

STELLAR STRUCTURE IN SCALAR-TENSOR SYMMETRIC TELEPARALLEL  
GRAVITY

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# SCALAR-TENSOR SYMMETRIC TELEPARALLEL GRAVITY<sup>2</sup>

## INTRODUCTION

Vanishing curvature and torsion

$$R^\sigma{}_{\rho\mu\nu} = \partial_\mu \Gamma^\sigma{}_{\nu\rho} - \partial_\nu \Gamma^\sigma{}_{\mu\rho} + \Gamma^\sigma{}_{\mu\lambda} \Gamma^\lambda{}_{\nu\rho} - \Gamma^\sigma{}_{\nu\lambda} \Gamma^\lambda{}_{\mu\rho} = 0, \quad (1)$$

$$T^\sigma{}_{\mu\nu} = \Gamma^\sigma{}_{\mu\nu} - \Gamma^\sigma{}_{\nu\mu} = 0, \quad (2)$$

and nonmetricity

$$Q_{\rho\mu\nu} = \nabla_\rho g_{\mu\nu} = \partial_\rho g_{\mu\nu} - \Gamma^\lambda{}_{\rho\mu} g_{\lambda\nu} - \Gamma^\lambda{}_{\rho\nu} g_{\mu\lambda}, \quad (3)$$

$$Q = -\frac{1}{4} Q_{\lambda\mu\nu} Q^{\lambda\mu\nu} + \frac{1}{2} Q_{\lambda\mu\nu} Q^{\lambda\mu\nu} + \frac{1}{4} Q_\mu Q^\mu - \frac{1}{2} Q_\mu \bar{Q}^\mu, \quad Q_\mu = Q^\nu{}_{\mu\nu}, \quad \bar{Q}_\mu = Q^\nu{}_{\nu\mu}. \quad (4)$$

Levi-Civita Riemann tensor and Ricci scalar

$$\mathring{R}^\sigma{}_{\rho\mu\nu} = \mathring{\nabla}_\nu L^\sigma{}_{\mu\rho} - \mathring{\nabla}_\mu L^\sigma{}_{\nu\rho} + L^\sigma{}_{\nu\lambda} L^\lambda{}_{\mu\rho} - L^\sigma{}_{\mu\lambda} L^\lambda{}_{\nu\rho}, \quad (5)$$

$$L^\sigma{}_{\mu\nu} = \frac{1}{2} (Q^\sigma{}_{\mu\nu} - Q_\mu{}^\sigma{}_\nu - Q_\nu{}^\sigma{}_\mu), \quad (6)$$

$$\mathring{R} = Q + \mathring{\nabla}_\mu (\bar{Q}^\mu - Q^\mu) \quad (7)$$

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<sup>2</sup>Bahamonde et al. 2022

# SCALAR-TENSOR SYMMETRIC TELEPARALLEL GRAVITY

## FIELD EQUATIONS

The action reads

$$S = \frac{1}{2\kappa^2} \int d^4x \sqrt{-g} \left( \mathcal{A}(\Phi)Q - \mathcal{B}(\Phi)g^{\alpha\beta} \partial_\alpha \Phi \partial_\beta \Phi - 2\mathcal{V}(\Phi) \right) + S_m [g, \Psi] . \quad (8)$$

Metric field equation<sup>3</sup>

$$\begin{aligned} \kappa^2 T_{\mu\nu} = & \mathcal{A}(\Phi) \overset{\circ}{G}_{\mu\nu} + 2 \frac{d\mathcal{A}(\Phi)}{d\Phi} P^\alpha{}_{\mu\nu} \partial_\alpha \Phi - \mathcal{B}(\Phi) \partial_\mu \Phi \partial_\nu \Phi \\ & + \frac{g^{\mu\nu}}{2} \left( \mathcal{B}(\Phi) g^{\alpha\beta} \partial_\alpha \Phi \partial_\beta \Phi + 2\mathcal{V}(\Phi) \right) , \end{aligned} \quad (9)$$

