



# The Einstein Telescope Project







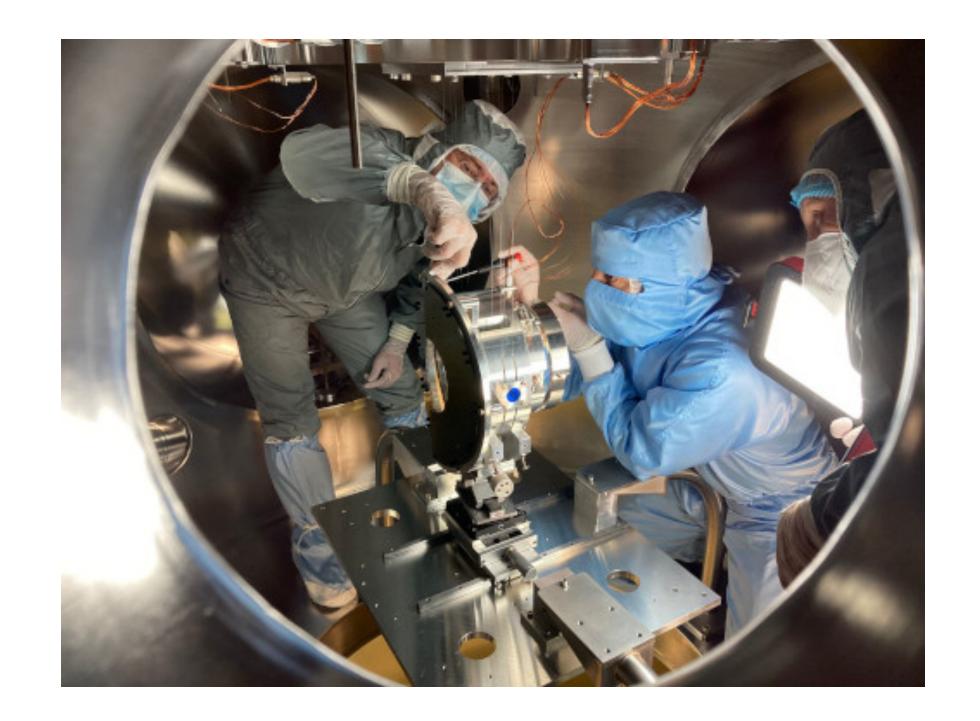
M. Martínez



Plan Complementario, Zaragoza, June 2024

# Outline

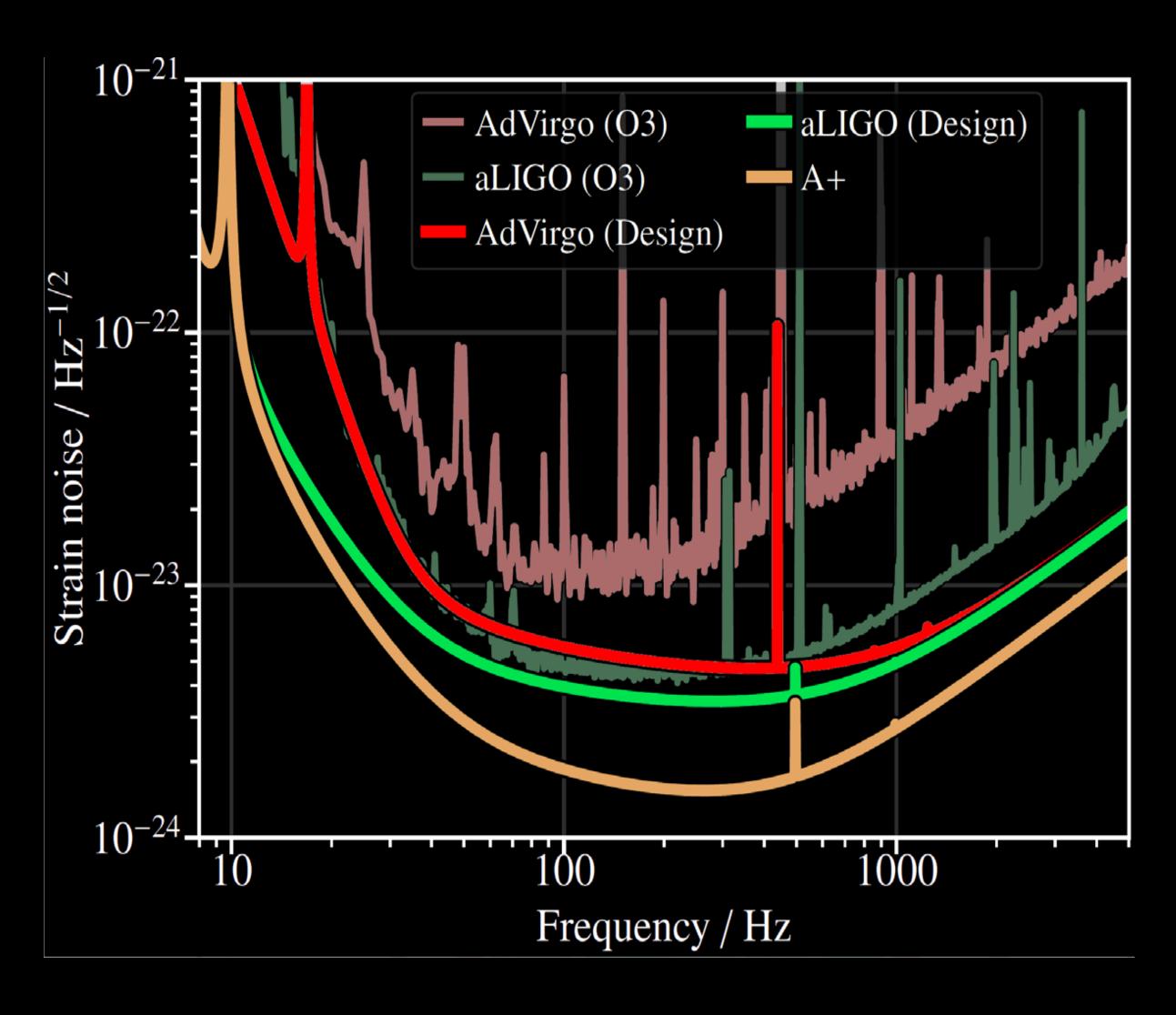
- The Einstein Telescope
  - Conceptual Design
  - Physics Potential
  - Time Scales, costs, location
  - R&D for its realization
- ET @ ESFRI 2021 roadmap
  - Preparatory Phase
  - Organization
  - Political Support
  - ET@CERN activities
  - Ongoing geometry discussions
  - ET-Spain
- Final notes

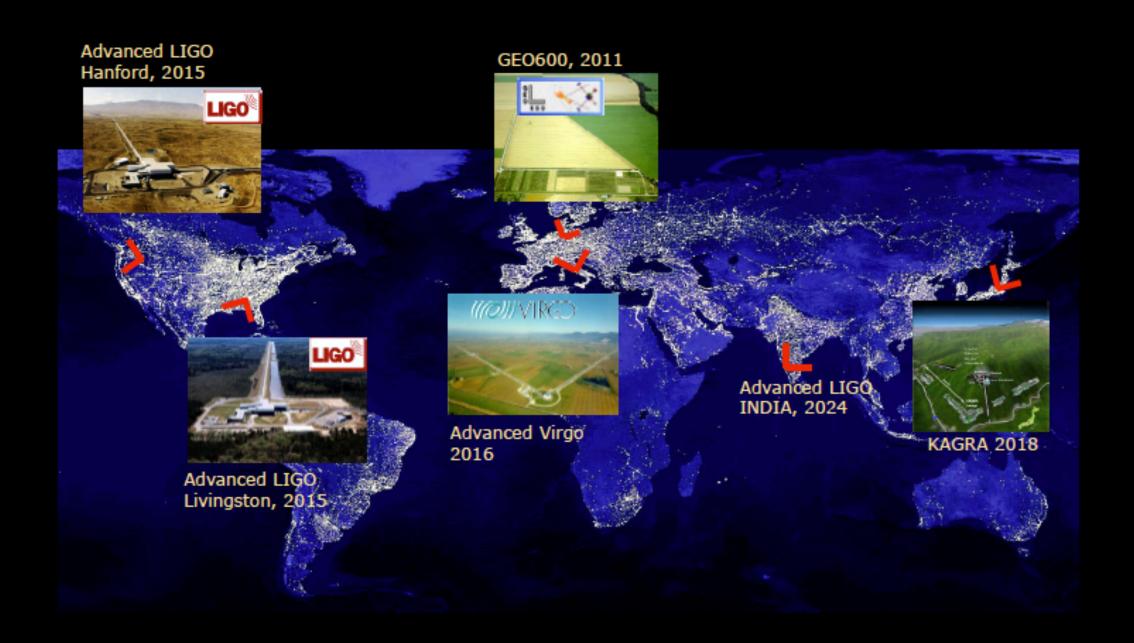


IFAE engineers installing instrumented baffle at Virgo

(A demonstrator of the novel technology... not for this talk)

# LVK sensitivity





In the next 6 years the current Interferometers will reach their design sensitivity...

Next Gen Network Cosmic Explorer

Einstein Telescope

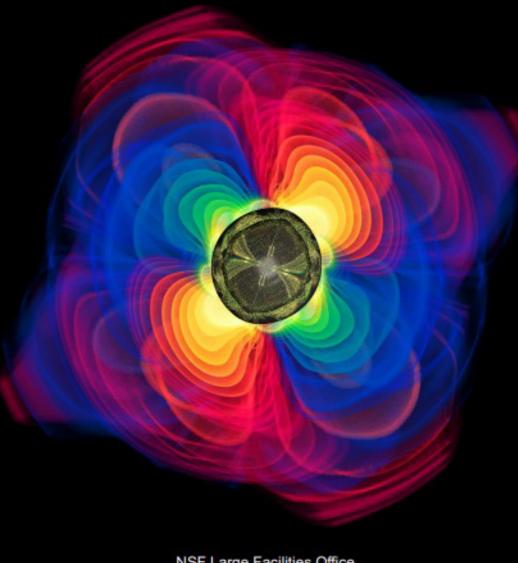


# Cosmic Explorer (USA)



#### RESEARCH INFRASTRUCTURE GUIDE

NSF guidance for full life-cycle oversight of Major Facilities and Mid-Scale Projects



NSF Large Facilities Office
Office of Budget, Finance and Award Management

NSF 21-107 December 2021

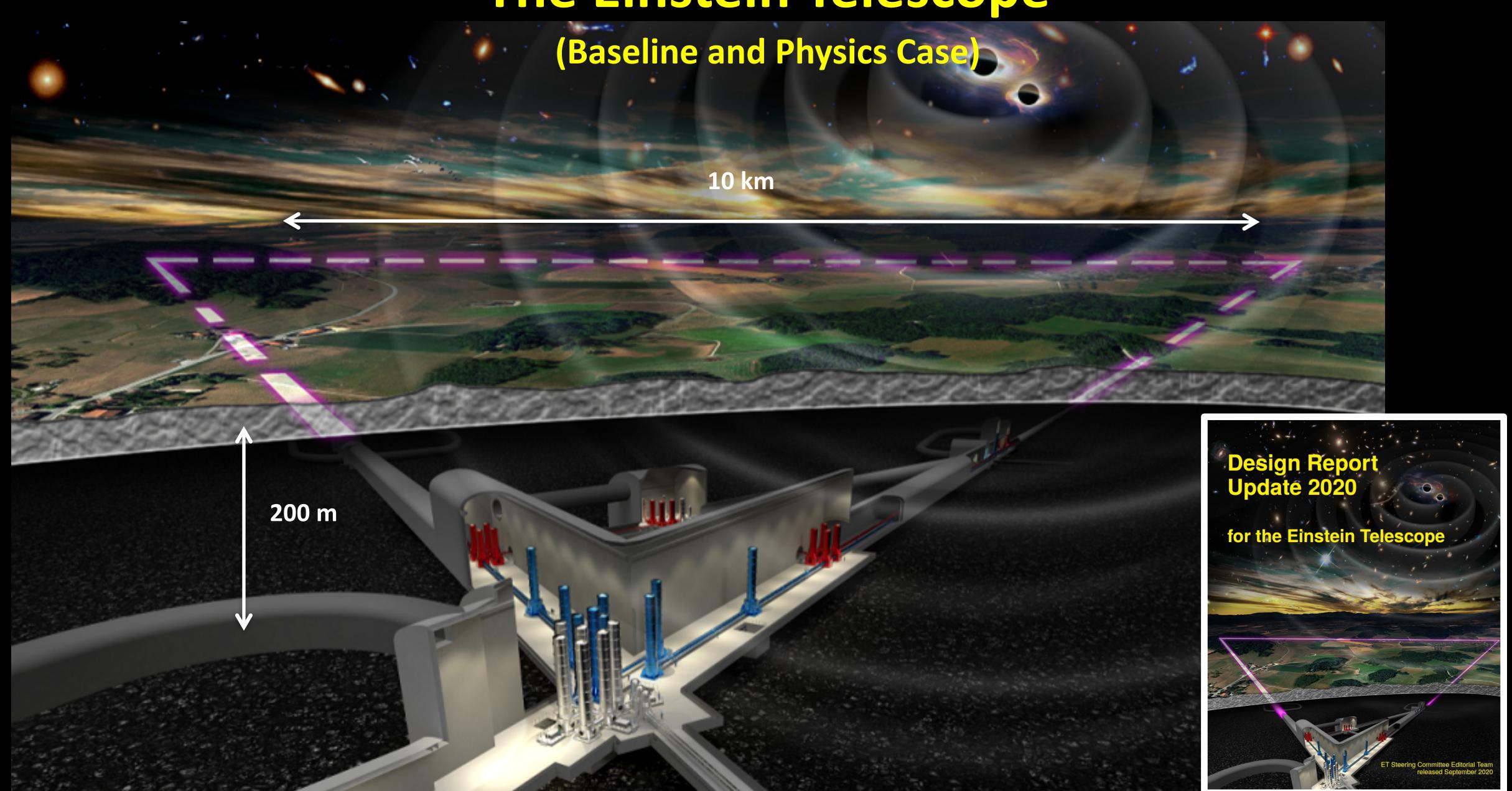
Credit: Scientific contact by Ed Seidel (eseidel@aci.mpg.de); simulations by Max Planck Institute for Gravitational Physics (Albert-Einstein-AEI); visualization by Werner Benger, Zuse Institute, Berlin (ZIB) and AEI. The computations were performed on NCSA's It.

