

Rare events searches and gaseous detector developments at CAPA

Álvaro Ezquerro Sastre
aezquerro@unizar.es



Universidad
Zaragoza

1542

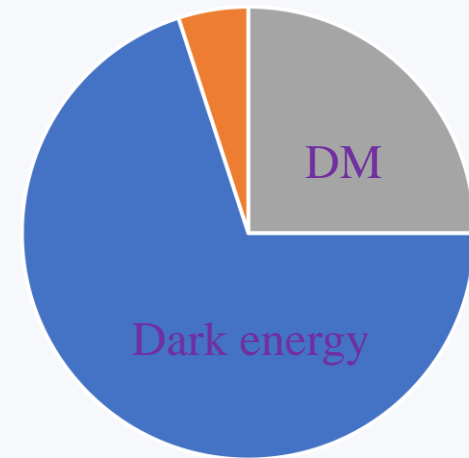


**Centro de Astropartículas y
Física de Altas Energías**
Universidad Zaragoza



CAPA activites

- CAPA: Centre for Astroparticle and High Energy Physics
- Research lines:
 - **Low background techniques and detector development**
 - Dark matter, axion physics and neutrino physics
 - Lattice gauge theory and field theory applications
 - Standard Model extensions and quantum gravity
 - Observational astrophysics and cosmology



Universe pie

CAPA activities

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- Research lines:
 - **Low background techniques and gaseous detector development**

non-gaseous (ANAIS)
in Ivan's talk

Low-mass WIMPs

Solar Axions

Diagram illustrating a WIMP interacting with an atom. The WIMP is shown as a black sphere with a white dot, moving towards the atom. The atom is depicted with a central nucleus and orbiting electrons.

Recoil spectra $m=1$ GeV

Recoil rate (keV/kg/day)

Energy (keV)

Legend: -Ar, -Ne, -Xe

$\sigma_{\text{WIMP-n}}^0 = 10^{-40} \text{ cm}^2$

Standard halo, SI

TREX-DM

Primakoff

γ

a

γ_{virtual}

$\times e, Ze$

CAST

Axion flux [$10^{-10} \text{ cm}^{-2} \text{ s}^{-1} \text{ keV}^{-1}$]

Axion energy [keV]