Scalar Field Equation

$$0 = 2\overset{\circ}{\square}\Phi \mathcal{B}(\Phi) + \frac{d\mathcal{B}}{d\Phi} g^{\alpha\beta} \partial_\alpha \Phi \partial_\beta \Phi + \frac{d\mathcal{A}}{d\Phi} Q - 2 \frac{d\mathcal{V}}{d\Phi} , \quad (10)$$

Connection Equation

$$0 = \frac{1}{2} \frac{d\mathcal{A}}{d\Phi} Q_\alpha K_\mu{}^\alpha + \frac{d^2\mathcal{A}}{d\Phi^2} \partial_\alpha \Phi K_\mu{}^\alpha + \frac{d\mathcal{A}}{d\Phi} \nabla_\alpha K_\mu{}^\alpha . \quad (11)$$

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$$P^\alpha{}_{\mu\nu} = -\frac{1}{4} (g_{\mu\nu} \overset{\circ}{Q}^\alpha + \delta^\alpha{}_\mu Q_\nu) - \frac{1}{4} Q^\alpha{}_{\mu\nu} + \frac{1}{4} Q_\mu{}^\alpha{}_\nu + \frac{1}{4} g_{\mu\nu} Q_\alpha , \quad K_\mu{}^\beta = \left[ \frac{1}{2} Q_\mu \delta^{\alpha\beta} - \frac{1}{2} \delta_\mu^\alpha Q^\beta - Q_\mu{}^\alpha{}_\beta + \delta_\mu^\alpha Q_\gamma{}^\beta \right] (\partial_\alpha \mathcal{A})$$

# SCALAR-TENSOR SYMMETRIC TELEPARALLEL GRAVITY

## STATIC AND SPHERICALLY SYMMETRIC GEOMETRY

### Connection<sup>4</sup>

- ▶ Two solution sets for a static and spherically symmetric connection with zero curvature and torsion
- ▶ First solution set
  - Three independent components  $\Gamma^{\phi}_{r\phi}$ ,  $\Gamma^t_{rr}$ ,  $\Gamma^r_{rr}$  and constraint for  $(\Gamma^{\phi}_{r\phi})'$
  - Effectively reduces to GR

$$\frac{3}{2} \frac{d\mathcal{A}}{d\Phi} \Phi'(r) = 0, \quad (12)$$

- ▶ Second solution set
  - Four independent components  $\Gamma^t_{\theta\theta}$ ,  $\Gamma^t_{rr}$ ,  $\Gamma^r_{\theta\theta}$ ,  $\Gamma^r_{rr}$  and constraints for  $(\Gamma^r_{\theta\theta})'$ ,  $(\Gamma^t_{\theta\theta})'$
  - Allows for solutions beyond GR

$$\frac{1}{2} [2c(k - 2c)\Gamma^t_{\theta\theta} + k] \frac{d\mathcal{A}}{d\Phi} \Phi'(r) = 0, \quad (13)$$

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<sup>4</sup>D'Ambrosio et al. 2022

# SCALAR-TENSOR SYMMETRIC TELEPARALLEL GRAVITY

## STATIC AND SPHERICALLY SYMMETRIC GEOMETRY

Connection, solution set 2 branch 2

$$\Gamma^{\rho}_{\mu\nu}(r, \theta) = \left[ \begin{array}{cccc} \left[ \begin{array}{cccc} 0 & 0 & 0 & 0 \\ 0 & \Gamma^t_{rr} & 0 & 0 \\ 0 & 0 & \Gamma^t_{\theta\theta} & 0 \\ 0 & 0 & 0 & \Gamma^t_{\theta\theta} \sin^2(\theta) \end{array} \right] & \left[ \begin{array}{cccc} 0 & 0 & 0 & 0 \\ 0 & \Gamma^r_{rr} & 0 & 0 \\ 0 & 0 & \Gamma^r_{\theta\theta} & 0 \\ 0 & 0 & 0 & \Gamma^r_{\theta\theta} \sin^2(\theta) \end{array} \right] \\ \left[ \begin{array}{cccc} 0 & 0 & 0 & 0 \\ 0 & 0 & -\frac{1}{\Gamma^r_{\theta\theta}} & 0 \\ 0 & -\frac{1}{\Gamma^r_{\theta\theta}} & 0 & 0 \\ 0 & 0 & 0 & -\sin(\theta) \cos(\theta) \end{array} \right] & \left[ \begin{array}{cccc} 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & -\frac{1}{\Gamma^r_{\theta\theta}} \\ 0 & 0 & 0 & \cot(\theta) \\ 0 & -\frac{1}{\Gamma^r_{\theta\theta}} & \cot(\theta) & 0 \end{array} \right] \end{array} \right] \quad (14)$$