http://dcc.cosmicexplorer.org/CE-P2100003/public

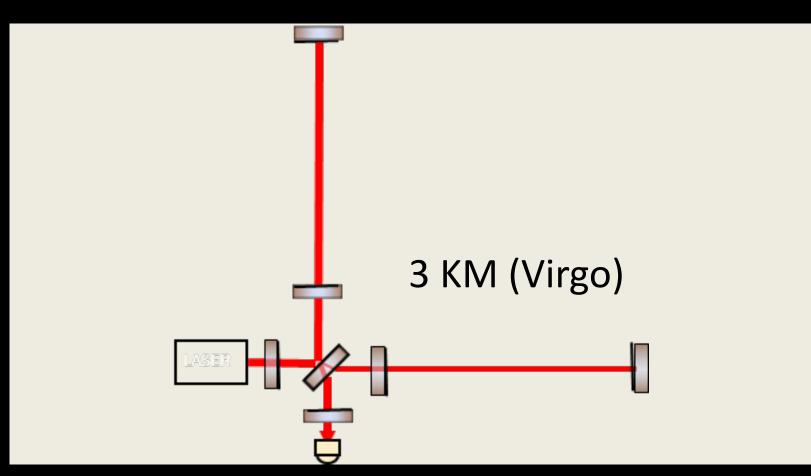
http://www.et-gw.eu/

# The Einstein Telescope

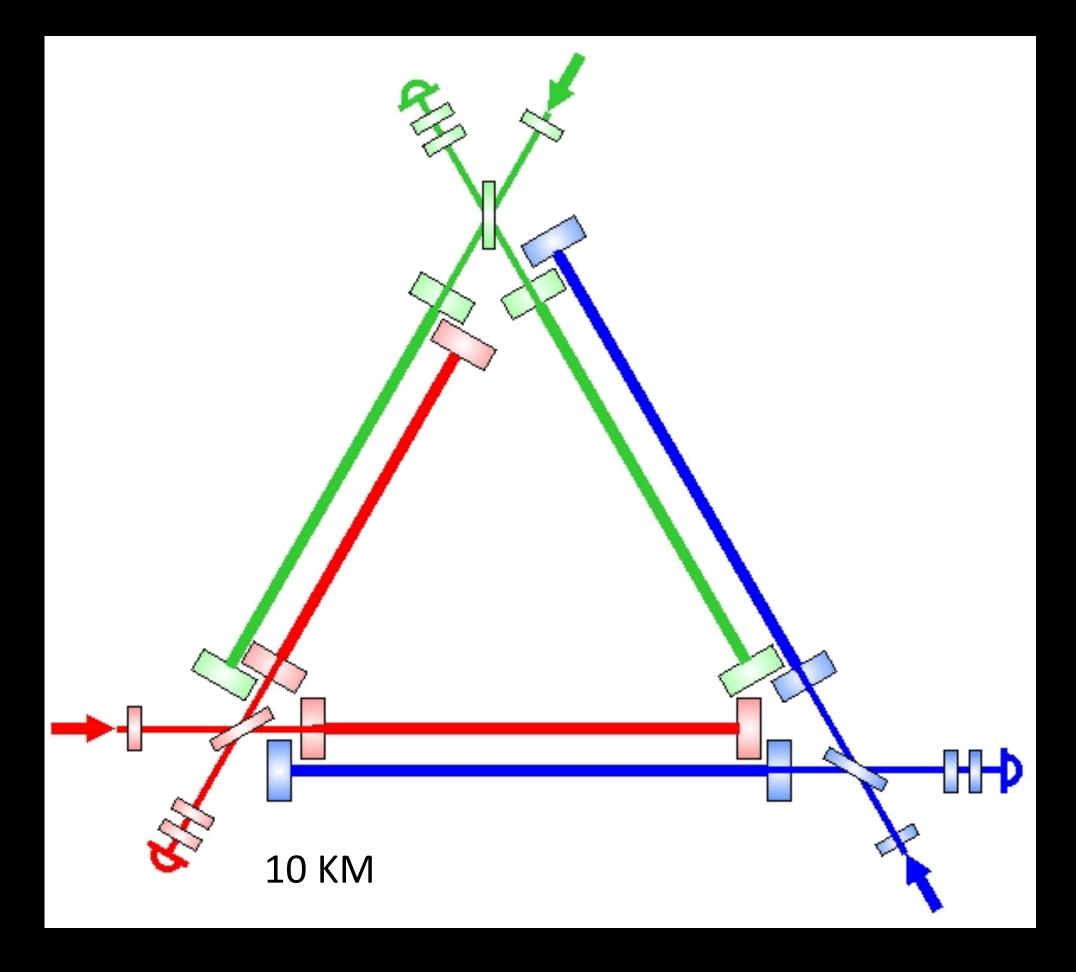
First CDR in 2010 - 2011

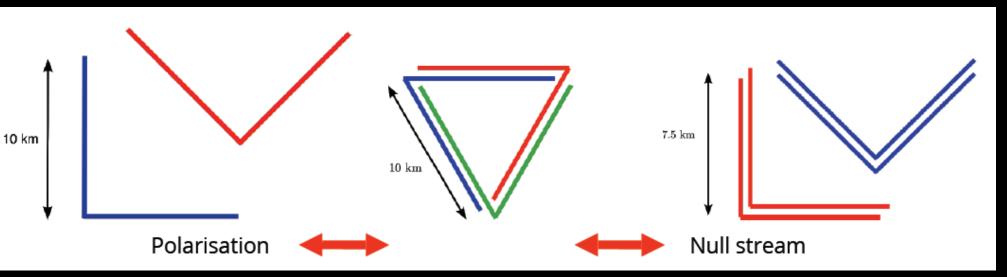


## 2nd Gen $\rightarrow$ 3rd Gen

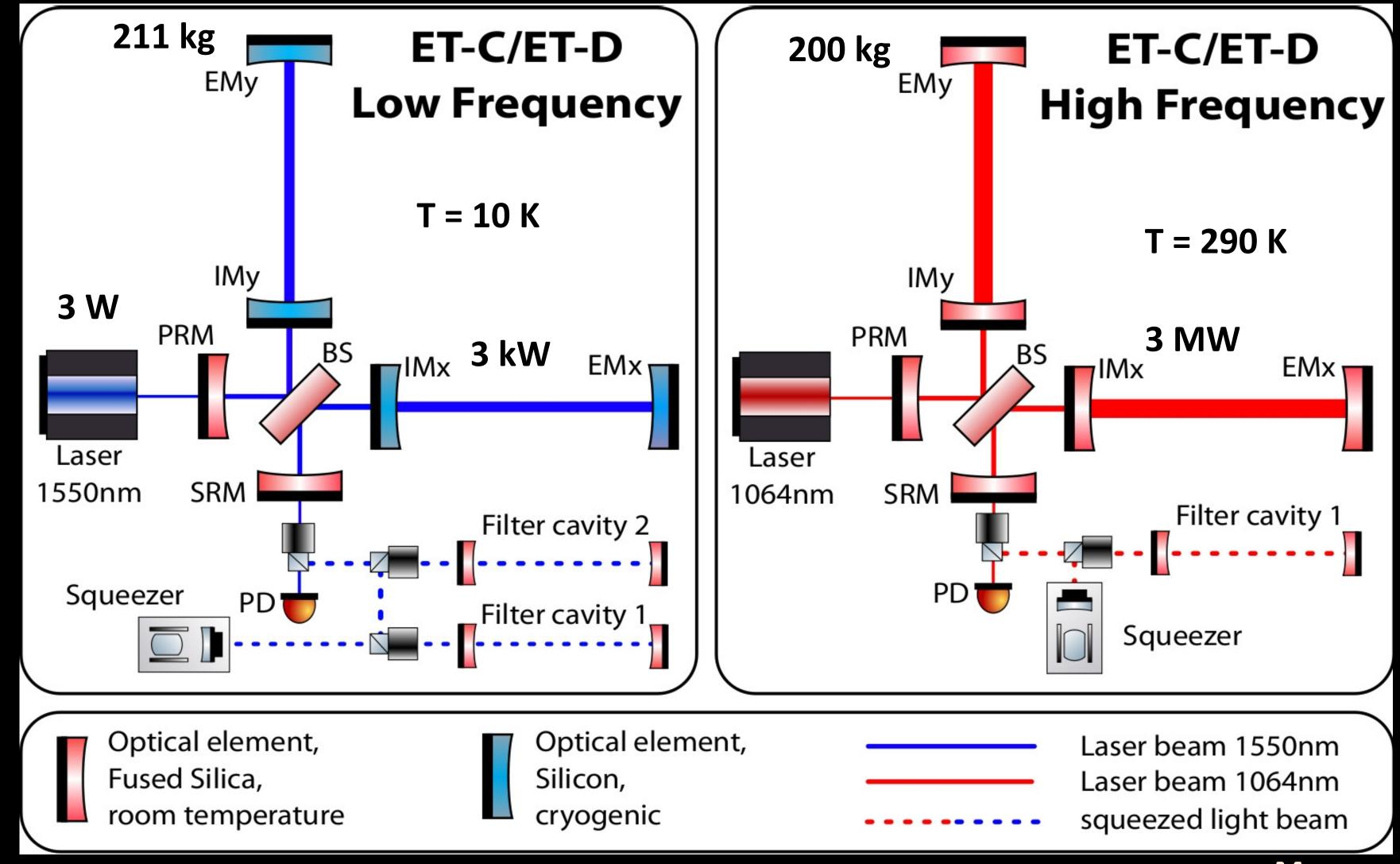


- Underground to suppress seismic and Newtonian noise at low frequencies
- Longer arms to increase sensitivity
- Triangular ITF configuration to
  - Measure polarization
  - Auto-calibration / null-stream
  - Redundancy
  - One single big infrastructure (optimization of costs)





2nd Gen -> ET



Underground Cryogenic Silicon mirrors

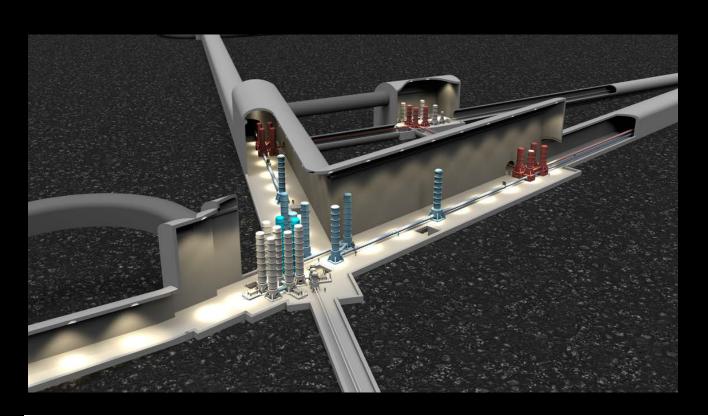
1550 nm (Si transparent)

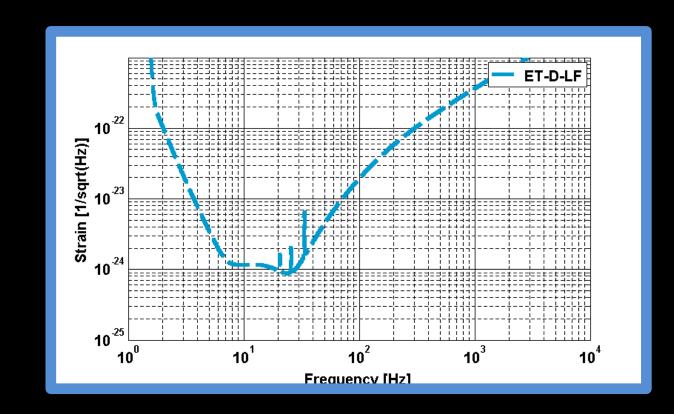
New optical coatings

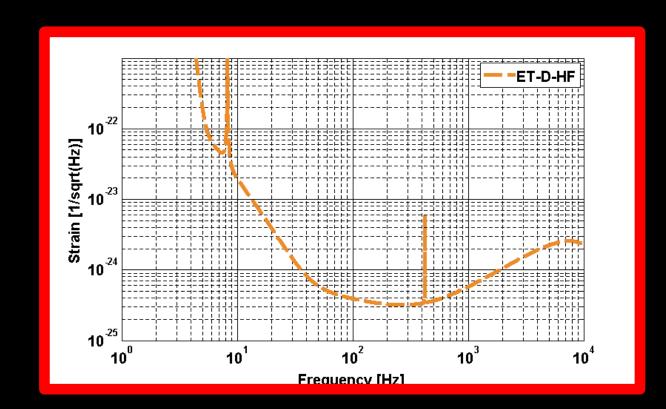
New suspensions / seismic controls

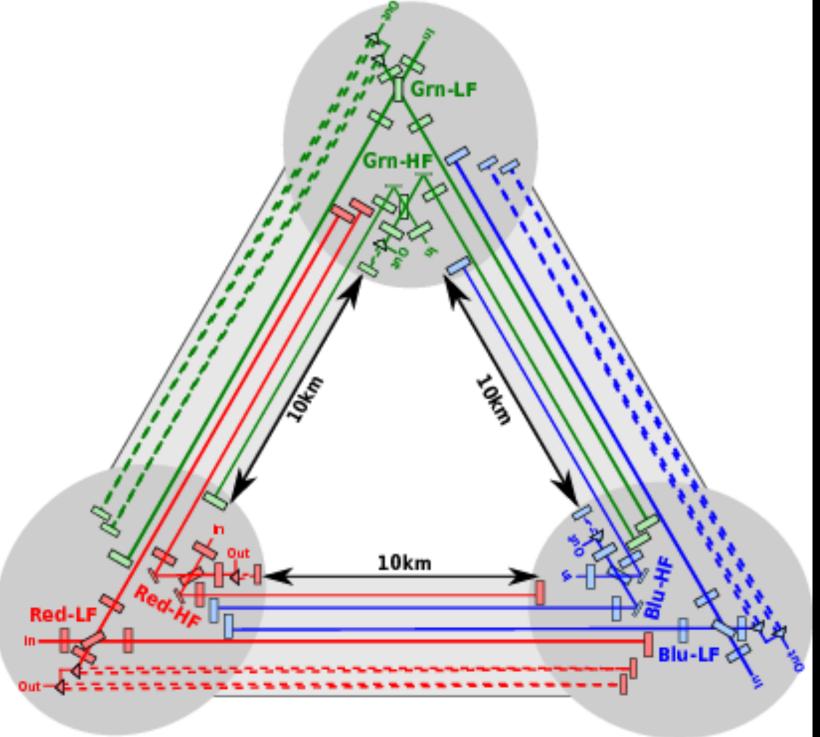
More powerful lasers
Larger fused silica mirrors
1064 nm (silica transparent)
New optical coatings
New thermal compensation systems

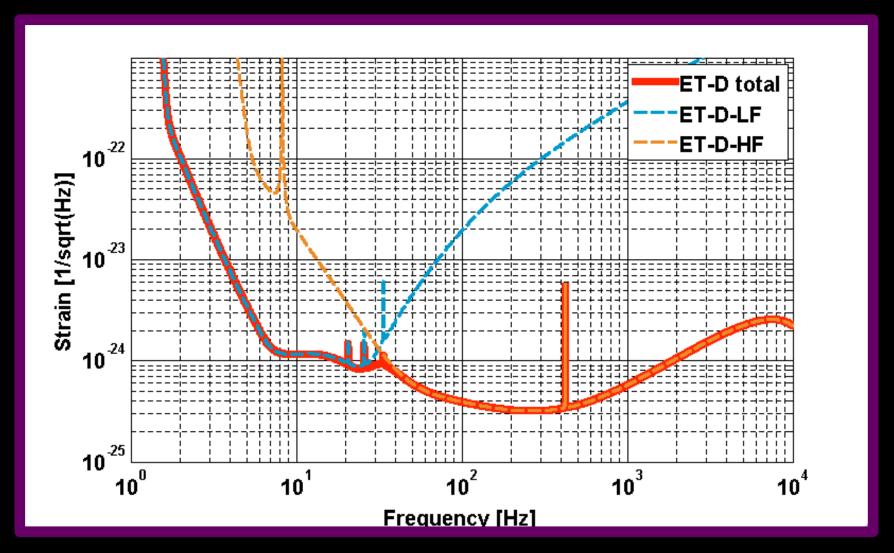
# Einstein Telescope (6 in 1) Xylophone











Each interferometer decoupled into 2 devices independent for the best sensitivity to low and high frequency

