

RADES project

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Axion dark matter

• Axions from astronomical sources (Baby IAXO from Sun)

• Axions produced on lab (ALPS II)



• Axions in the galactic halo as dark matter





Axion dark matter

Axion's parameter space



Haloscope concept



ADMX



Power:

$$P_a = g_{a\gamma}^2 \rho_{\rm DM} \frac{\beta}{1+\beta} \frac{1}{m_a} B^2 V Q_L G^2,$$

Haloscope concept



Scanning rate:

$$dt = \Delta v_a \left(\frac{\frac{S}{N}k_B T_{sys}}{P_d}\right)^2 = \frac{m_a}{Q_a} \left(\frac{\frac{S}{N}k_B T_{sys}}{P_d}\right)^2$$

Is used as figure of merit:

$$\frac{dm_a}{dt} = \frac{Q_a}{Q_l} \left(\frac{P_d}{\frac{S}{N}k_B T_{sys}}\right)^2 = Q_a Q_l \kappa^2 g_{a\gamma}^4 \left(\frac{\rho_a}{m_a}\right)^2 B_e^4 C^2 V^2 \left(\frac{S}{N}k_B T_{sys}\right)^{-2}$$

Rades setups





Future experiments



BabyIAXO - Rades



Scaled version 10 times smaller



Future experiments

MPP Munch - Rades



Vertical cut tuning Frequency range 8-9 GHz









Counting photons







Dark count rate for SPD: 1-100 mHz

 10^{3}

 $\lambda_{\mathrm{thresh}}$

10⁶

 10°







Dilution refrigerator











Dark photon sensitivity plot

 Dixit et al. Searching for dark matter with a superconducting qubit, Phys. Rev. Lett. 126





Thank you very much!!





Background slides

Two tone spectroscopy





Rabi chevrons



Single shot



Characteristic times (Relaxation)



Pulse sequence

Characteristic times (decoherence)



*P.Krantz, et.al., Appl. Phys. Rev. 6, 021318 (2019)







Ramsey chevrons (thermal photons in the cavity)





Detection protocol



ERC sensitivity bands



200-500 MHz

8-18 GHz