with the relations

$$\frac{d\Gamma^r_{\theta\theta}}{dr} = -\Gamma^r_{\theta\theta}\Gamma^r_{rr} - 1, \quad \frac{d\Gamma^t_{\theta\theta}}{dr} = -\Gamma^r_{\theta\theta}\Gamma^t_{rr} - \frac{\Gamma^t_{\theta\theta}}{\Gamma^r_{\theta\theta}}. \quad (15)$$

Metric

$$ds^2 = -g_{tt}(r)dt^2 + g_{rr}(r)dr^2 + r^2d\Omega^2. \quad (16)$$

# SCALAR-TENSOR SYMMETRIC TELEPARALLEL GRAVITY

## STATIC AND SPHERICALLY SYMMETRIC FIELD EQUATIONS

Metric field equations

$$-8\pi G\rho = \frac{\mathcal{B}\Phi'^2}{2g_{rr}} + \frac{\Phi' \left( r^2 + 2r\Gamma^r_{\theta\theta} + (\Gamma^r_{\theta\theta})^2 g_{rr} \right) d\mathcal{A}}{r^2\Gamma^r_{\theta\theta}g_{rr}} \frac{d\mathcal{A}}{d\Phi} - \frac{(rg'_{rr} + g_{rr}^2 - g_{rr}) \mathcal{A}}{r^2g_{rr}^2} + \mathcal{V} \quad (17)$$

$$8\pi Gp = -\frac{\mathcal{B}\Phi'^2}{2g_{rr}} - \frac{\Phi' \left( r^2 - (\Gamma^r_{\theta\theta})^2 g_{rr} \right) d\mathcal{A}}{r^2\Gamma^r_{\theta\theta}g_{rr}} \frac{d\mathcal{A}}{d\Phi} + \frac{(rg'_{tt} - g_{rr}g_{tt} + g_{tt}) \mathcal{A}}{r^2g_{rr}g_{tt}} + \mathcal{V} \quad (18)$$

$$8\pi Gp = \frac{\mathcal{B}\Phi'^2}{2g_{rr}} + \frac{\Phi' (r\Gamma^r_{\theta\theta}g'_{tt} + 2rg_{tt} + 2\Gamma^r_{\theta\theta}g_{tt}) d\mathcal{A}}{2r\Gamma^r_{\theta\theta}g_{rr}g_{tt}} \frac{d\mathcal{A}}{d\Phi} + \mathcal{V} \\ - \frac{\left( -2rg_{rr}g_{tt}g''_{tt} + rg_{rr} (g'_{tt})^2 + rg_{tt}g'_{rr}g'_{tt} - 2g_{rr}g_{tt}g'_{tt} + 2g_{tt}^2g'_{rr} \right) \mathcal{A}}{4rg_{rr}^2g_{tt}^2} \quad (19)$$