## ET Science

arXiv:1912.02622

ET Science Blue Book will be ready by September 2024

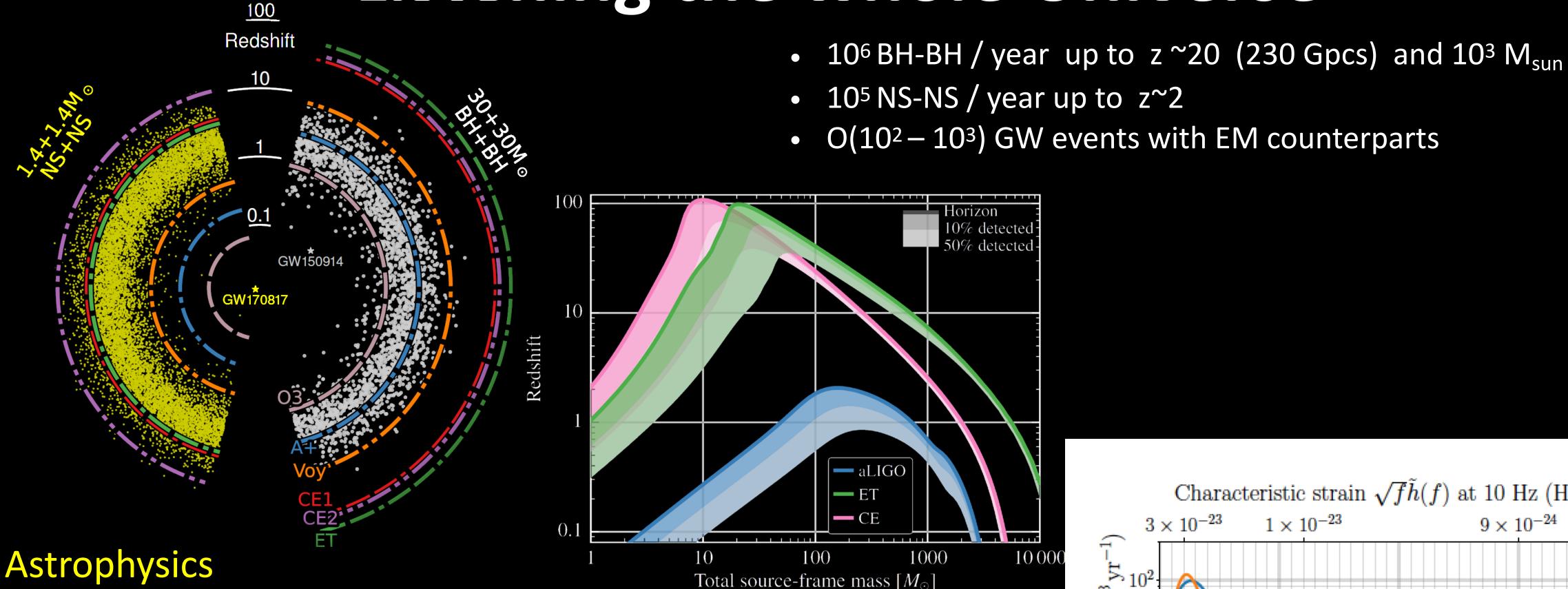
#### Astrophysics

- Black hole properties
- origin (stellar vs. primordial)
- evolution, demography
- Neutron star properties
- demography, equation of state
- Multi-messenger astronomy
- joint GW/EM observations (GRB, kilonova,...)
- multiband GW detection (LISA)
- Detection of new astrophysical sources
- core collapse supernovae
- isolated neutron stars
- stochastic background of astrophysical origin

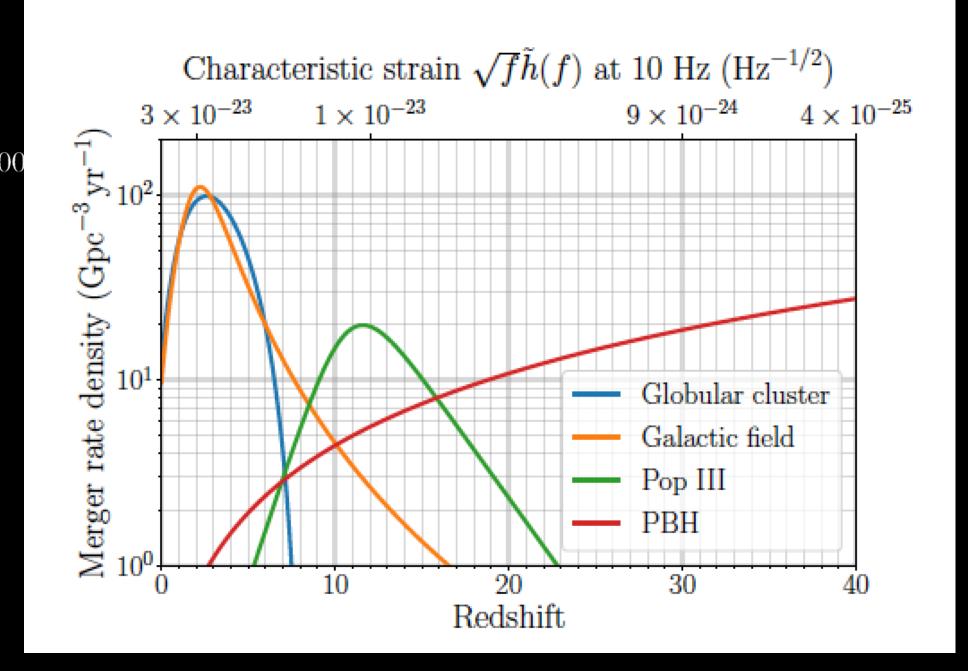
#### Fundamental Physics, Cosmology, HEP

- Testing the nature of gravity
- perturbative regime: inspiral phase of BBH, post-Newtonian expansion
- strong field regime: physics near BH horizon exotic compact objects
- QCD
- interior structure of neutron stars probe ultra-high temperatures and densities
- exotic states of matter
- Dark Matter / New Particles
- primordial BHs
- axions, dark matter accreting on compact objects
- Modified Cosmology
- Dark Energy equation of state
- modified GW propagation @ cosmological scales
- Stochastic backgrounds of cosmological origin and HEP
- inflation, first-order phase transitions
- cosmic strings, domain walls...

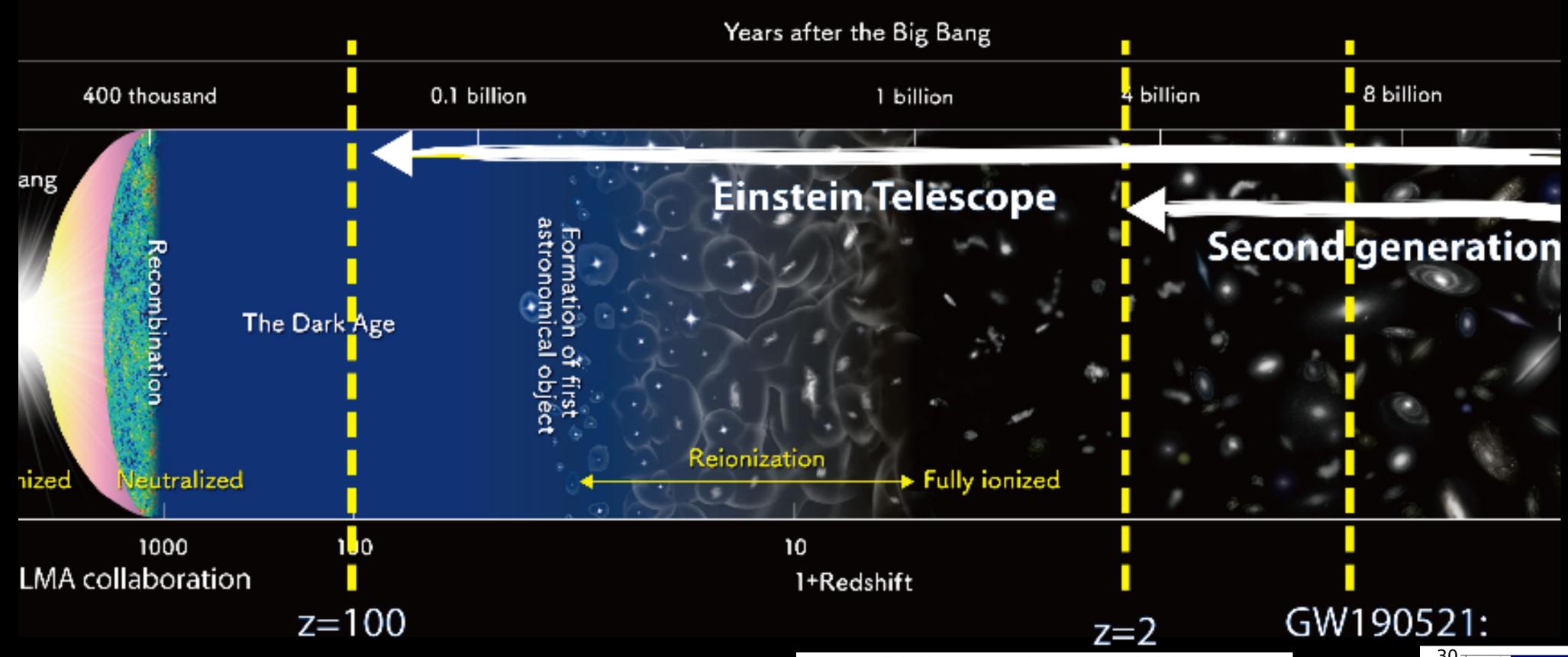
# Listening the whole Universe



- BH demography and evolution
- Primordials? Stellar?
- Are BHs part of the dark matter?
- Supernovae, Pulsars, Stochastic signals
- Properties of neutron stars
- Multi Messenger: Optical, Neutrinos, Gamma Rays



#### Detection horizon for black-hole binaries

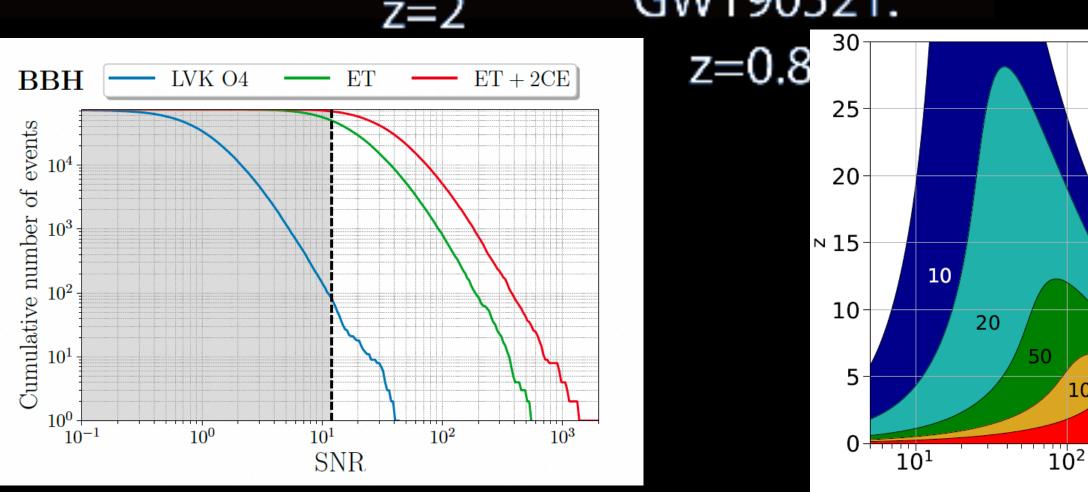


Huge rate of detections (about 1 per minute) Extended redshift coverage up to the Dark Age