$\langle E_a \rangle = 4.2 \text{ keV}$

1 X-Ray cluster Low energy

$\times \left(\frac{g_{a\gamma}}{10^{-10} \text{ GeV}^{-1}} \right)^2$

CAST

Inclination System

Support Frame

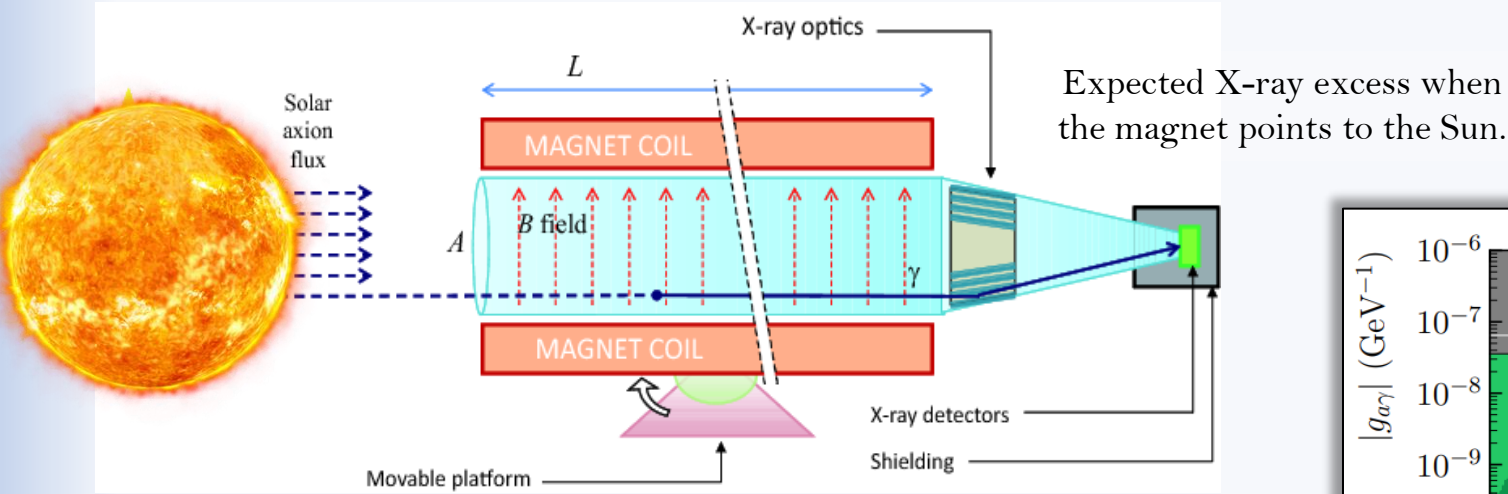
Flexible Lines

Rotating Disk

Services

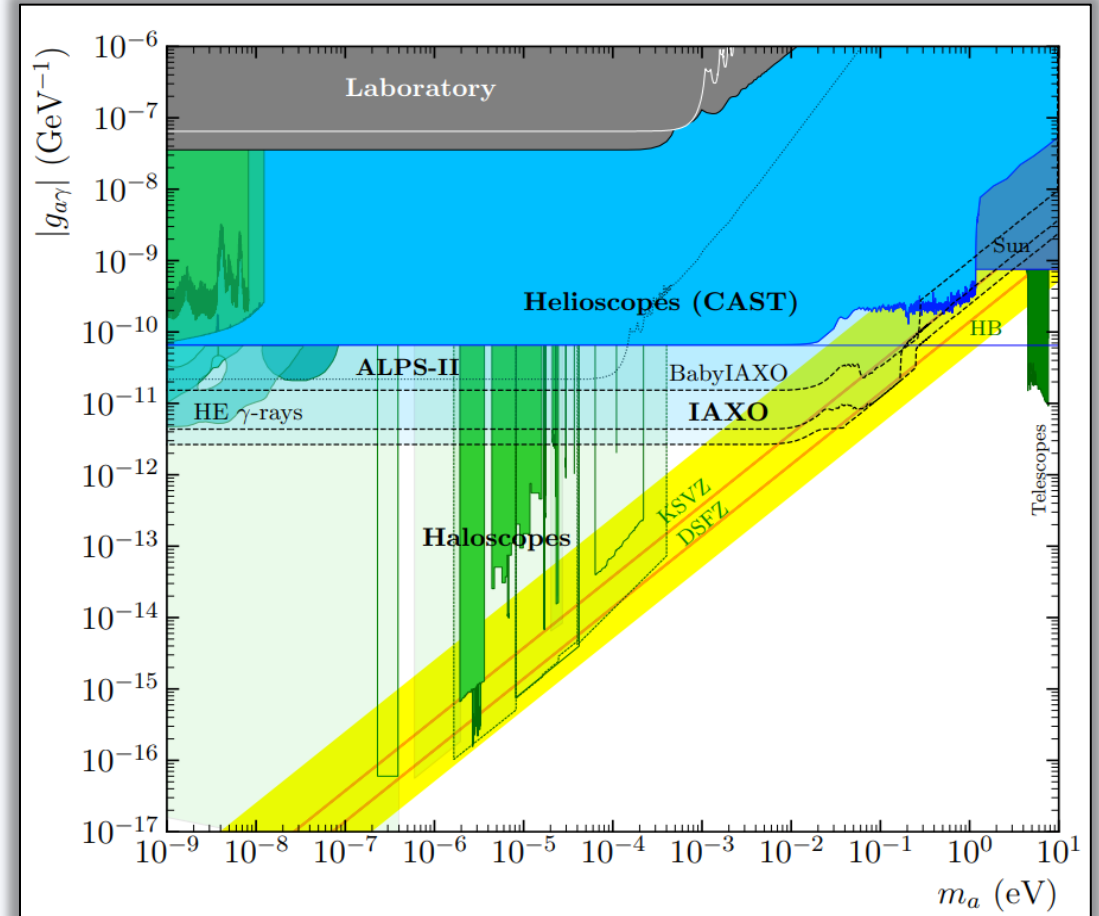
Rotation System

IAXO helioscope



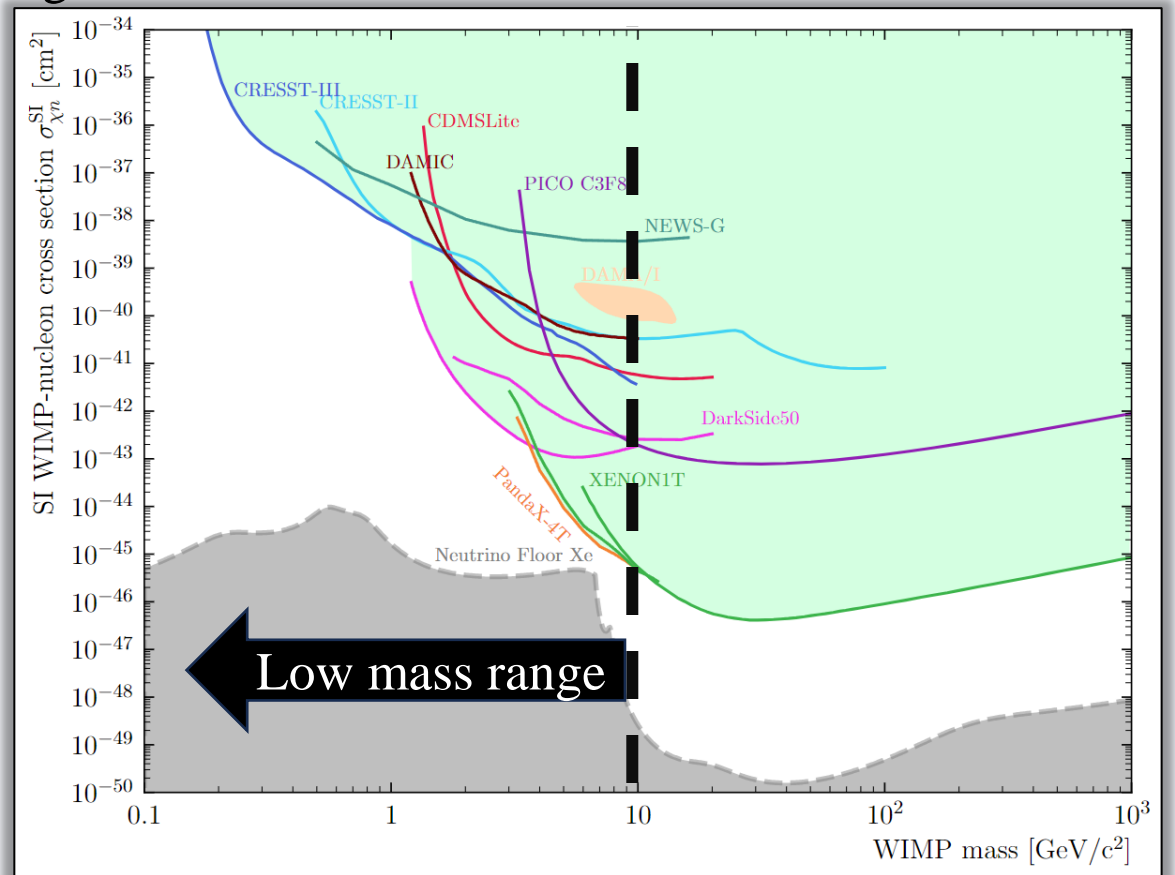
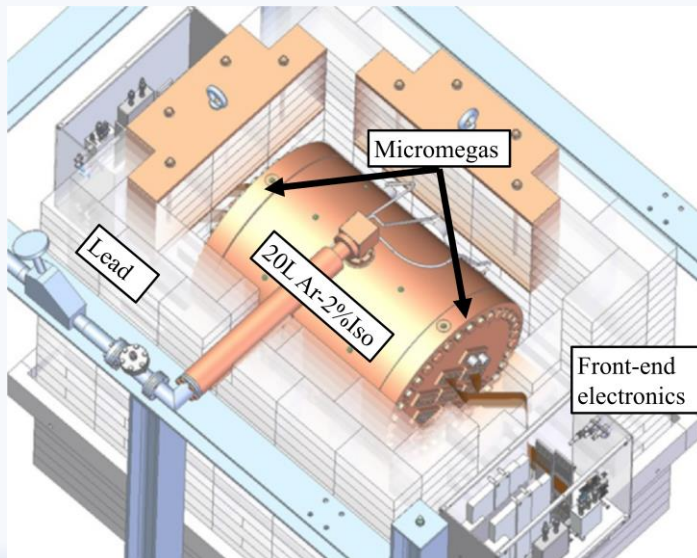
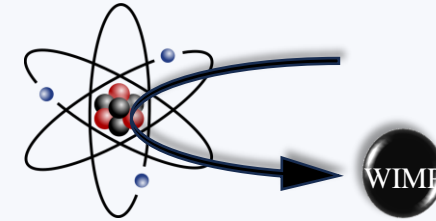
Sensitivity parameters:

- Magnet
 - Magnetic field
 - Length
- Optics
 - Efficiency
 - Area of the focused spot
- Detector
 - Efficiency
 - Background level
- Time tracking the sun



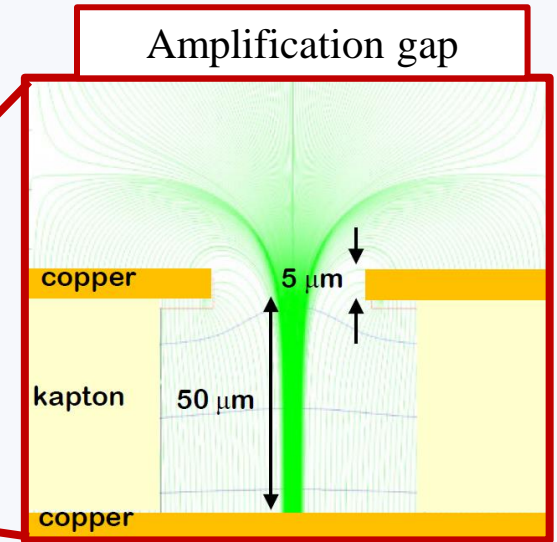
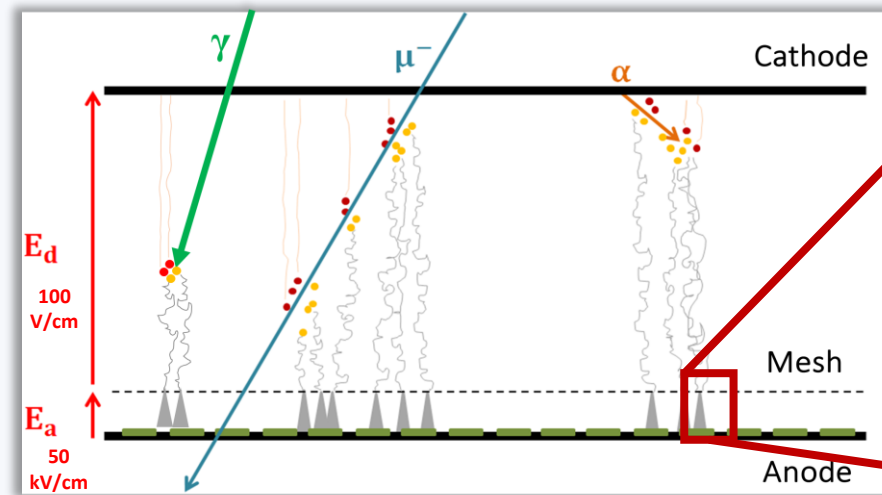
TREX-DM experiment

- Gaseous TPC of cylindrical copper vessel
 - Active volumen of 20L up to 10bar
 - Copper, lead, polyethylene ceiling&water shielding
- Located at LSC (2400 m.w.e.)
- Goal:
 - Low energy threshold < 1 keV
 - Low background level ~ 1 (keV kg day) $^{-1}$.

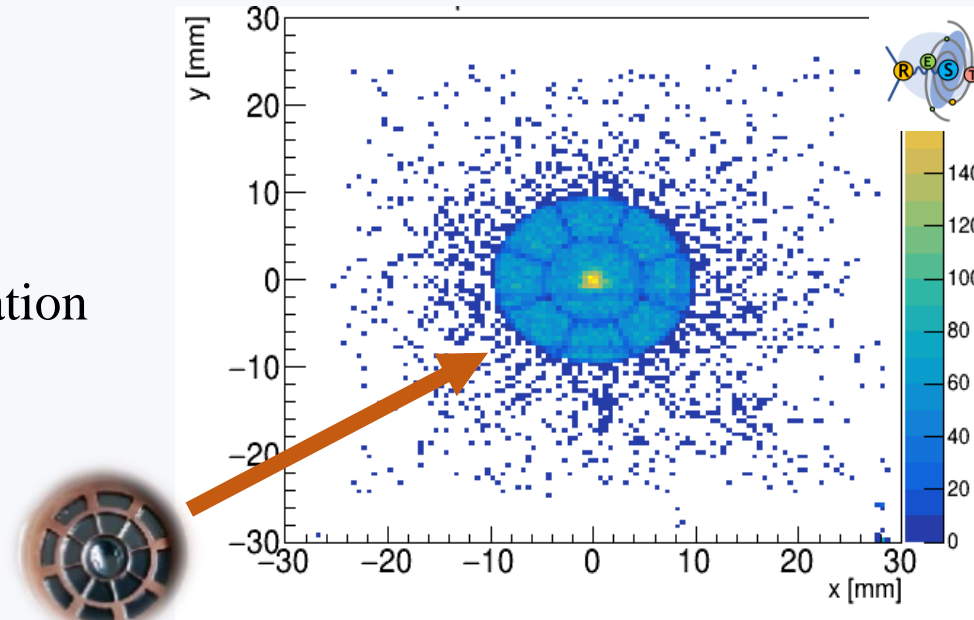


Micromegas detector

- Readout plane for gaseous Time Projection Chamber
- Low intrinsic radioactivity
- Absolute gain 10^3 - 10^4



- Topological information



IAXO-D Micromegas prototypes for

IAXO detector prototype

- Ar or Xe + quencher, 1.4 bar, 3cm drift
- Shielding (active & passive)