# SCALAR-TENSOR SYMMETRIC TELEPARALLEL GRAVITY

## STATIC AND SPHERICALLY SYMMETRIC FIELD EQUATIONS

Connection equation

$$0 = \frac{2 \left( r^2 - (\Gamma^r_{\theta\theta})^2 g_{rr} \right) \Phi'^2}{r^2 \Gamma^r_{\theta\theta} g_{rr}} \frac{d^2 \mathcal{A}(\Phi)}{d\Phi^2} + \left( 2r^2 g_{rr} g_{tt} \Phi'' + r^2 g_{rr} \Phi' g'_{tt} - r^2 g_{tt} \Phi' g'_{rr} + 4r g_{rr} g_{tt} \Phi' - 2 (\Gamma^r_{\theta\theta})^2 g_{rr}^2 g_{tt} \Phi'' - (\Gamma^r_{\theta\theta})^2 g_{rr}^2 \Phi' g'_{tt} - (\Gamma^r_{\theta\theta})^2 g_{rr} g_{tt} \Phi' g'_{rr} - 4 \Gamma^r_{\theta\theta} g_{rr}^2 g_{tt} \Phi' (\Gamma^r_{\theta\theta})' \right) \frac{1}{r^2 \Gamma^r_{\theta\theta} g_{rr}^2 g_{tt}} \frac{d\mathcal{A}}{d\Phi} \quad (20)$$

Scalar field equation

$$0 = -2 \frac{d\mathcal{V}(\Phi)}{d\Phi} + \frac{\Phi'^2}{g_{rr}} \frac{d\mathcal{B}(\Phi)}{d\Phi} + \frac{(2r g_{rr} g_{tt} \Phi'' + r g_{rr} \Phi' g'_{tt} - r g_{tt} \Phi' g'_{rr} + 4 g_{rr} g_{tt} \Phi') \mathcal{B}(\Phi)}{r g_{rr}^2 g_{tt}} - \mathcal{Q} \frac{d\mathcal{A}}{d\Phi} \quad (21)$$

# TOLMAN-OPPENHEIMER-VOLKOFF EQUATION<sup>5</sup>

## MODIFIED GRAVITY THEORIES

- ▶ Modified gravity theory that can be written as

$$\sigma(\Psi^i) \left( \mathring{G}_{\mu\nu} - W_{\mu\nu} \right) = \kappa^2 T_{\mu\nu} , \quad (22)$$

$$\mathring{G}_{\mu\nu} = \kappa^2 T_{\mu\nu}^{eff} = \frac{\kappa^2}{\sigma} T_{\mu\nu} + W_{\mu\nu} , \quad (23)$$

where  $\kappa^2 = 8\pi G$  and such that  $\mathring{\nabla}_\mu T_{eff}^{\mu\nu} = 0$

- ▶ Perfect fluid

$$T_{\mu\nu} = (\rho + p) u_\mu u_\nu + p g_{\mu\nu} , \quad u_\mu u^\mu = -1 , \quad (24)$$

- ▶ Static and spherically symmetric geometry  $u_0 = -\sqrt{g_{tt}}$ .

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<sup>5</sup>Wojnar and Velten 2016



# TOLMAN-OPPENHEIMER-VOLKOFF EQUATION

## MODIFIED GRAVITY THEORIES

The components of the Ricci tensor read

$$R_{tt} = -\frac{g''_{tt}}{2g_{rr}} + \frac{g'_{tt}}{4g_{rr}} \left( \frac{g'_{rr}}{g_{rr}} + \frac{g'_{tt}}{g_{tt}} \right) - \frac{g'_{tt}}{rg_{rr}} = -\frac{\kappa^2}{2\sigma} (\rho + 3p) g_{tt} + W_{tt} + \frac{g_{tt}W}{2}, \quad (25)$$

$$R_{rr} = \frac{g''_{tt}}{2g_{tt}} - \frac{g'_{tt}}{4g_{tt}} \left( \frac{g'_{rr}}{g_{rr}} + \frac{g'_{tt}}{g_{tt}} \right) - \frac{g'_{rr}}{rg_{rr}} = -\frac{\kappa^2}{2\sigma} (\rho - p) g_{rr} + W_{rr} - \frac{g_{rr}W}{2}, \quad (26)$$

$$R_{\theta\theta} = -1 - \frac{r}{2g_{rr}} \left( -\frac{g'_{rr}}{g_{rr}} + \frac{g'_{tt}}{g_{tt}} \right) + \frac{1}{g_{rr}} = -\frac{\kappa^2}{2\sigma} (\rho - p) r^2 + W_{\theta\theta} - \frac{r^2W}{2}. \quad (27)$$