- Test for primordial BH origin
- Cosmology & Cosmography

Many events with very large Signal-to-Noise ratios

Precision tests of GR predictions and detailed BH studies

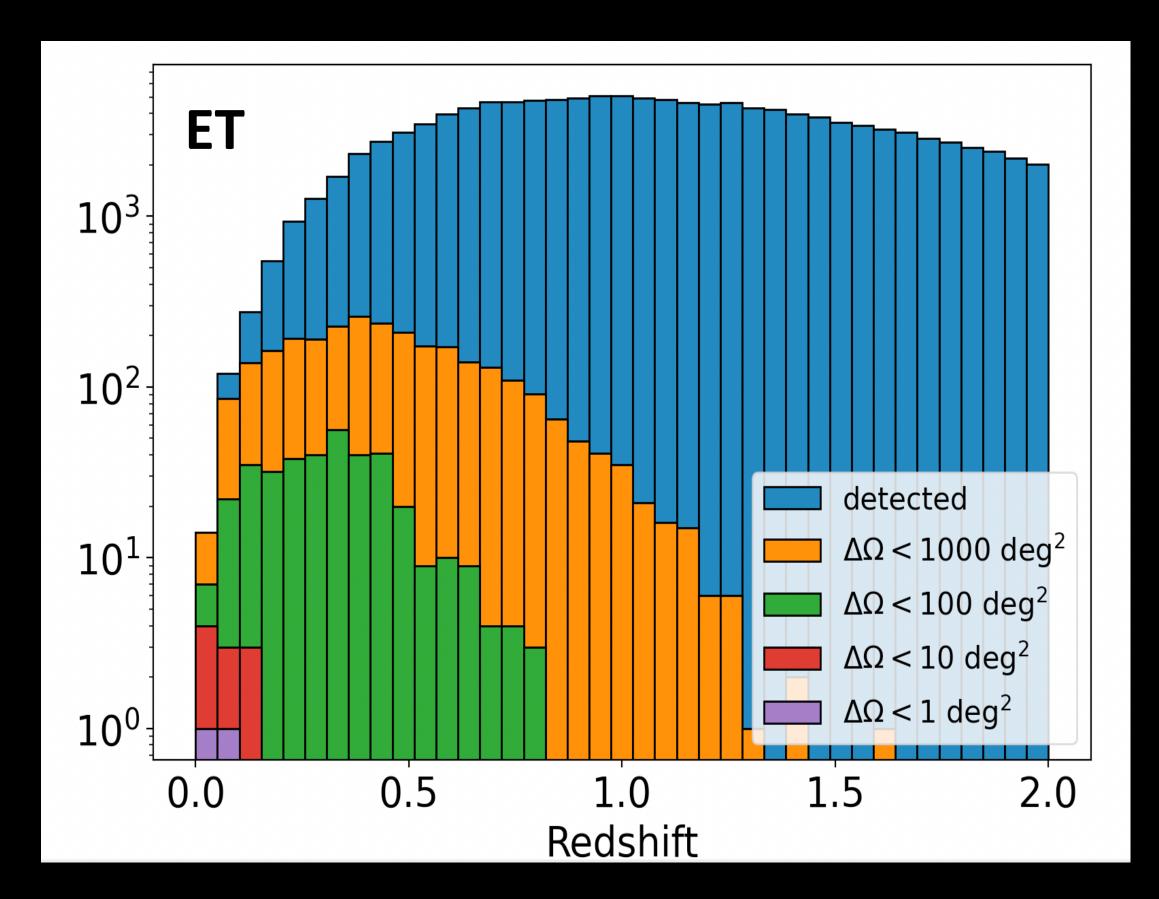


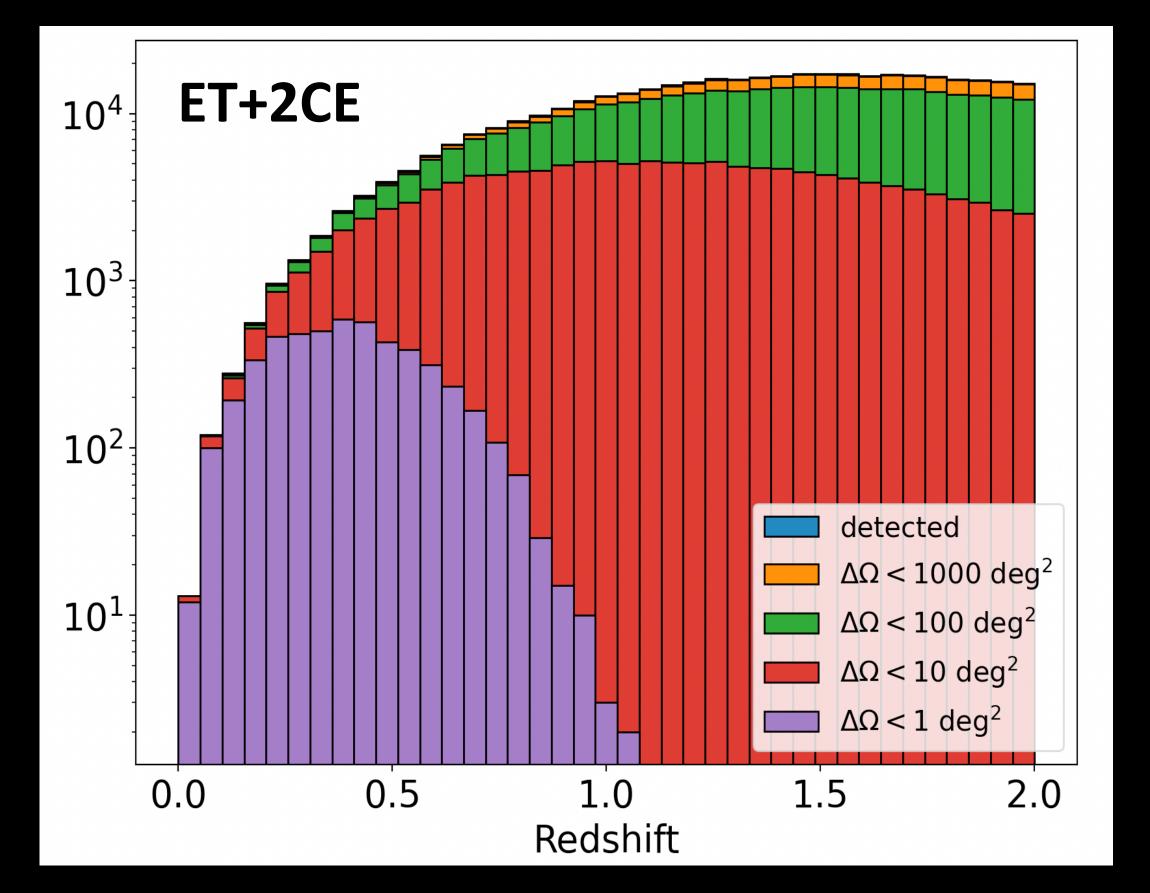
103

 $M(M_{\odot})$ 

 $10^{4}$ 

## Sky localization

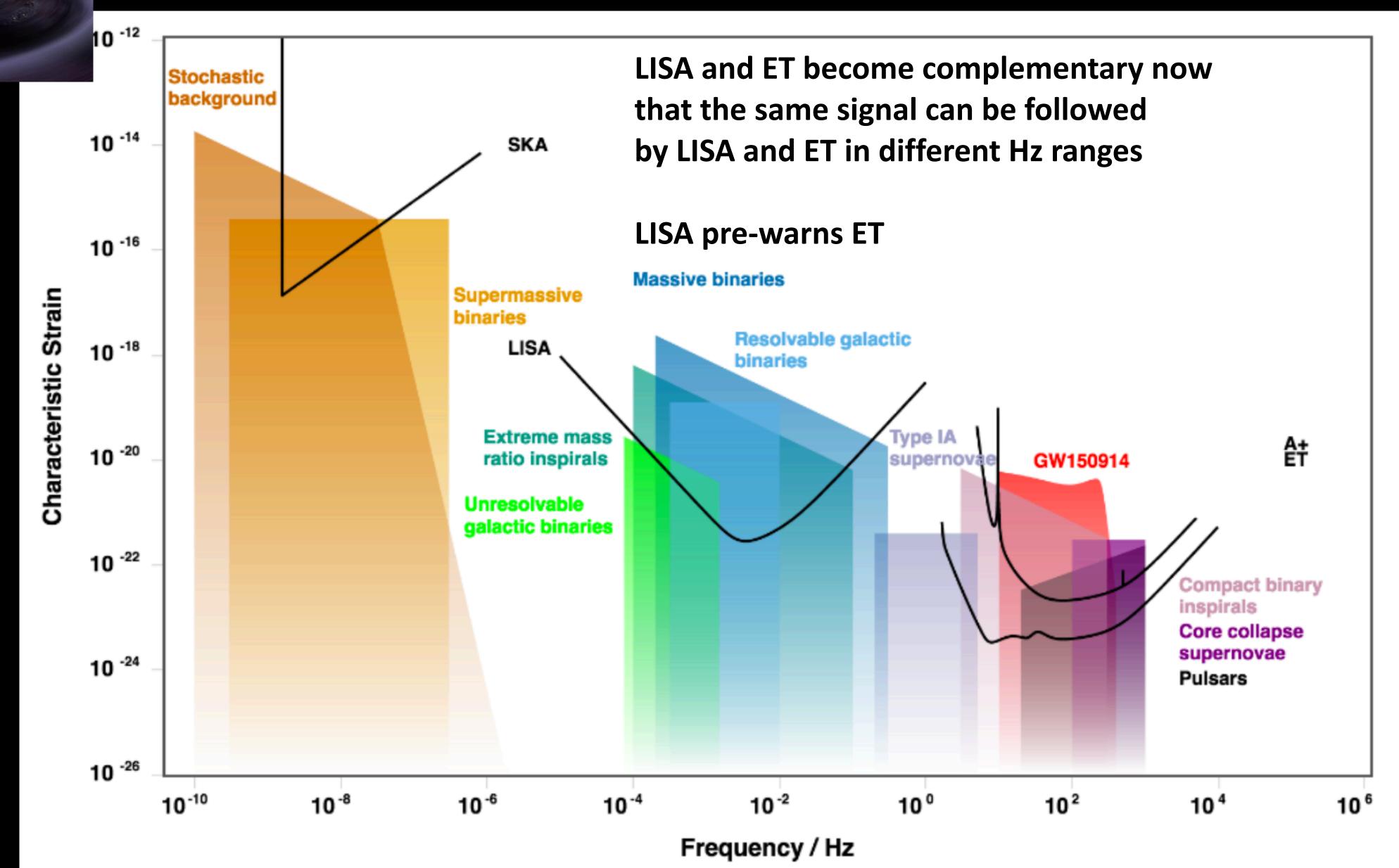




ET only configuration would allow for O(100) events / year with a sky-localizations (90% CL) < 100 deg<sup>2</sup>

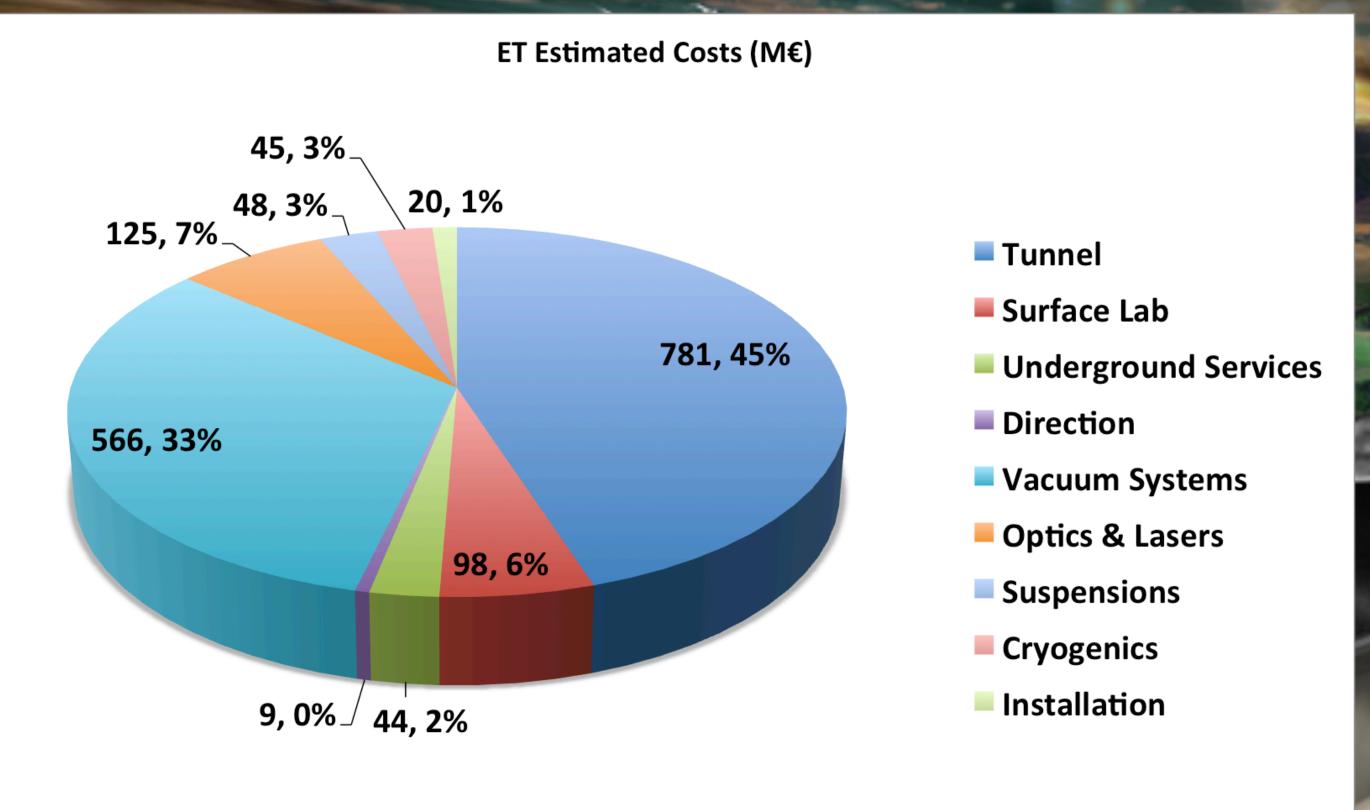
ET + 2 CE configuration would allow for O(1000) events / year with a sky-localizations (90% CL) < 1 deg<sup>2</sup>

# Complementarity



http://www.et-gw.eu/ The Einstein Telescope (As EU singular infrastructure) 10 km Design Report Update 2020 200 m for the Einstein Telescope

# Estimated cost (EU accounting)



#### Preparatory phase (170M€)

- 1. Site qualification (funded)
- 2. Site preparation (50 60 M€)Covered by host country
- 3. R&D on technology (95 M€) (funded)

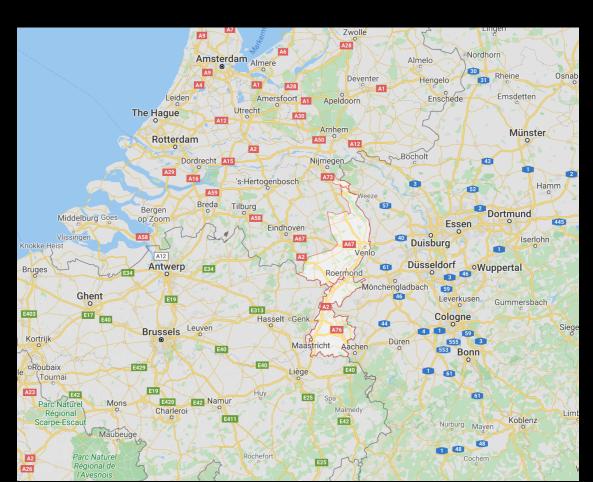
Construction: 1900 M€ (in 10 years)

M&O:37M€/year

Host country is expected to contribute with > 50% of the total cost

#### O(50 M€) investment Lab in construction

## Locations?



30 M€ investment ETparthfinder

Intensive studies

- @ Limburg,
- @ Sardinia
- @ Saxony

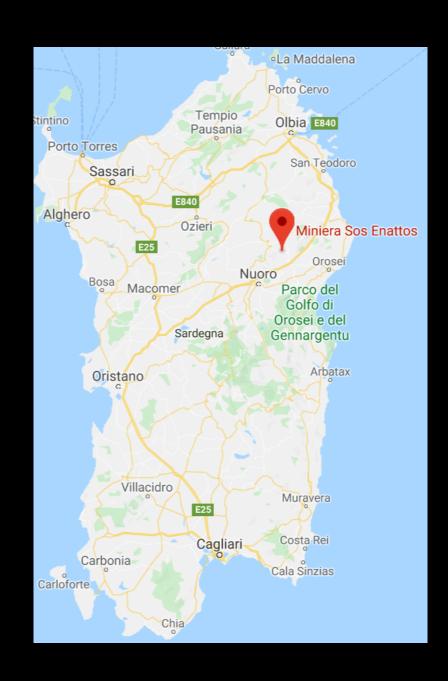
For characterize seismic, environmental noise, etc ...

- @ Limburg area (border NL-B-D)
- Promoted by Nikhef



- @ Germany is very present in ET and ETpathfinder They foresee a large investment in the following years
- → Exploring Saxony as a possibility
- → Ongoing geological characterization of the site

#### O(50 M€) investment Lab in construction



- @ Sardinia
- → Promoted by INFN



## Rising Construction Funds

In the Netherlands a formal request of 900M€ for ET@ Maastricht has been approved by the Science Minister to the NL Government

Italy approved a 50M€ project for enabling technologies and additional 950M€ for supporting ET@ Italy has been secured

Time to discuss the level of financial involvement by other EU countries in ET for the following decade

#### Einstein Telescope in Euregio Meuse-Rhine (EMR)



Connected institutions in: Belgium, Germany & the Netherlands

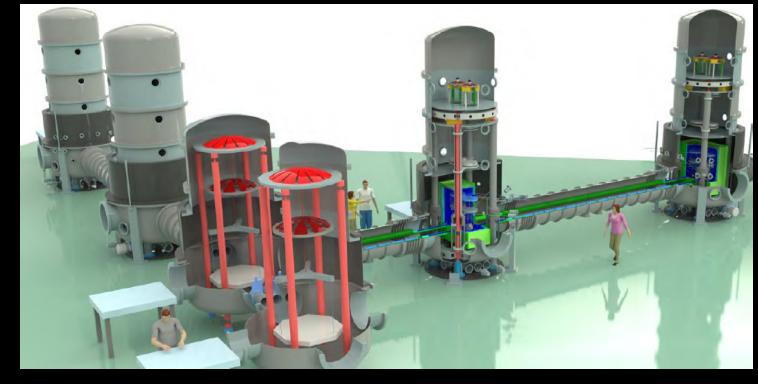


as well as possible Dutch share towards ET realization

#### ETIC – Einstein Telescope Infrastructure Consortium



# Ongoing R&D for ET (examples)





@ Maastricht small-scale prototype in order to study the operations in cryogenics with silicon optics at 1550 nm and with mirrors up to 100 kg (relevant for ET-LF)



R&D in active mitigation of seismic / Newtonian noise

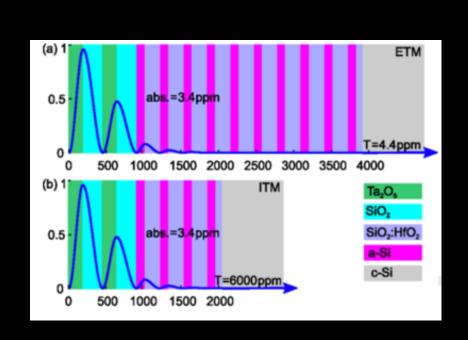
R&D on IR photo sensors

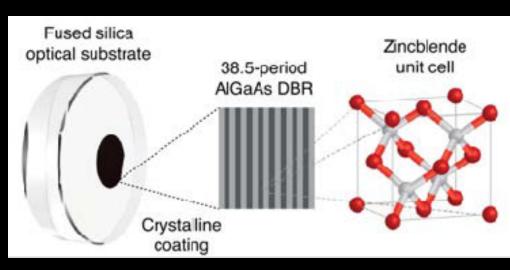
@ Rome and other R&D places in cryogenic suspension

R&D for the production of mirrors up to 200kg based on silica or silicon of high purity and homogeneity.