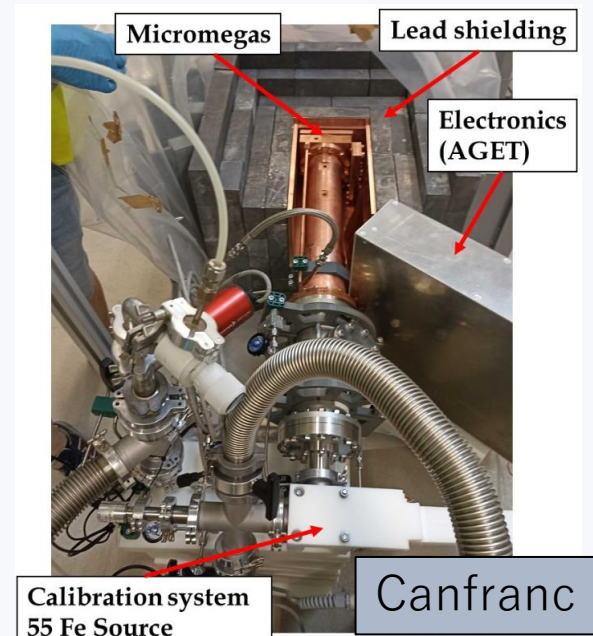
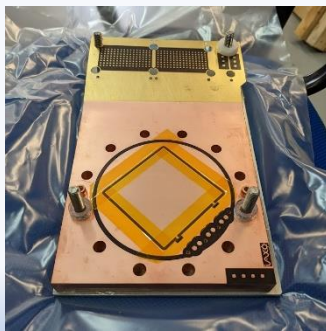
Goals

- Background level: 10^{-7} - 10^{-8} counts $\text{keV}^{-1} \text{cm}^{-2} \text{s}^{-1}$
- Energy threshold: ~ 0.1 keV

Micromegas detector

- Same design as CAST XRT-MM detector
- AGET-based electronics (radiopure version planned)

6cm x 6cm
2x 124 strips
0.475mm width,
0.5mm pitch



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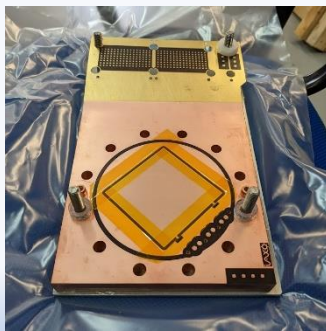
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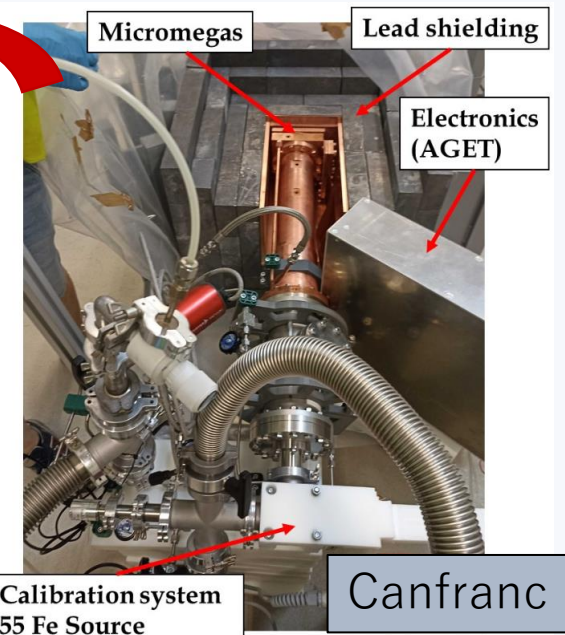
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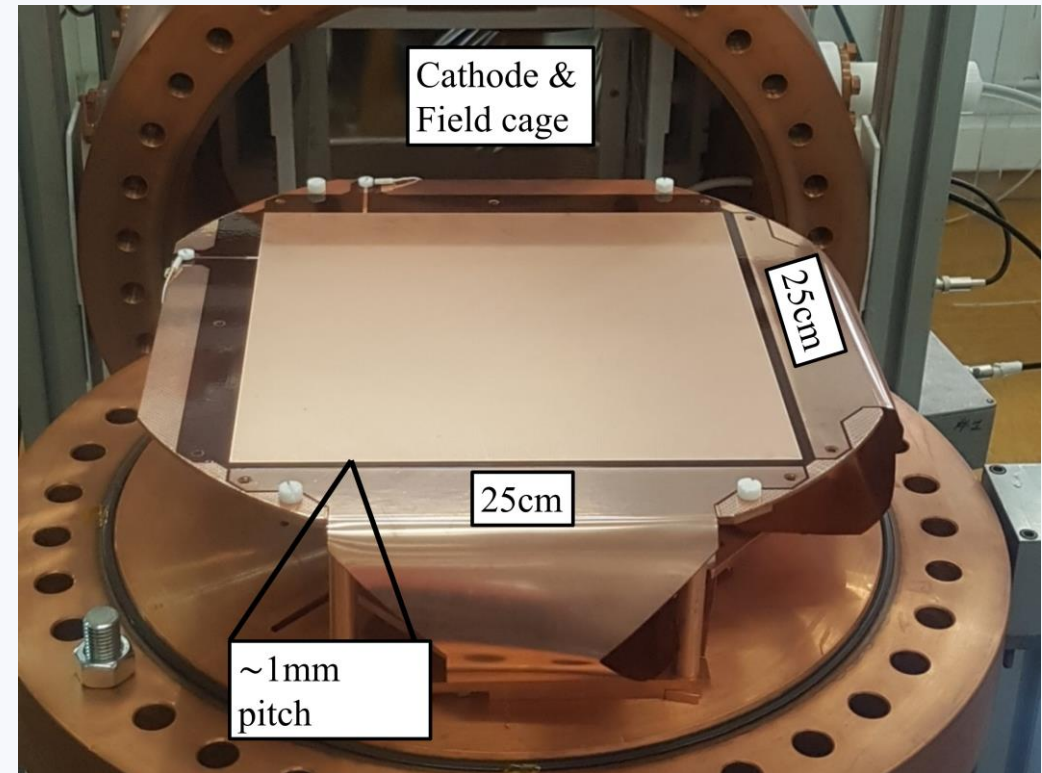
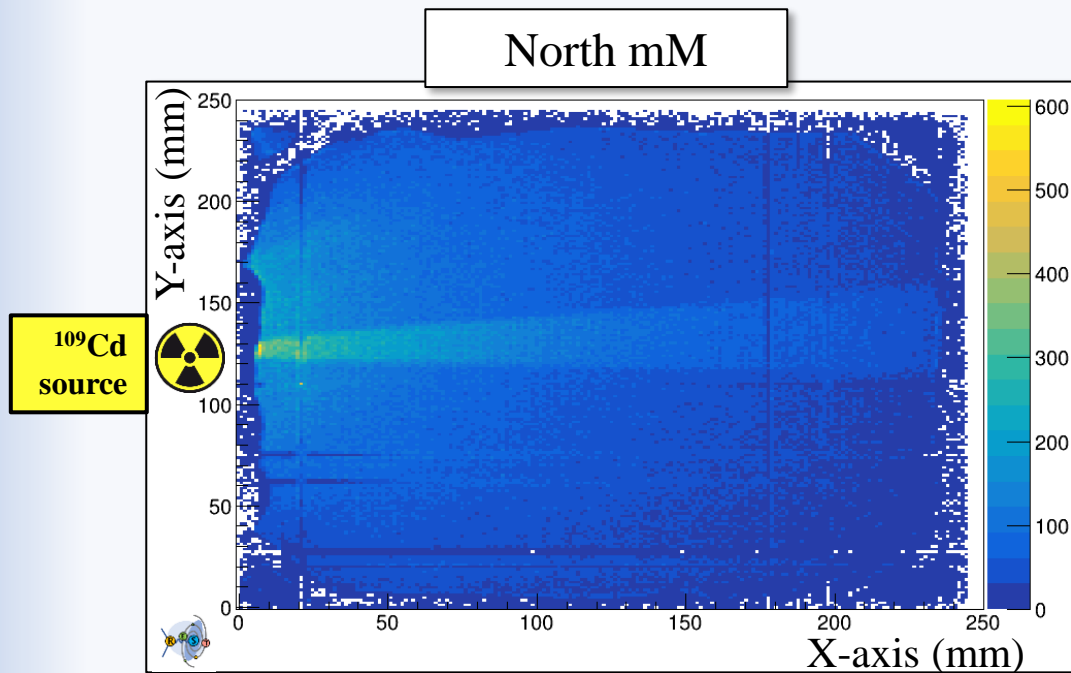