Combining these equations to

$$\frac{R_{tt}}{2g_{tt}} + \frac{R_{rr}}{2g_{rr}} + \frac{R_{\theta\theta}}{r^2} \rightarrow g_{rr}(r) = \left( 1 - \frac{2GM(r)}{r} \right)^{-1}, \quad (28)$$

and using  $\overset{\circ}{\nabla}_{\mu} T^{\mu\nu}_{eff} = 0$  as well as the metric component relations in (27)

$$\frac{g'_{tt}}{g_{tt}} = \frac{g_{rr} - 1}{r} + \frac{\kappa^2 r g_{rr}}{\sigma} \Pi, \quad \frac{g'_{rr}}{g_{rr}} = \frac{1 - g_{rr}}{r} + \frac{\kappa^2 r g_{rr}}{\sigma} Q \quad (29)$$

# TOLMAN-OPPENHEIMER-VOLKOFF EQUATION

## MODIFIED GRAVITY THEORIES

### TOV equation

$$\begin{aligned} \left(\frac{\Pi}{\sigma}\right)' &= -\frac{GM}{r^2} \left(\frac{Q}{\sigma} + \frac{\Pi}{\sigma}\right) \left(1 + \frac{4\pi r^3 \Pi}{\sigma \mathcal{M}}\right) \left(1 - \frac{2GM}{r}\right)^{-1} \\ &\quad + \frac{2}{\kappa^2 r} \left(\frac{W_{rr}}{g_{rr}} - \frac{W_{\theta\theta}}{r^2}\right) \end{aligned} \quad (30)$$

with effective energy density and pressure

$$Q(r) := \rho(r) + \frac{\sigma(r)W_{tt}(r)}{\kappa^2 g_{tt}(r)}, \quad \Pi(r) := p(r) + \frac{\sigma(r)W_{rr}(r)}{\kappa^2 g_{rr}(r)}, \quad (31)$$

and mass

$$\frac{d\mathcal{M}(r)}{dr} = 4\pi r^2 \frac{Q(r)}{\sigma(r)}. \quad (32)$$

# TOLMAN-OPPENHEIMER-VOLKOFF EQUATION

## SCALAR-TENSOR SYMMETRIC TELEPARALLEL GRAVITY

Metric field equation

$$\begin{aligned}\kappa^2 T_{\mu\nu} &= \sigma(\Psi^i) \left( \overset{\circ}{G}_{\mu\nu} - W_{\mu\nu} \right) , \\ \kappa^2 T_{\mu\nu} &= \mathcal{A} \overset{\circ}{G}_{\mu\nu} + 2 \frac{d\mathcal{A}}{d\Phi} P^\alpha{}_{\mu\nu} \partial_\alpha \Phi - \mathcal{B} \partial_\mu \Phi \partial_\nu \Phi + \frac{g^{\mu\nu}}{2} \left( \mathcal{B} g^{\alpha\beta} \partial_\alpha \Phi \partial_\beta \Phi + 2\mathcal{V} \right) ,\end{aligned}\quad (33)$$

such that  $\sigma(\Psi^i) = \mathcal{A}(\Phi)$  and

$$W_{\mu\nu} = -\frac{1}{\mathcal{A}} \left[ 2 \frac{d\mathcal{A}}{d\Phi} P^\alpha{}_{\mu\nu} \partial_\alpha \Phi - \mathcal{B} \partial_\mu \Phi \partial_\nu \Phi + \frac{g^{\mu\nu}}{2} \left( \mathcal{B} g^{\alpha\beta} \partial_\alpha \Phi \partial_\beta \Phi + 2\mathcal{V} \right) \right]. \quad (34)$$