R&D in optical coatings reflective with low absorption and reduced thermal noise.





AlGaAs crystalline coatings

E T EINSTEIN
TELESCOPE Major R&D Facilities in ET (incomplete) Glasgow • Careful shielding for thermal fluctuations, acoustics, and E-M tpathfinder in Maastricht Hamburg Interreg Vlaanderen-Neder Main goal: Sub-SQL interferometry provide a testbed for ET technol Studies of vibration isolation / control concepts and qualify them in low 2 FPMI interferometers Hannover Amsterdam Maastricht incomplete Louvain -TEST : Einstein Telescope EMR Site and Technology lest facility for experimental investigations of the He-II based ET-LF payload cooling concept Karlsruhe See Monday talk by MA - Laboratoire des Matériaux Avancés LAUS: Centro per Applicazioni sulle Unde gravitazionali Test the limits of active seismic isolation in an underground environment Inter-platform motion control Underground environmental monitoring Test new approaches to controls optimization SAR-GRAV Laboratory 🏥 FSC 👊 🐿 uniss 🖬 IGEA SPA 🕼 📽 SAR-GRAV hosts ET activities as well as Sos Enatos It is planned to test at least O Cavern that should host the partially a preliminary version of Archimedes experiment the double-suspended inverted pendulum in a quiet Q underground environment. O See Monday talk See Monday talk by Ettore Majorana by Enrico Calloni U.S. Navy, NGA, GEBCO Google Earth

selection

the

From

## ETpathfinder(s)

**ETparthfinder** 



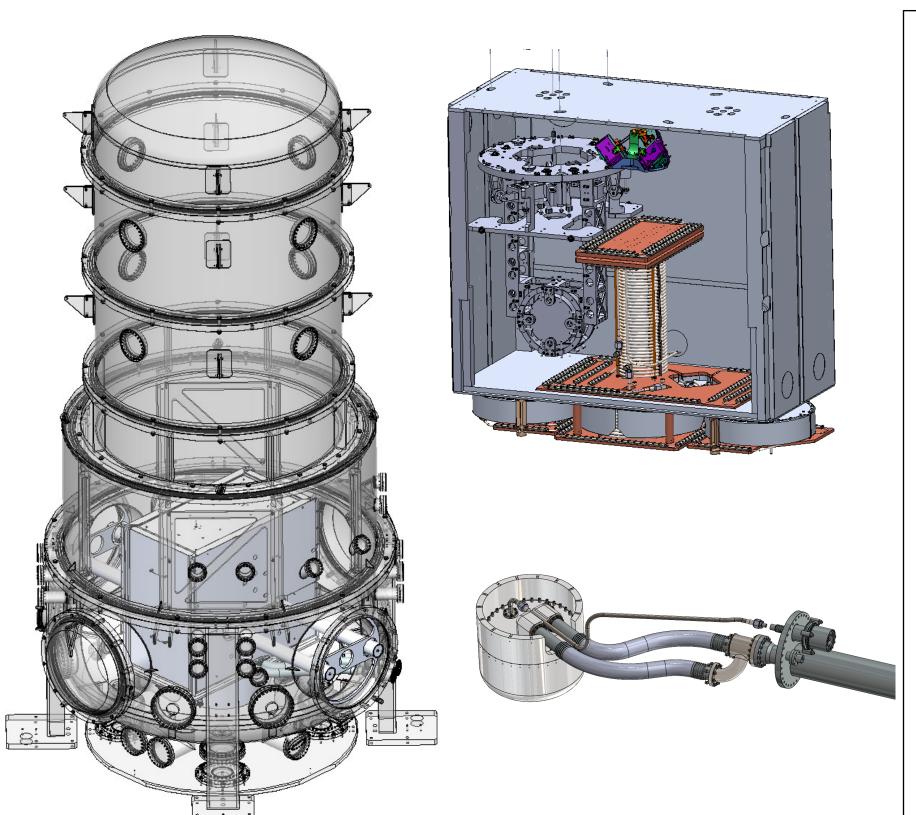
A collaboration established with Etpathfinder @ Maastricht

PC will cover partially IFAE redesigned the cryo-shielding [paid by Nikhef] the hardware cost at

à IFAE will contribute to its installation in 2024

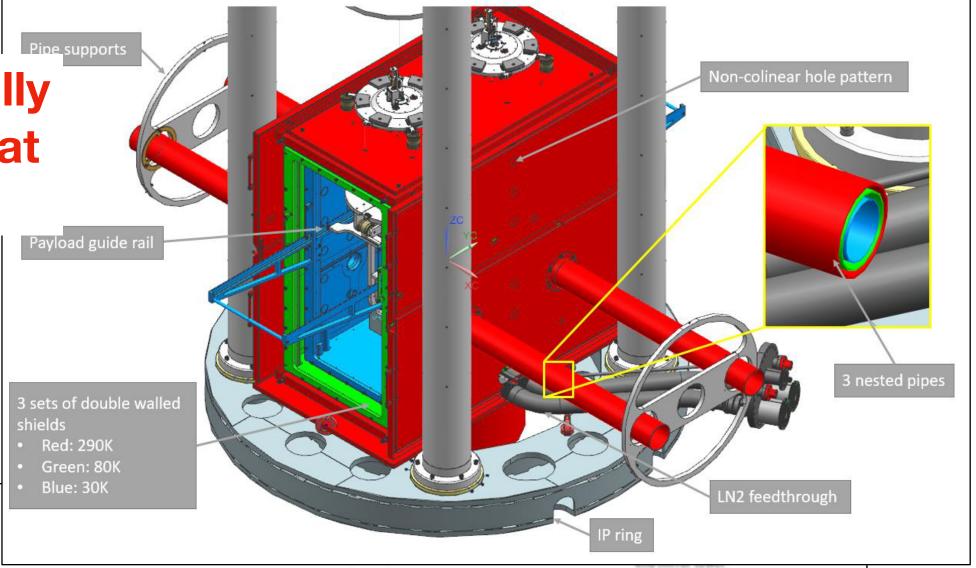
à Pre-alignment & monitoring of the mirror surface at 80 - 290 K (instrumented baffle with sensors for  $\lambda$  = 1550 nm )

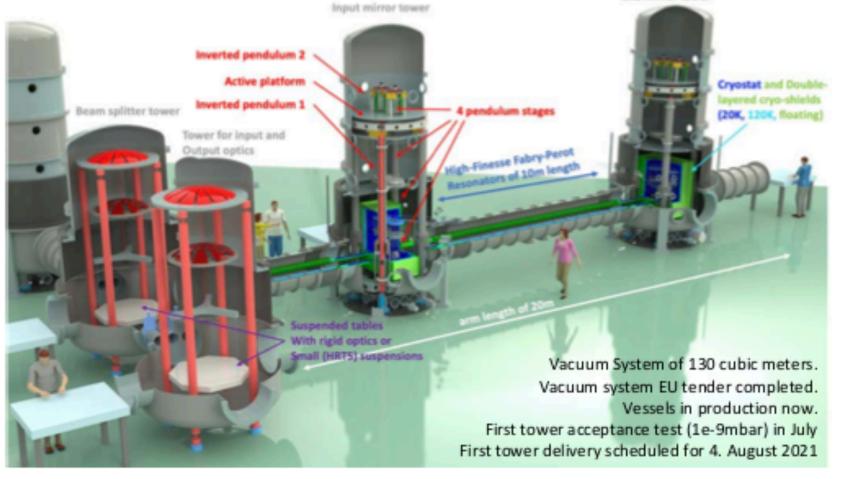
R&D on InGaAs photosensors with AR-coatings will start soon



## ETpathfinder

- New facility for testing 3G technology in a low-noise, full-interferometer setup.
- Key aspects: Silicon mirrors (3 to 100+kg), cryogenics (cryogenic liquids and sorption coolers, water/ice management), "new" wavelengths (1550 and 2090nm), new coatings ...
- Start with 2 FPMI, one 120K and one 15K.
- 16 official partners from NL/B/G/FR + a few more involved, but not yet official partners (like AEI, KIT, Bham, Cardif, Barcelona etc).
- Initial capital funding of 14.5 Meuro (no personpower).
- Detailed Design Report available at apps.et-gw.eu/tds/?content=3&r=17177
- Open for everyone interested to join.
- www.etpathfinder.eu









Slide from Nikhef SAC Meeting

Nikhef SAC, 2021-05-26

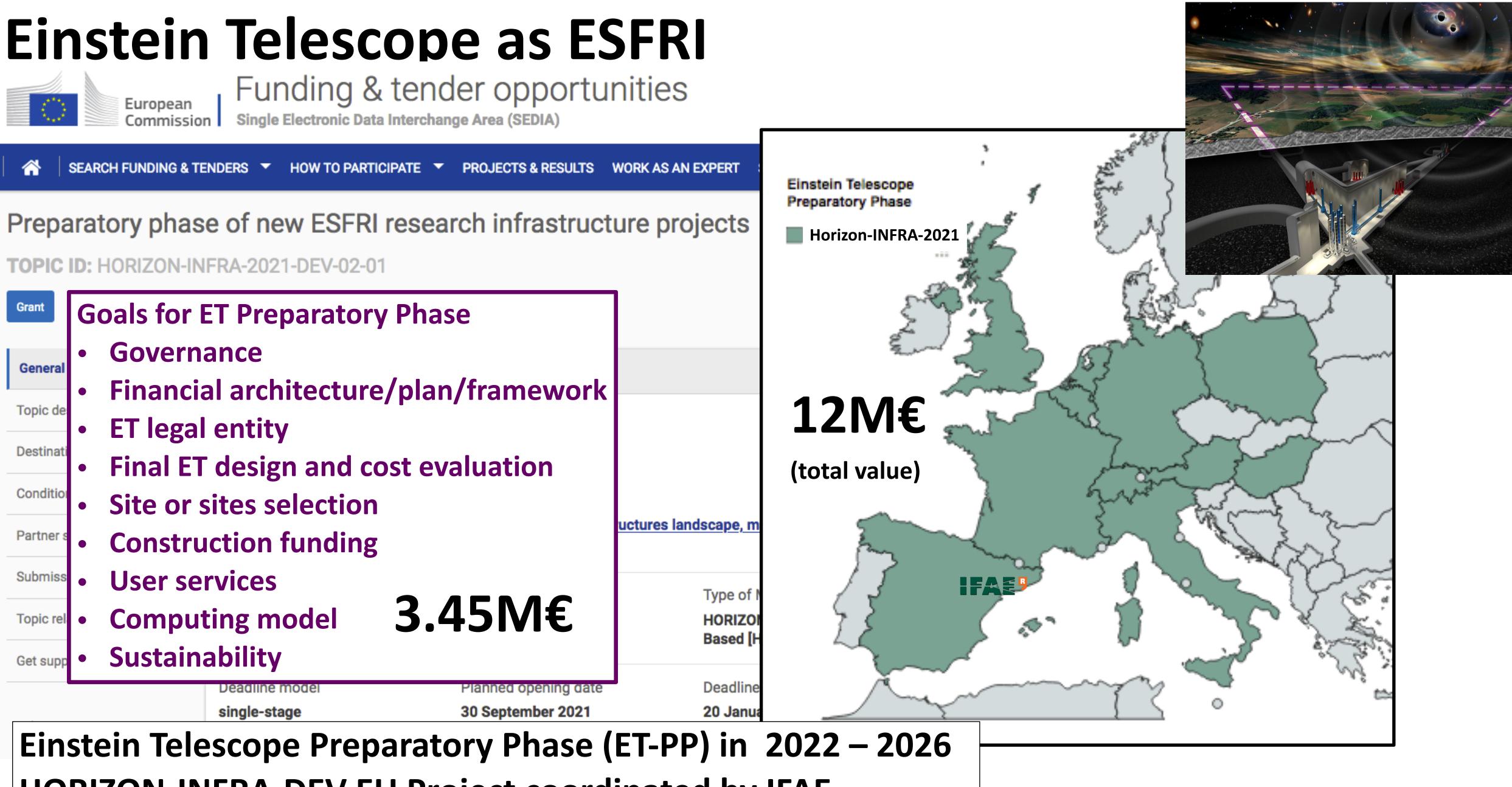
[S.Hild]

# ET on the ESFRI roadmap ET THESCOPE



## ET Consortium ET CA signed by 41 institutions INFN and Nikhef are the coordinators of the consortium Funding expected in the next months by the governments in the frontline EU funding for the Preparatory Phase in 2022 Irlanda Polonia ( Germania Repubblica Portogallo Spagna Grecia

Google



**HORIZON-INFRA-DEV EU Project coordinated by IFAE** 

Project started 1st September 2022 (https://etpp.ifae.es)

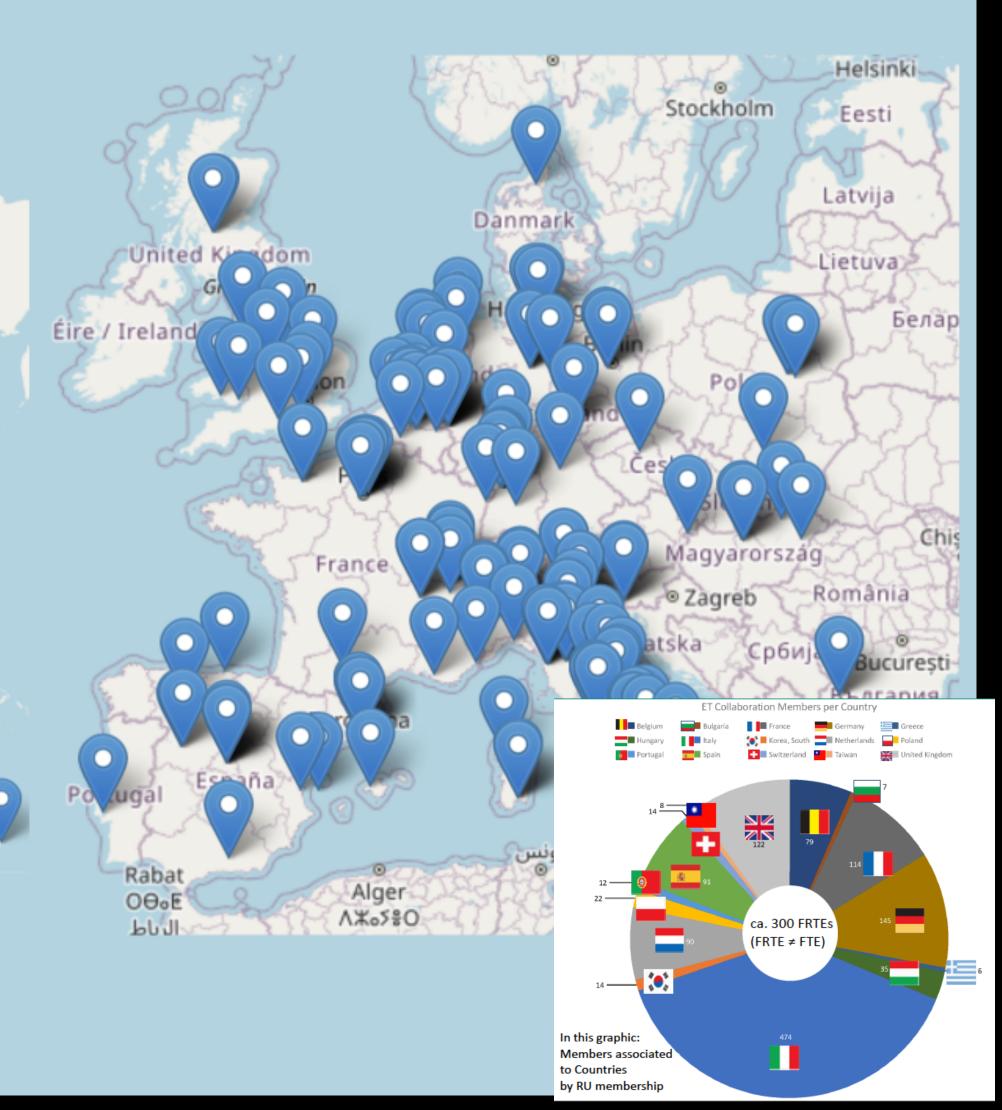
## The Einstein Telescope Collaboration



- 85 Research Units (+1 request pending)
- 1568 members (24/11/2023 15:29)
- Total: 226 Institutions in 25 Countries