Achieved !!
(preliminary)



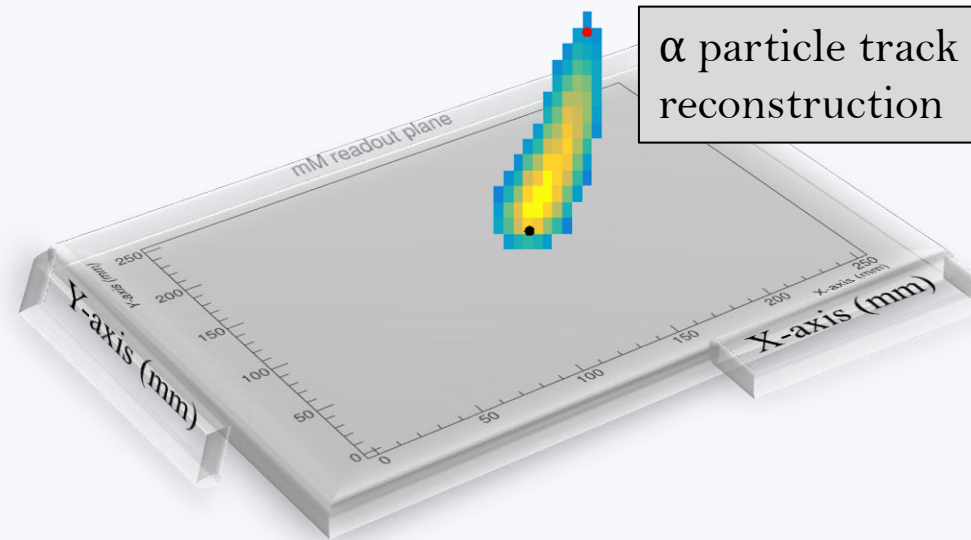
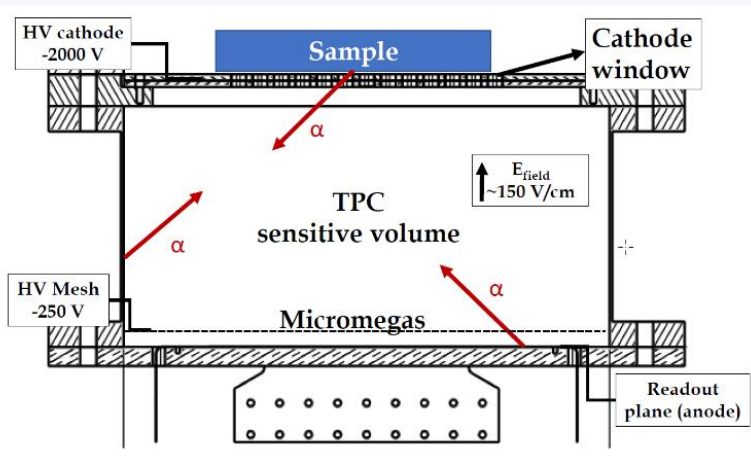
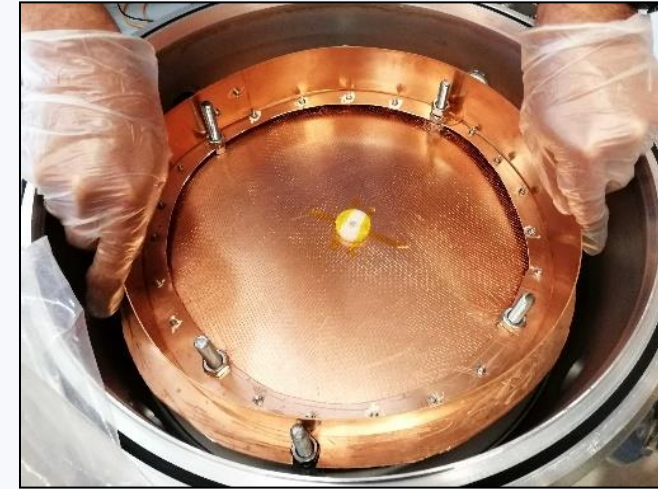
TREX-DM Micromegas

- 1 cathode and 2 readout planes
- Largest surface produced (*micro-bulk*) Micro-Mesh Gaseous Structure (**Micromegas**)
 - 512 channels: 256 X strips, 256 Y strips



AlphaCAMM (Alpha CAMera Micromegas)

- Gaseous TPC with a segmented mM (25cm x 25cm)
- Measure ^{210}Pb surface contamination of flat samples down to 100 nBq/cm^2
- Track reconstruction (origin and end) allows to identify the alphas coming from the sample

