\mu e_\beta^\nu \eta_{\mu\nu} = g_{\alpha\beta}, \quad \omega_\alpha^{\mu\nu} = e_\beta^\mu \partial_\alpha e^{\beta\nu} + e_\beta^\mu \Gamma_{\alpha\gamma}^\beta e^{\gamma\nu}. \quad (28)$$

In (2+1)D, they neatly combine into a **gauge field** – with values in the **local isometry algebra \mathfrak{g}** (3D Poincaré or (Anti-)de Sitter) – which is the Cartan connection

$$A = -\frac{1}{2} \epsilon_{\nu\sigma}^\mu \omega_\alpha^{\nu\sigma} J_\mu dx^\alpha + e_\alpha^\mu P_\mu dx^\alpha, \quad (29)$$

where J_μ, P_μ are generators of \mathfrak{g} . Then, the Einstein-Hilbert action **can be written as** the Chern-Simons theory action

$$S = \frac{1}{16\pi G} \int \left(\langle dA \wedge A \rangle + \frac{1}{3} \langle A \wedge [A, A] \rangle \right) \quad (30)$$

if the **scalar product** on \mathfrak{g} is given by

$$\langle J_\mu, P_\nu \rangle = \eta_{\mu\nu}, \quad \langle J_\mu, J_\nu \rangle = \langle P_\mu, P_\nu \rangle = 0. \quad (31)$$

How to couple a conical defect

The **massive point-particle** spacetime interval looks as in vacuum:

$$ds^2 = (1 - \Lambda r^2) dt^2 - (1 - \Lambda r^2)^{-1} dr^2 - r^2 d\tilde{\phi}^2, \quad (32)$$

when the polar angle is rescaled to $\tilde{\phi} := (1 - 4Gm) \phi$ (a conical defect). Similarly, **spin** $\neq 0$ introduces a jump in the time coordinate.

If spacetime has the topology $\mathbb{R} \times \mathcal{S}$, the field A may be expressed as $A = A_t dt + A_S$ and the action of **gravity with a particle** is

$$S = \int dt L = \frac{1}{16\pi G} \int dt \int_S \langle \dot{A}_S \wedge A_S \rangle - \int dt \langle c_0 h^{-1} \dot{h} \rangle + \int dt \int_S \left\langle A_t \left(\frac{1}{8\pi G} F_S - h c_0 h^{-1} \delta^2(\vec{x} - \vec{x}_*) dx^1 \wedge dx^2 \right) \right\rangle. \quad (33)$$

Mass $\neq 0$ and spin of a particle are encoded by $\mathfrak{g} \ni c_0 = m J_0 + s P_0$, while a **gauge group** element h acting via $h c_0 h^{-1} = \mathbf{p} + \mathbf{j}$ determines its momentum $\mathbf{p} = p^\mu J_\mu$ and (generalized) angular momentum $\mathbf{j} = j^\mu P_\mu$.

Relating the gravitational and particle DOFs

A_t acts as a Lagrange multiplier imposing a **constraint on the curvature** of spatial connection $F_S = dA_S + [A_S, A_S]$:

$$F_S = 8\pi G h c_0 h^{-1} \delta^2(\vec{x} - \vec{x}_*) dx^1 \wedge dx^2. \quad (34)$$

From $F_S = R_S + T_S + C_S$ (C_S is the cosmological-constant term), it follows that the spatial **Riemann curvature and torsion** are given by

$$\begin{aligned} R_S &= -C_S + 8\pi G \mathbf{p} \delta^2(\vec{x} - \vec{x}_*) dx^1 \wedge dx^2, \\ T_S &= 8\pi G \mathbf{j} \delta^2(\vec{x} - \vec{x}_*) dx^1 \wedge dx^2, \end{aligned} \quad (35)$$

i.e. **they vanish** (on the constant background $R_S = -C_S$) everywhere except a **singularity at the particle's** worldline.

Alternatively, \mathfrak{g} can be equipped with the scalar product

$$\langle \mathbf{J}_\mu, \mathbf{P}_\nu \rangle = 0, \quad \langle \mathbf{J}_\mu, \mathbf{J}_\nu \rangle = -\Lambda^{-1} \langle \mathbf{P}_\mu, \mathbf{P}_\nu \rangle = \eta_{\mu\nu}. \quad (36)$$

If our action is defined using it, \mathbf{j} generates R_S and \mathbf{p} generates T_S .