Large Collaboration

ET member database



- A Large Collaboration (comparable to a LHC experiment)
  Requires a proper Governance /Financial Model
  - Internal organisation in place
  - Bylaws already in place

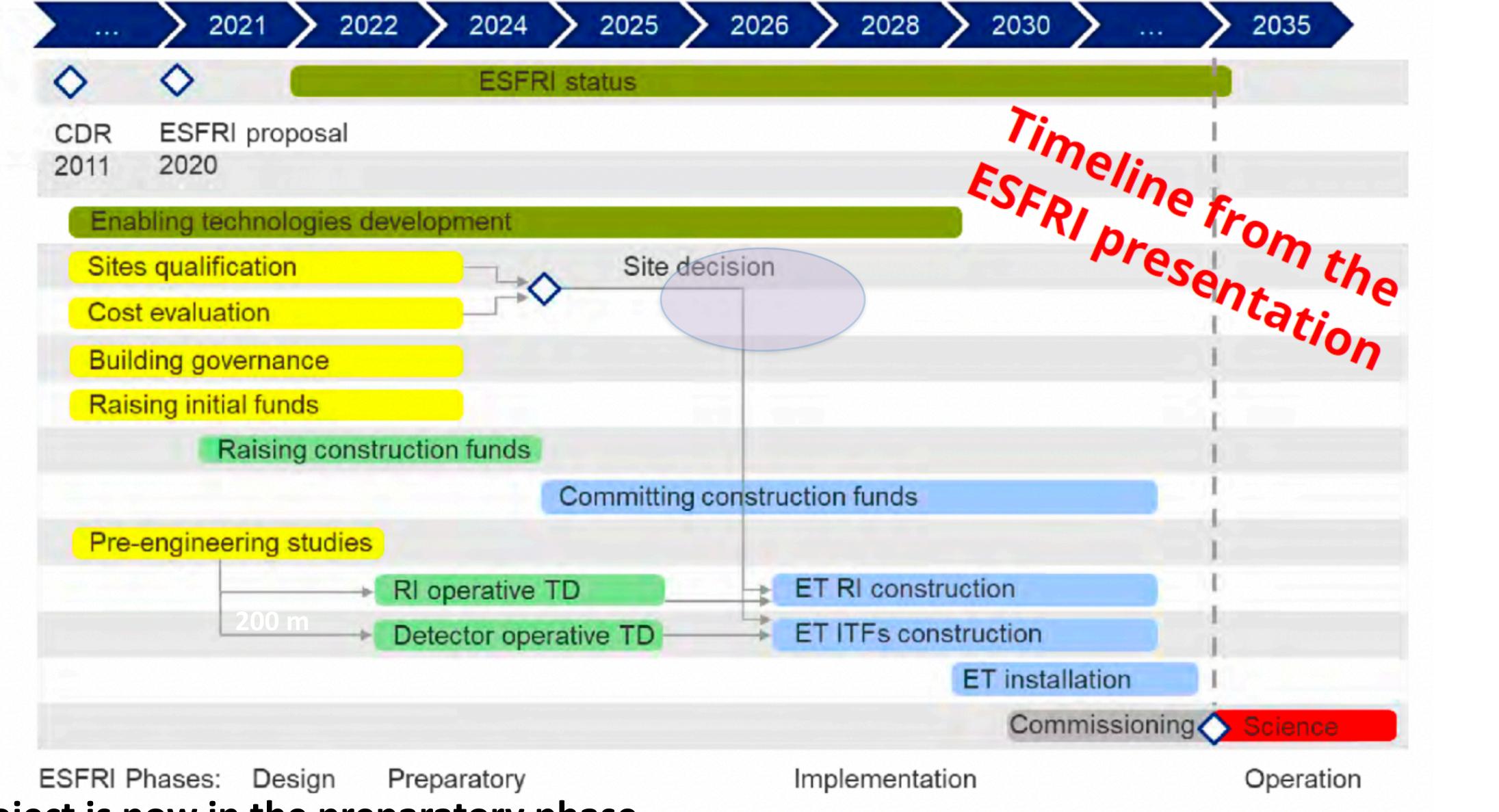
# ET experiment and ET Organization

ETO (ET Organisation) Board of governmental representatives **Board of Scientific** Coordinators representatives **ET Directorate** Engineering Project office department Administrative ET-PP office

ET Collaboration Collaboration board **Executive board** Specific boards ISB OSB **EIB** SCB SSB

ET-PP is naturally framed inside ETO chart

ET research infrastructure, services, and vacuum system under ETO supervision ET Scientific Collaboration dedicated to experiment design & scientific exploitation

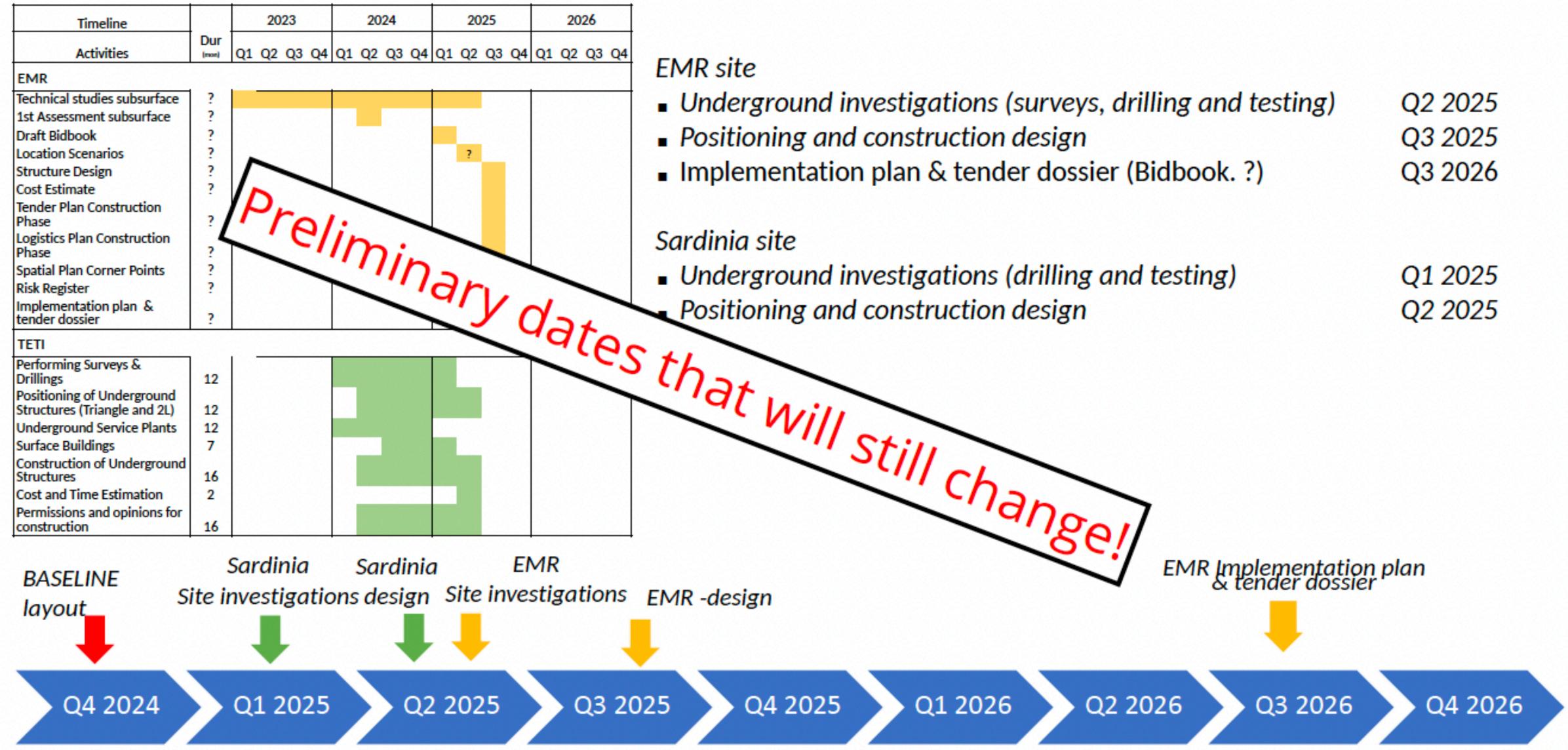


#### ET project is now in the preparatory phase

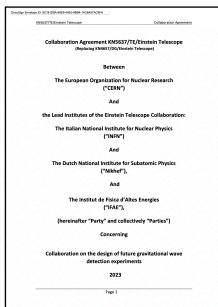
- —> We know already ESFRI roadmap schedule was too optimistic based on simplified assumptions
- —> We are in the process to define a new schedule and roadmap —> to be ready by early 2025

## EMR-TETI civil engineering - Timelines

## Example 1: Civil engineering (with CERN)







The Einstein Telescope is a 'recognised experiment' at CERN. We can access support provided it is cost neutral to CERN.

- **Vacuum pipe**: governed by an MOU between CERN, INFN, Nikhef and IFAE, CERN has started a dedicated activity to deliver the technical design report for the vacuum pipe in 2025. Currently a prototyp station is being built at CERN, supported by a large and efficient team.
- **Civil engineering**: an extension to the MOU has been agreed: CERN will provide consultancy and technical support towards the creation of the TDR for the civil engineering and technical infrastructure for 2026.
- **Document management**: project management requires specific tools, we are investigating the use of the CERN tool EDMS. CERN is providing support for a pilot study which has now started.
- **Engineering support**: technical designs at CERN are usually done by a large interdisciplinary team, including for example the safety group. We have organised several discussions with the relevant teams and are now formulating a plan for dedicated support for the design of the technical infrastructure.