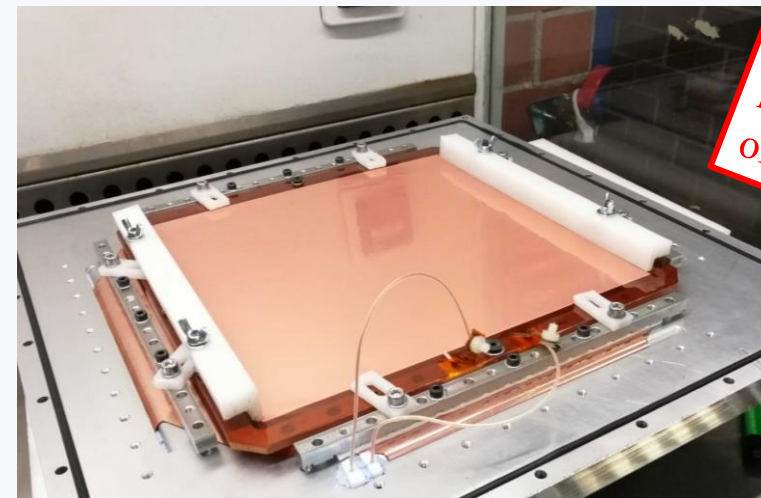
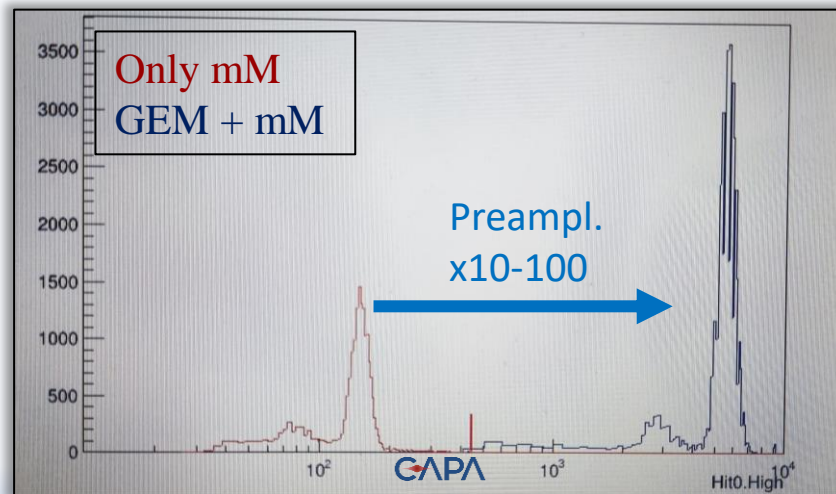
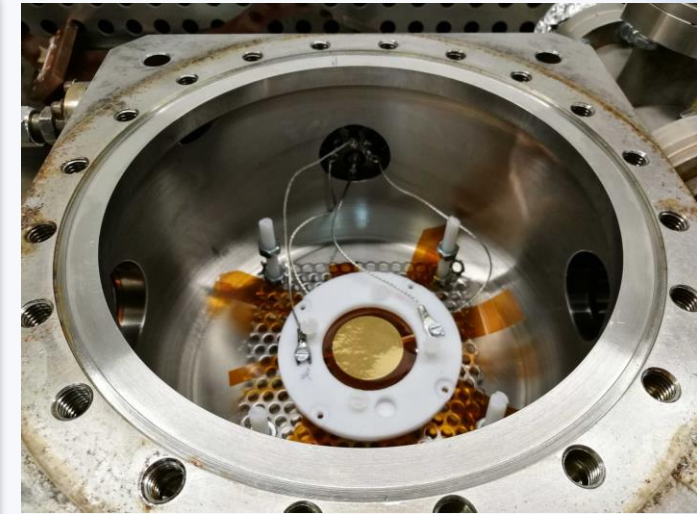
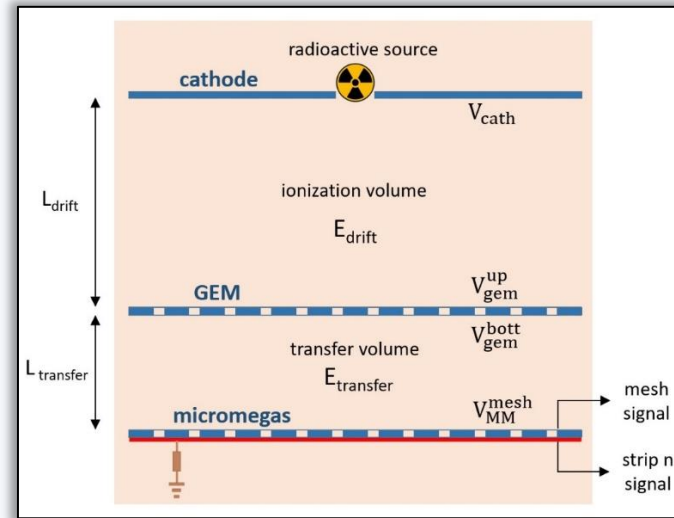


K. Altenmuller et al, 2022 JINST 17 P08035

Micromegas + GEM

- Improve gain by adding a Gas Electron Multiplier (**GEM**) amplification stage on top of the Micromegas planes
 - increase signal-to-noise ratio
- Preliminary results in test set-up shows x10-100 extra amplification factor

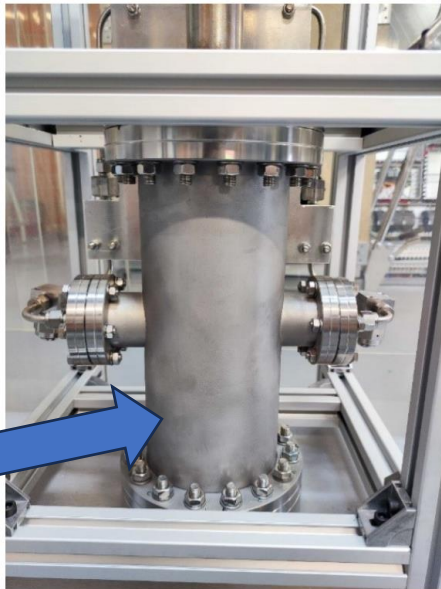
→ E_{thr} lowered down to single electron (20 eV_{ee})



GEM already installed at Left side of TREX-DM !!!