For the connection solution, set 2 branch 2, in a static and spherically symmetric geometry

$$W_{tt}(r) = \frac{g_{tt} \Phi'}{g_{rr} \mathcal{A}} \left\{ \frac{\mathcal{B} \Phi'}{2} + \frac{r^2 + 2r\Gamma^r{}_{\theta\theta} + (\Gamma^r{}_{\theta\theta})^2 g_{rr}}{r^2 \Gamma^r{}_{\theta\theta}} \frac{d\mathcal{A}}{d\Phi} \right\} + \frac{g_{tt} \mathcal{V}}{\mathcal{A}} , \quad (35)$$

$$W_{rr}(r) = \frac{\Phi'}{\mathcal{A}} \left\{ \frac{\mathcal{B} \Phi'}{2} + \frac{r^2 - (\Gamma^r{}_{\theta\theta})^2 g_{rr}}{r^2 \Gamma^r{}_{\theta\theta}} \frac{d\mathcal{A}}{d\Phi} \right\} - \frac{g_{rr} \mathcal{V}}{\mathcal{A}} , \quad (36)$$

$$W_{\theta\theta}(r) = -\frac{r^2 \Phi'}{g_{rr} \mathcal{A}} \left\{ \frac{\mathcal{B} \Phi'}{2} + \left( \frac{g'_{tt}}{2g_{tt}} + \frac{1}{\Gamma^r{}_{\theta\theta}} + \frac{1}{r} \right) \frac{d\mathcal{A}}{d\Phi} \right\} - \frac{r^2 \mathcal{V}}{\mathcal{A}} , \quad (37)$$

# TOLMAN-OPPENHEIMER-VOLKOFF EQUATION

## SCALAR-TENSOR SYMMETRIC TELEPARALLEL GRAVITY

'The effective energy density and pressure read

$$Q(r) = \rho(r) - \frac{\Phi'}{\kappa^2 r g_{rr}} \left\{ \frac{r \mathcal{B} \Phi'}{2} + \left( 2 + \frac{r}{\Gamma^r_{\theta\theta}} + \frac{g_{rr} \Gamma^r_{\theta\theta}}{r} \right) \frac{d\mathcal{A}}{d\Phi} \right\} - \frac{\mathcal{V}}{\kappa^2}, \quad (38)$$

$$\Pi(r) = p(r) - \frac{\Phi'}{\kappa^2 r g_{rr}} \left\{ \frac{r \mathcal{B} \Phi'}{2} + \left( \frac{r}{\Gamma^r_{\theta\theta}} - \frac{g_{rr} \Gamma^r_{\theta\theta}}{r} \right) \frac{d\mathcal{A}}{d\Phi} \right\} + \frac{\mathcal{V}}{\kappa^2}, \quad (39)$$

and hence the mass equation is

$$\frac{d\mathcal{M}(r)}{dr} = \frac{4\pi r^2}{\mathcal{A}} \left[ \rho(r) - \frac{\Phi'}{\kappa^2 r g_{rr}} \left\{ \frac{r \mathcal{B} \Phi'}{2} + \left( 2 + \frac{r}{\Gamma^r_{\theta\theta}} + \frac{g_{rr} \Gamma^r_{\theta\theta}}{r} \right) \frac{d\mathcal{A}}{d\Phi} \right\} - \frac{\mathcal{V}}{\kappa^2} \right]. \quad (40)$$

# TOLMAN-OPPENHEIMER-VOLKOFF EQUATION

## SCALAR-TENSOR SYMMETRIC TELEPARALLEL GRAVITY

$$\left(\frac{\Pi}{\sigma}\right)' = -\frac{GM}{r^2} \left(\frac{Q}{\sigma} + \frac{\Pi}{\sigma}\right) \left(1 + \frac{4\pi r^3 \Pi}{\sigma \mathcal{M}}\right) \left(1 - \frac{2GM}{r}\right)^{-1} \quad (41)$$

$$+ \frac{2}{\kappa^2 r} \left(\frac{W_{rr}}{g_{rr}} - \frac{W_{\theta\theta}}{r^2}\right),$$

$$p' = -\frac{GM}{r^2} \left(\rho + p - \frac{\mathcal{B}(\Phi')^2}{\kappa^2 g_{rr}} - \frac{2\Phi' \frac{d\mathcal{A}}{d\Phi}}{\kappa^2 \Gamma^r_{\theta\theta} g_{rr}} - \frac{2\Phi' \frac{d\mathcal{A}}{d\Phi}}{\kappa^2 r g_{rr}}\right) \left(1 - \frac{2GM}{r}\right)^{-1} \quad (42)$$