Manufacturing

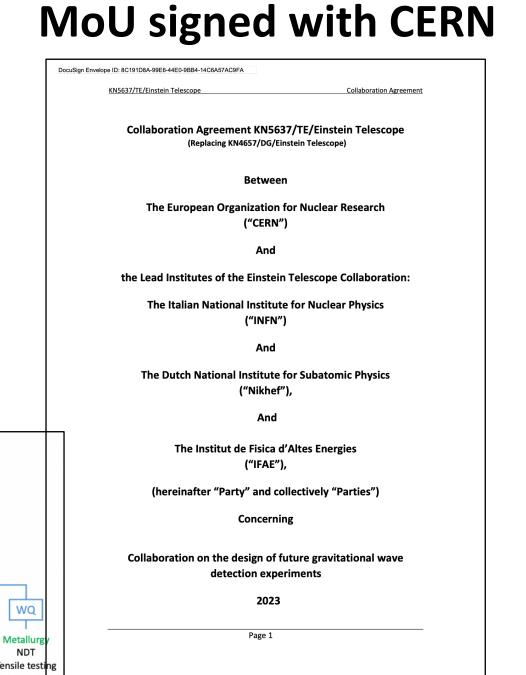
316LN end covers

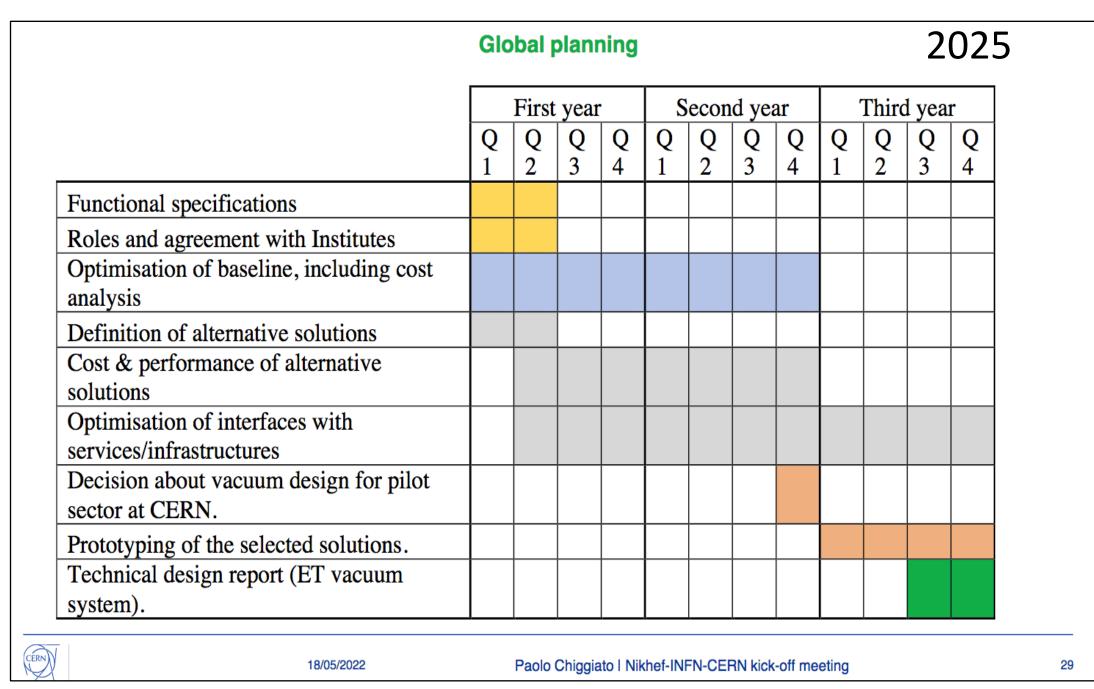
## ET & CE vacuum pipe design

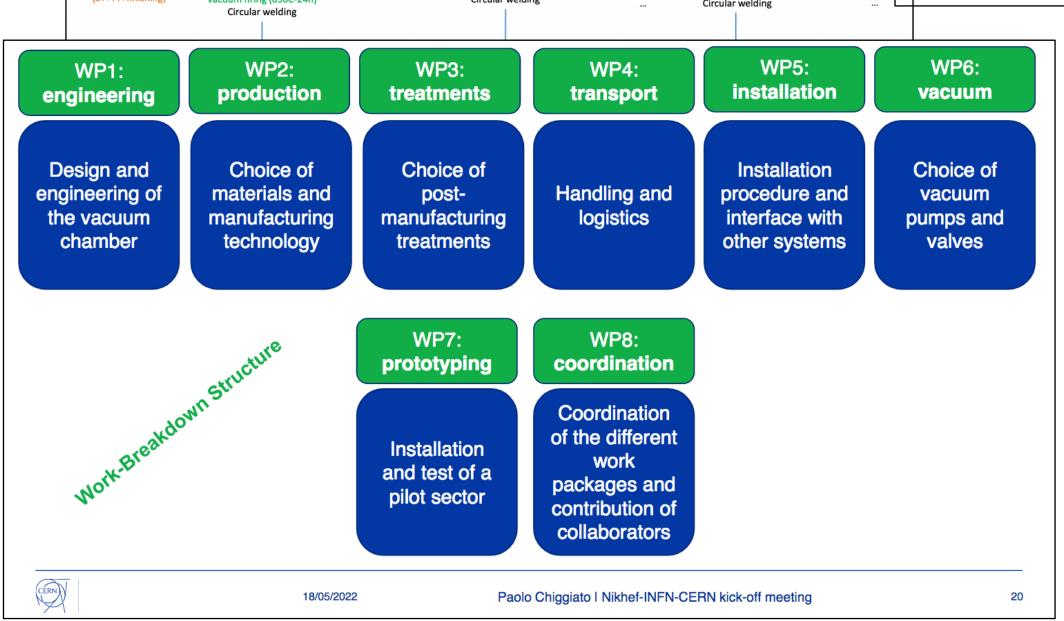












**AISI 430** 

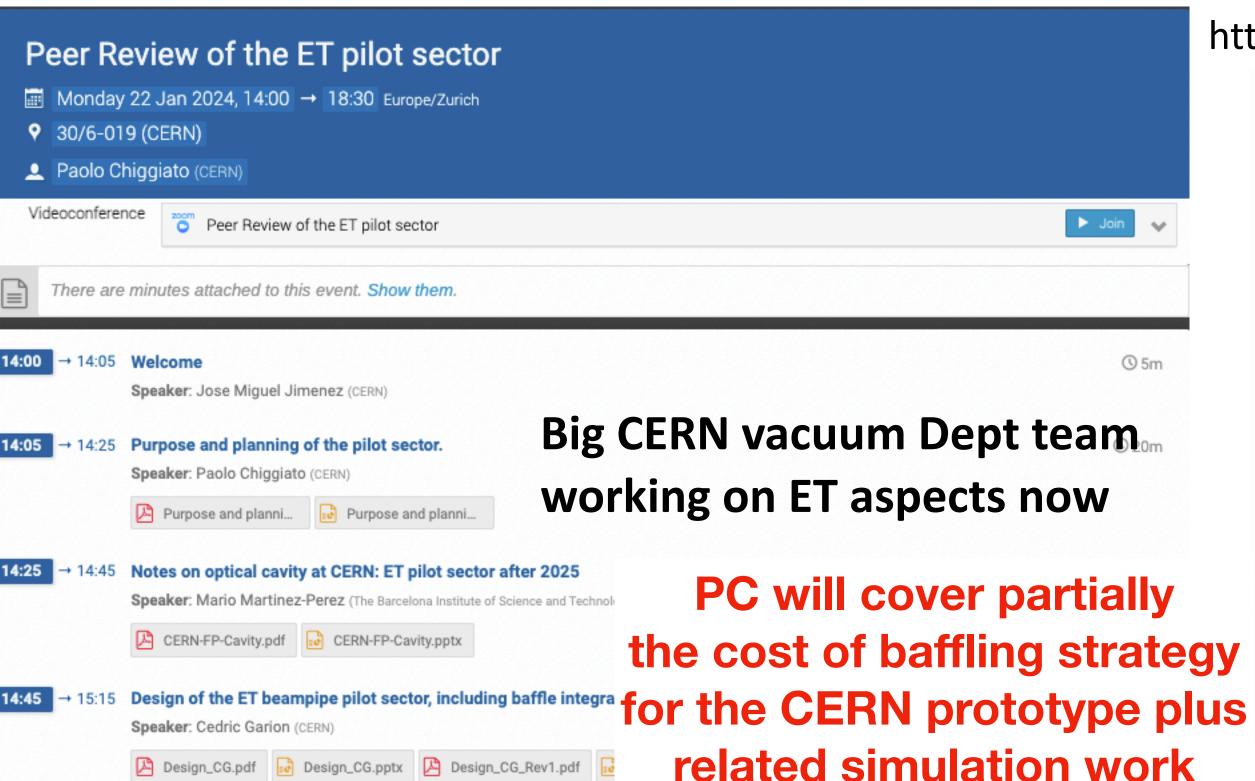
Manufacturing

S315MC

On-going effort led by CERN on the design of ET vacuum pipe (1/3 of the total ET cost) —> Will deliver a TDR in 2 years

- Physics requirements
- Vacuum / Cryo Technology
- Civil Infra-structure
- Cost Reduction/Optimization
- Prototyping

PC will cover partially the cost of baffling strategy for the CERN prototype plus related simulation work

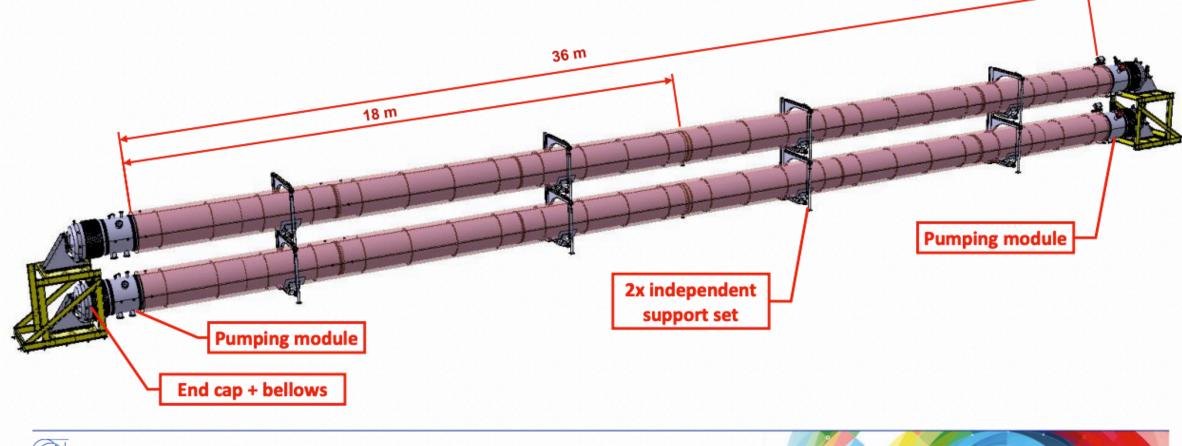


https://indico.cern.ch/event/1360696/

#### Mechanical layout of the ET pilot sector at CERN

- Two independent vacuum chambers of around 18m will be transported and installed individually.
- Each pipe is suspended on independent supports and can be aligned vertically and laterally.
- Each endcap can be aligned laterally, vertically and longitudinally.





s for witness samples to implement

#### Pilot sector: Choice of Material and

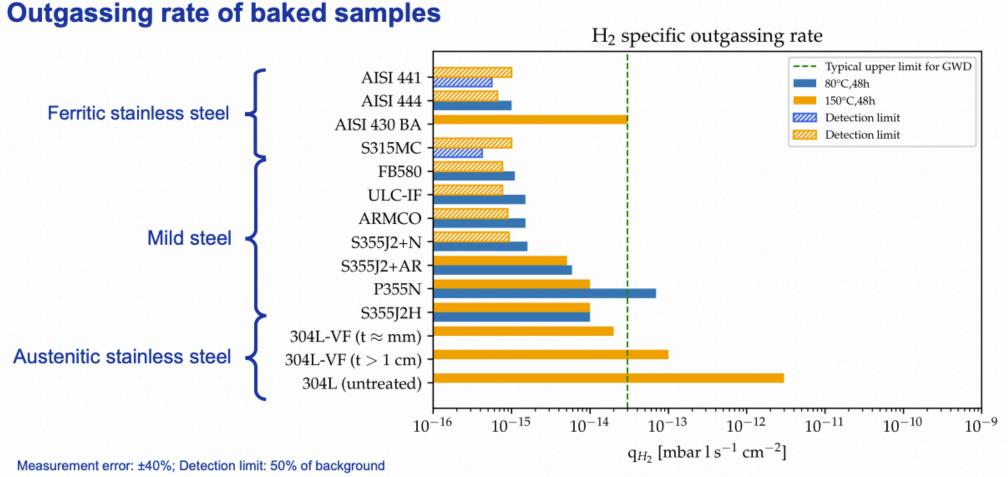
Materials	ET Vacuum requirements	Manufacturability (Welding and forming)	Corrosion resistance
AISI 304 L	Good	Good	High
S 315 MC	Good	Good	Low
AISI 441	Good	Good	High



\* Quantity of material needed for two chambers

#### Vacuum characterization of ferritic alloys

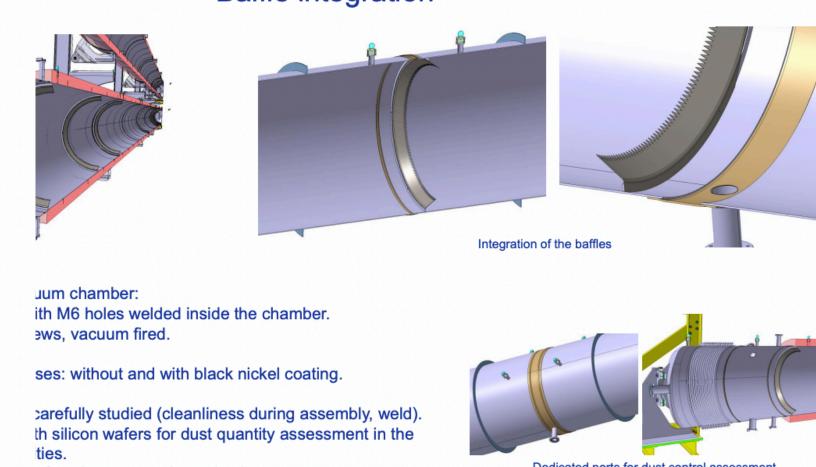
Carlo Scarcia | Review of the ET pilot sector



January 22nd, 2024

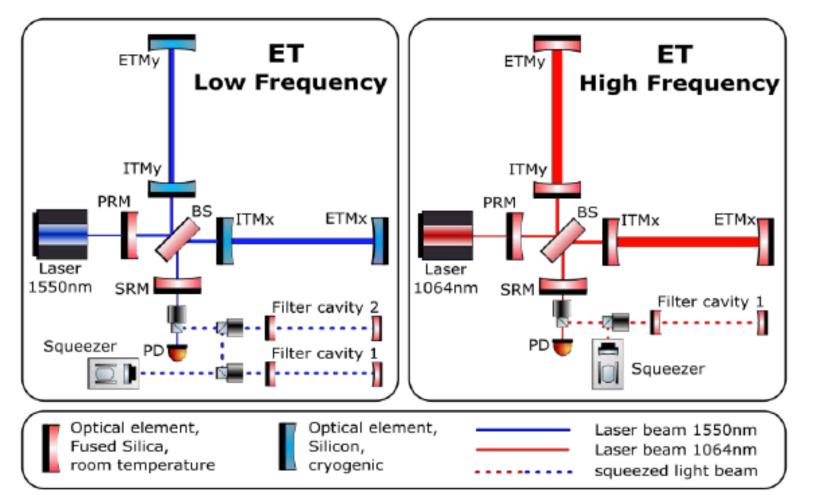
#### Baffle integration

January 22nd, 2024





# ET stray light simulations



							_	
-								
IFO	λ	mode	mirror Ø	$R_{\rm C}$	$w_0$	<i>z</i> <sub>0</sub>	w	g-factor
ET-HF	1064 nm	TEM <sub>00</sub>	62 cm	5070 m	1.42 cm	5000 m	12.0 cm	0.95
ET-LF	1550 nm	TEM <sub>00</sub>	45 cm	5580 m	2.9 cm	5000 m	9.0 cm	0.63

Running optical simulations to determine the best strategy for stray light mitigation in ET

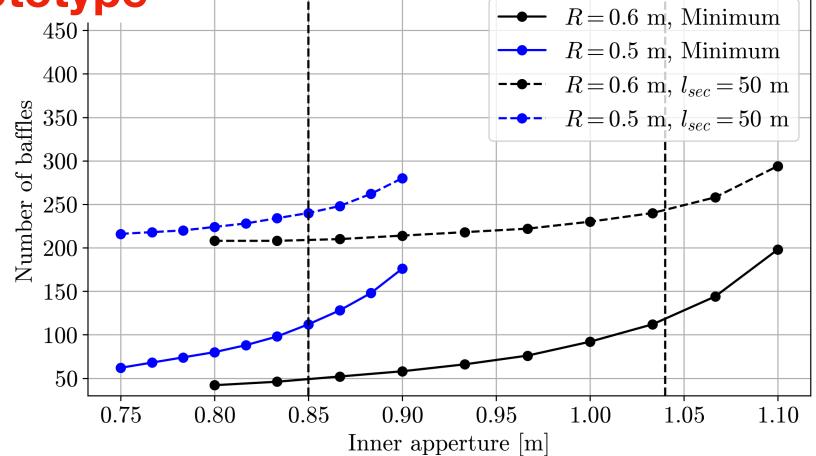
- -> Desire to extend it to CE
- -> Running common ET/CE meetings

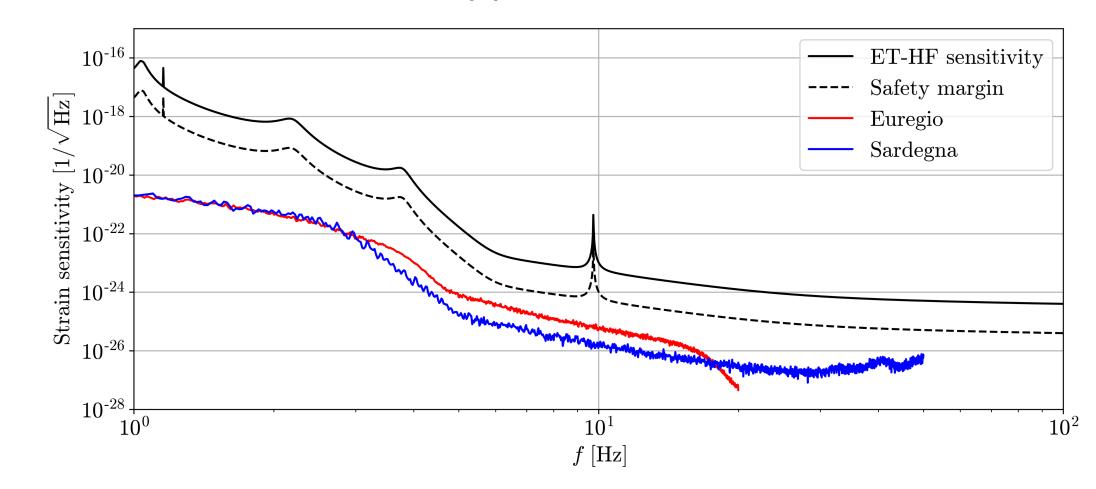
A joint effort with CERN on the very details of the vacuum pipe design and the baffling integration strategy, eventually including active monitoring inside the cavity