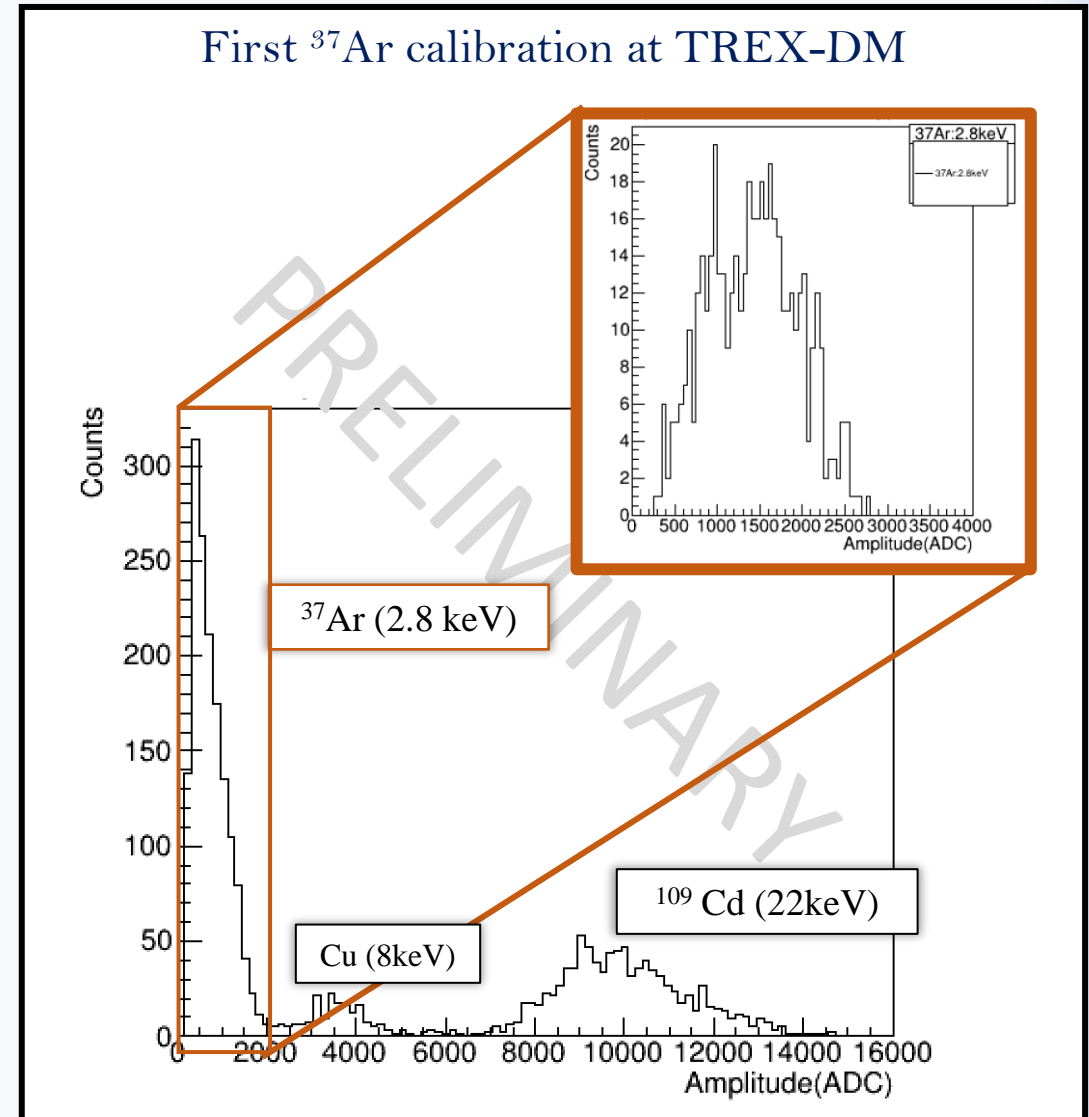
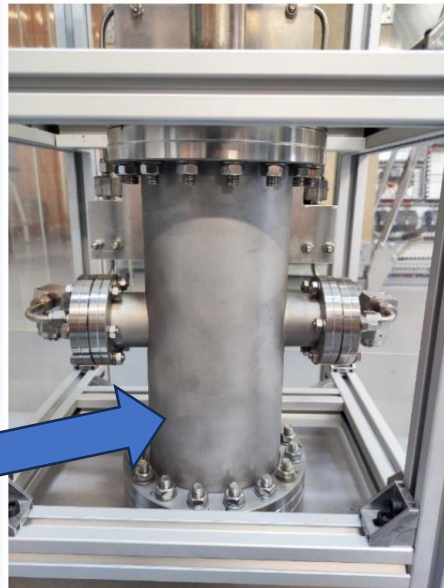
Low energy calibration (^{37}Ar)

- ^{37}Ar : 2.8 keV (90%), 0.27keV (9%)
- Gas source → homogeneous illumination
- Irradiated CaO powder at CNA (Sevilla)
 - 6 hours
 - ~ 1 kBq