$$\times \left(\frac{4\pi r^3}{\mathcal{A}\mathcal{M}} \left(p - \frac{\mathcal{B}(\Phi')^2}{2\kappa^2 g_{rr}} + \frac{\mathcal{V}}{\kappa^2} - \frac{\Phi' \frac{d\mathcal{A}}{d\Phi}}{\kappa^2 \Gamma^r_{\theta\theta} g_{rr}} + \frac{\Gamma^r_{\theta\theta} \Phi' \frac{d\mathcal{A}}{d\Phi}}{\kappa^2 r^2}\right) + 1\right)$$

$$+ \frac{1}{\kappa^2 \mathcal{A}} \left(\Phi' \frac{d\mathcal{A}}{d\Phi} + \frac{r\mathcal{B}\Phi'^2}{2} + \frac{r\Phi' \frac{d\mathcal{A}}{d\Phi}}{\Gamma^r_{\theta\theta}}\right) \left(-\frac{\mathcal{B}(\Phi')^2}{2g_{rr}} + \mathcal{V} - \frac{\Phi' \frac{d\mathcal{A}}{d\Phi}}{\Gamma^r_{\theta\theta} g_{rr}} + \frac{\Gamma^r_{\theta\theta} \Phi' \frac{d\mathcal{A}}{d\Phi}}{r^2}\right)$$

$$+ \frac{p}{\mathcal{A}} \left(\Phi' \frac{d\mathcal{A}}{d\Phi} - \frac{r\mathcal{B}\Phi'^2}{2} - \frac{r\Phi' \frac{d\mathcal{A}}{d\Phi}}{\Gamma^r_{\theta\theta}}\right) + \frac{(\Phi')^2 \frac{d\mathcal{A}}{d\Phi}}{\kappa^2 r \mathcal{A}} \left(\frac{d\mathcal{A}}{d\Phi} - \Phi'\right)$$

$$+ \frac{1}{\kappa^2} \left(\frac{\mathcal{B}\Phi'}{2r} + \Phi' \frac{d\mathcal{A}}{d\Phi} \left(\frac{1}{r\Gamma^r_{\theta\theta}} + \frac{1}{r^2}\right)\right) \left(\frac{1}{g_{rr}} - 1\right)$$

## REFERENCES

-  Bahamonde, Sebastian et al. (Aug. 2022). “Black hole solutions in scalar-tensor symmetric teleparallel gravity”. In: *Journal of Cosmology and Astroparticle Physics* 2022.08, p. 082. DOI: 10.1088/1475-7516/2022/08/082.
-  D’Ambrosio, Fabio et al. (Jan. 2022). “Black holes in  $f(Q)$  gravity”. In: *Physical Review D* 105.2. DOI: 10.1103/physrevd.105.024042.
-  Doneva, Daniela D. et al. (2022). *Scalarization*. DOI: 10.48550/ARXIV.2211.01766.
-  Kalita, Surajit and Banibrata Mukhopadhyay (2018). “Modified Einstein’s gravity to probe the sub-and super-Chandrasekhar limiting mass white dwarfs: a new perspective to unify under-and over-luminous type Ia supernovae”. In: *Journal of Cosmology and Astroparticle Physics* 2018.09, p. 007.
-  Kalita, Surajit, Lupamudra Sarmah, and Aneta Wojnar (2023). “Metric-affine effects in crystallization processes of white dwarfs”. In: *Physical Review D* 107.4, p. 044072.
-  Olmo, Gonzalo J., Diego Rubiera-Garcia, and Aneta Wojnar (Sept. 2020). “Stellar structure models in modified theories of gravity: Lessons and challenges”. In: *Physics Reports* 876, pp. 1–75. DOI: 10.1016/j.physrep.2020.07.001.
-  Wojnar, Aneta and Hermano Velten (Dec. 2016). “Equilibrium and stability of relativistic stars in extended theories of gravity”. In: *The European Physical Journal C* 76.12. DOI: 10.1140/epjc/s10052-016-4549-z.