-> IFAE will build baffle prototypes

PC covered the cost of simulations for baffling strategy



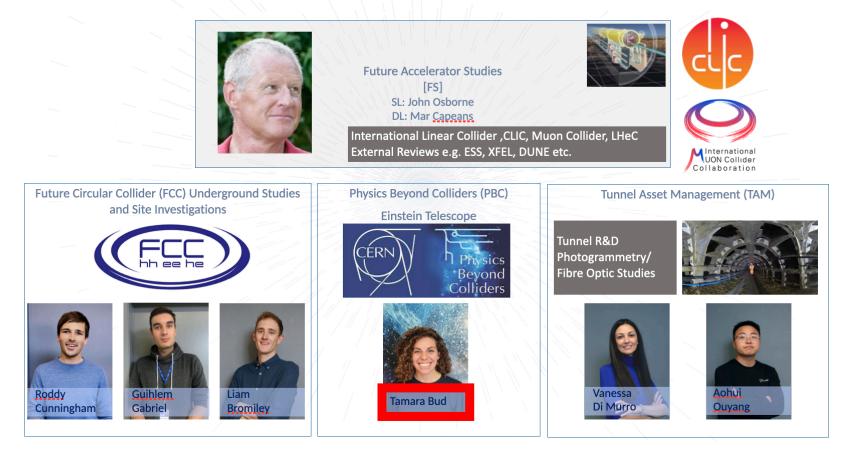




M. Andrés-Carcasona et al., Phys. Rev. D 108, 102001 (2023)

# ET Civil Eng. workshop at CERN 29-30th April 2024

#### **SCE-SAM-Future Studies Section**





- Successful workshop at CERN, hosting colleagues from ETO, INFN, Nikhef, IFAE, Local Teams, Amberg, Tractebel, Rocksoil.
- Seed questions from the local teams addressed, and hosted a lively discussion
- Identified challenges and considered next steps together
- Decided that a baseline report addressing some of the design complexities of the project would be a positive next step

## Ongoing layout discussion



#### Two scenarios

- D of 10 km
- D of 15 km



Two scenarios

- 2L of 15 km
- 2L of 20 km

2L misaligned of 45°

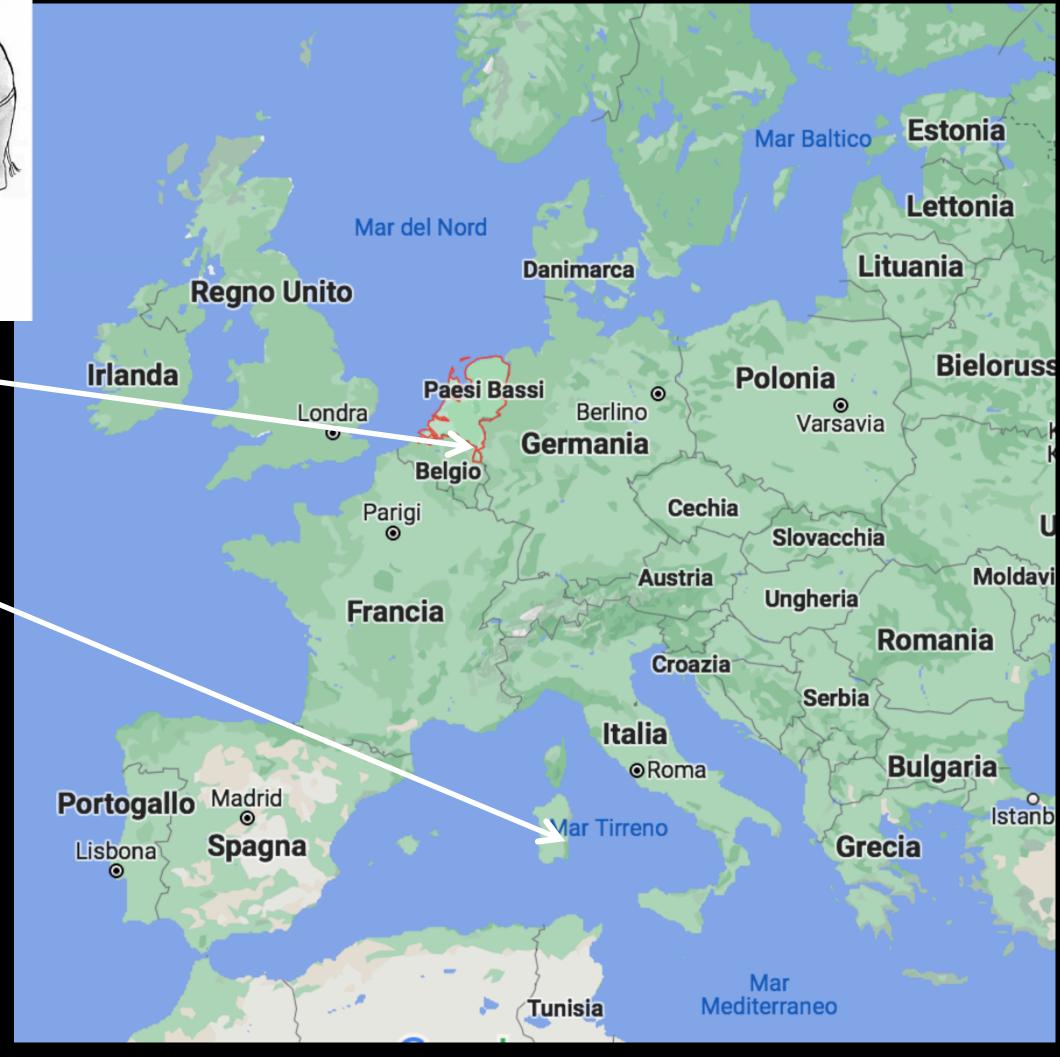
Full sensitivity: HF+LF

Only HF

Always underground

Scientific community made a study of physics potential comparing ET-baseline (triangle) vs 2L configurations



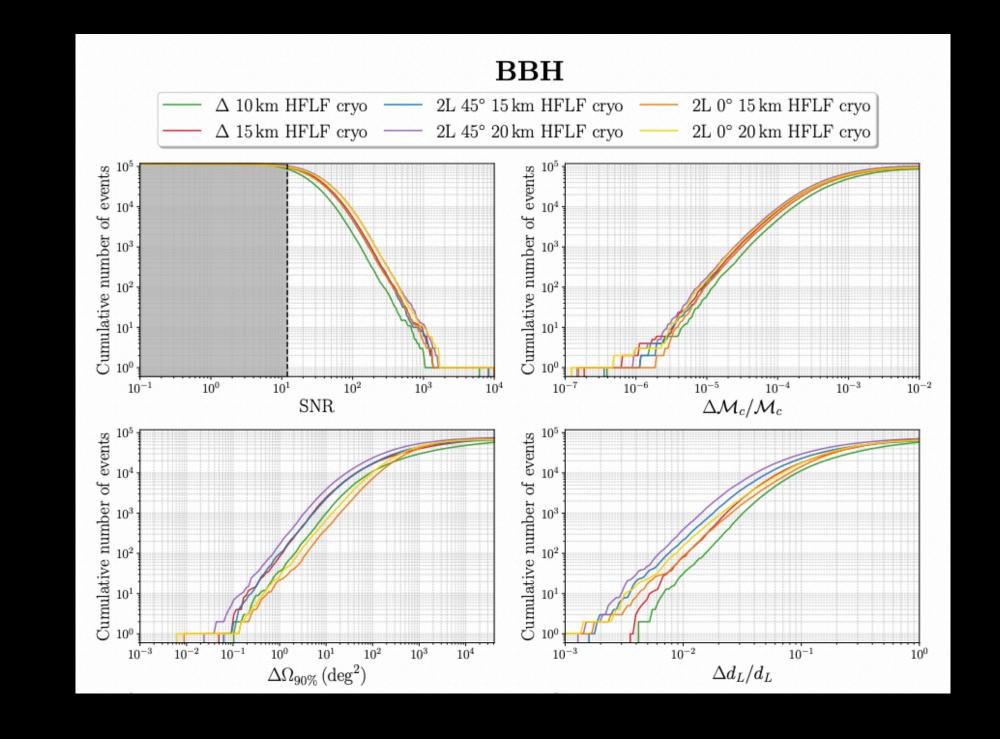


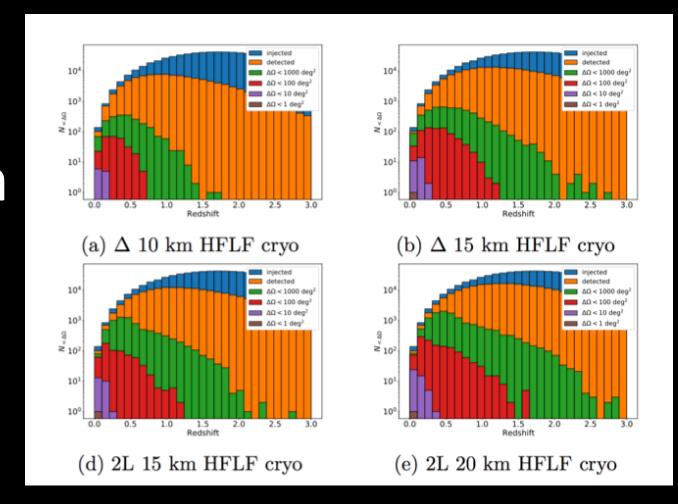
- -> ETO received the mandate to present full cost of risk analysis
- -> Will compare underground 10 km ET triangle vs underground 2L 15km

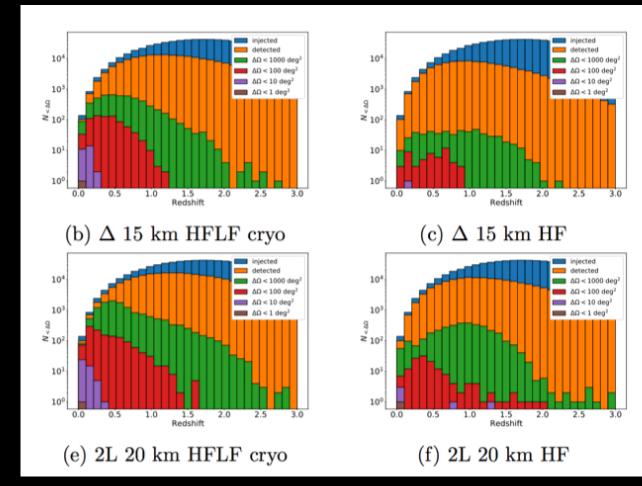
# Comparative studies (only physics potential)

- The conclusions are the expected ones
  - Longer arms perform better
  - Only 1L is not an option (dependent on overall network)
    - Either one site Triangle or 2 sites L
- NO LF translates into reduction of well localised events (more severe for triangular configuration)
  - 2L HF 15 km comparable to HFLF Δ10km
- Only LF makes BNS pre-meger alerts possible —> impact on multi messenger
- Concerns about possible correlations in the  $\Delta$  compromising stochastic searchers

arXiv:2303.15923 Branchesi, Maggiore et al. 2023





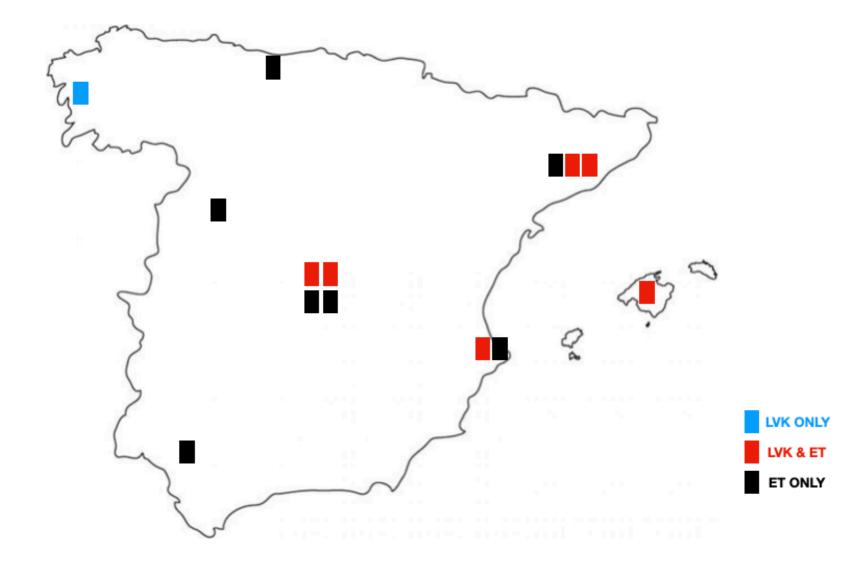


# Map of Spanish involvement in LVK/ET

Group	LVK	ET	Hardware	Computing	Note
CIEMAT	Virgo	Y	Y	YY	Still qualifying in Virgo
ICE	LISA	Y	YY-LISA	N	Not allocated FTEs for HD
IFAE	Virgo	Y	YY	YY	
IFIC		Y			TH
UIB	LIGO	Y		Υ	TH/Analysis
USC	LIGO	Ν			TH/Analysis
USAL		Y			TH
IFT	Virgo	Y			TH/Analysis
UCM		Y			TH
UCAN		Y			TH
ICCUB	Virgo	Y		Υ	TH/Analysis
UV	Virgo	Y			TH/Analysis
IEM		Y			TH/Analysis
IAA		Y			TH/Analysis

After a INFN-IFAE dedicated discussion in June 2023 of possible synergies... the message has been propagated to all the Spanish HEP community showing the R&D opportunities (including plenary talks in HEP all hands CPAN meetings)

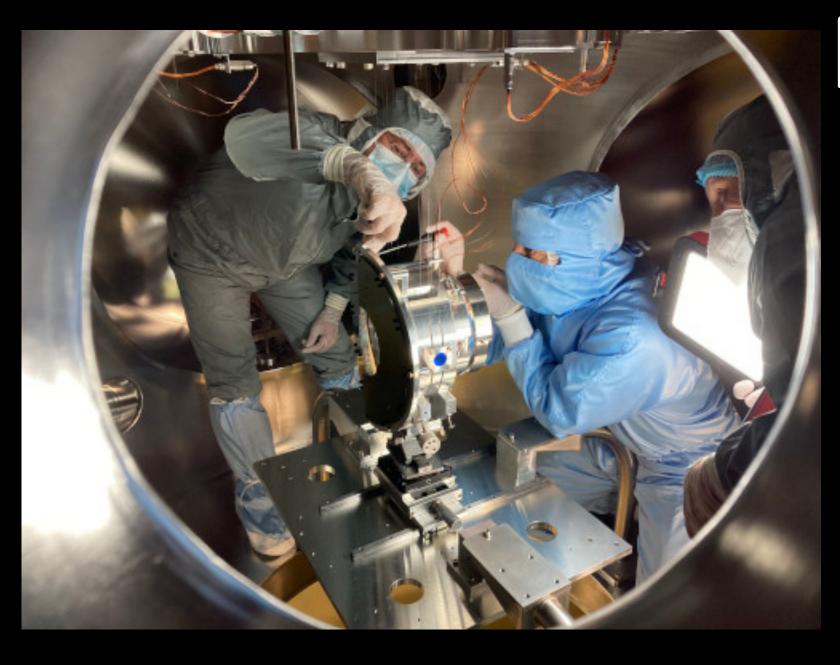
We made the effort of visiting ICFO (x2); ICMAB; ICREA... etc... Scheduled a visit to CLPU (Pulsed Lasers) in Salamanca in 2 weeks



- There are many groups interested in ET, some of them from HEP, some of them detector builders
- A majority has no hardware capacities
- Most of them are Theory and GW analysis oriented
- As today there is no consolidated large enough experimental community in GW is Spain.

We are actively exploring synergies with other research centers in Spain to improve the situation

Close contact with UPC (Applied Optics)



## Final notes

- ET is the leading EU 3G project today and is regarded as a flagship
- Together with CE will drive the GW field in 203X
- Strong proposals for hosting the ET infrastructure in place including already secured money for construction
- Intensive R&D activity in all fronts (somehow focused on LF)
- In the following years the ET project will progress towards a site selection(s), final design and governance, costs and risk studies, aiming for starting construction before the end of the decade [?]
- ET-PP EU Preparatory Phase Project is a tool to build coherence in the process (never easy...) and to glue "competing teams"
- The re-discussion of ET geometry saga might slow down the progress as it is putting ET in a "non-projected quantum state of geometry and location(s)"
- ET-Spain needs to grow and get experimental community more involved.