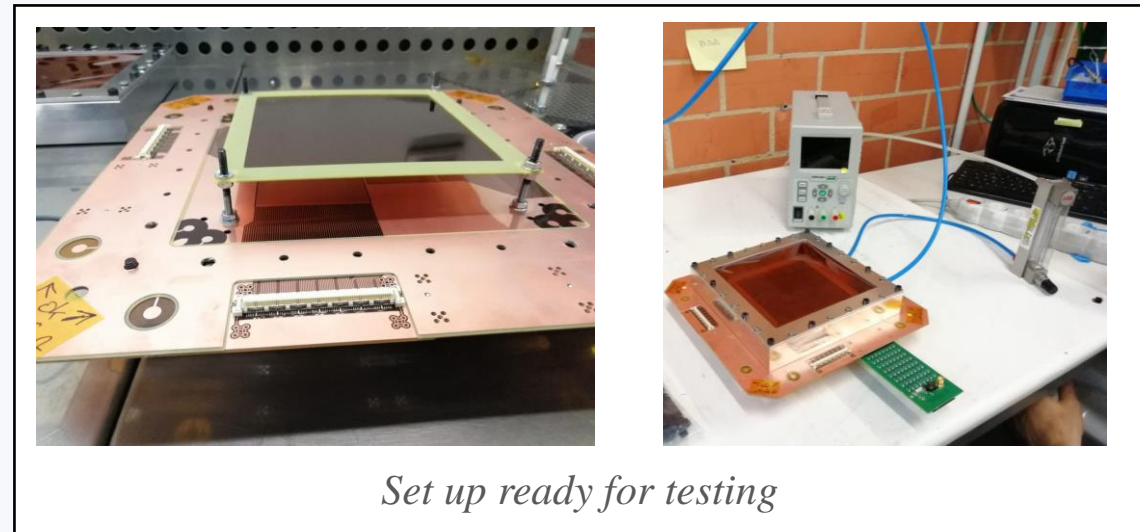
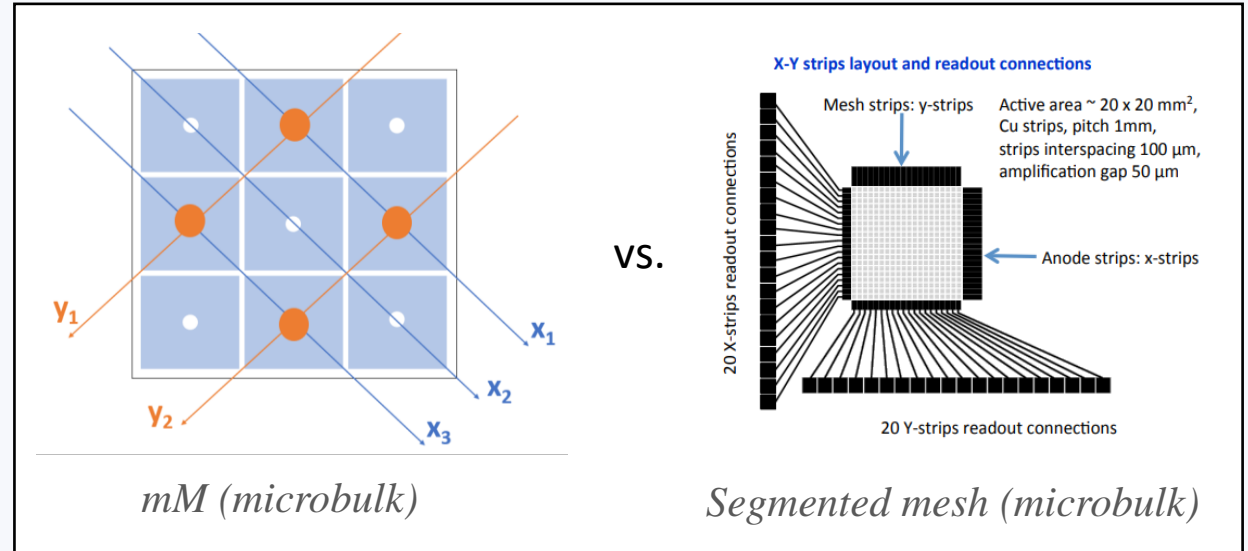
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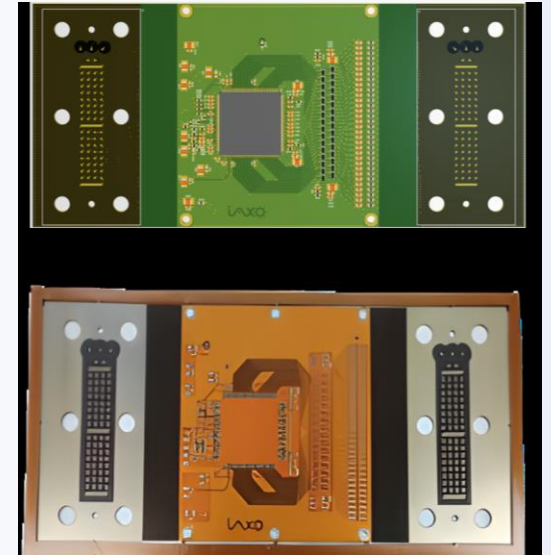
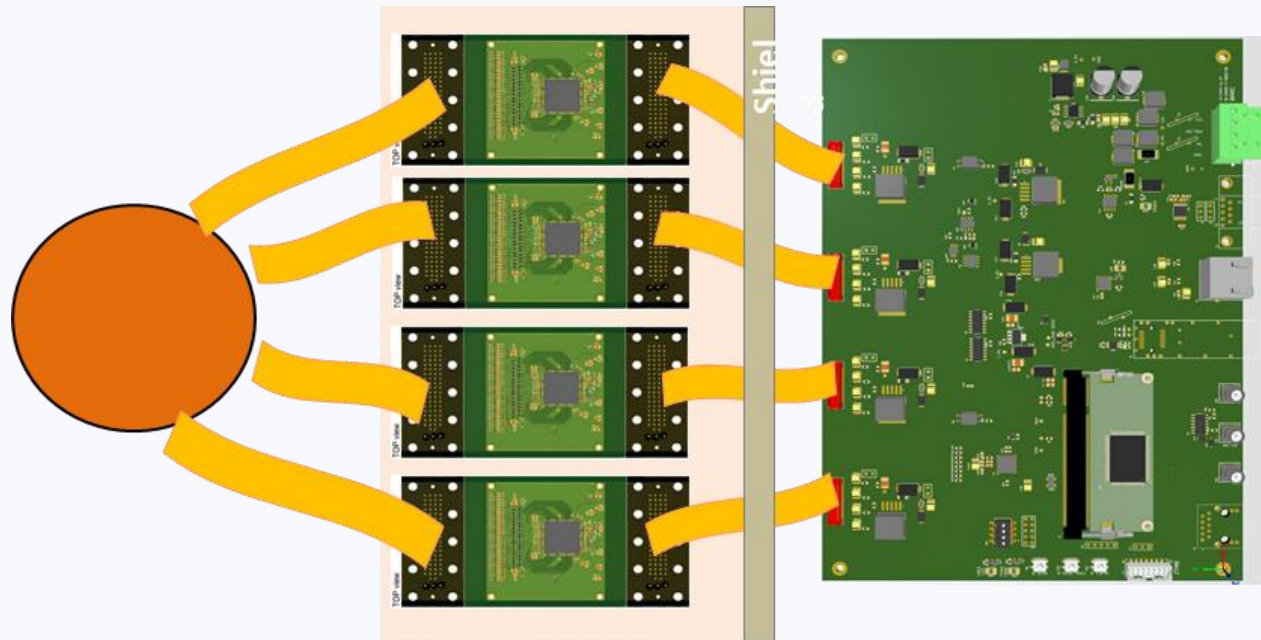
Segmented mesh

- ✓ Simpler construction process
- ✓ Less material budget
- ✓ Better spatial resolution
- ❖ Dedicated electronics to power the channels of the mesh and read the signal



Radiopure electronics

- Usually readout electronics outside the shielding
 - Longer cables, higher noise/threshold
- Development of radiopure FE boards
- Complements with non-radiopure BE



✓ Limandes
mM-FEC

✓ FEC PCB
❖ FEC components

✓ Limandes
FEC-BEC ❖ BEC

CAPA

ICCUB

Institut de Ciències del Cosmos
UNIVERSITAT DE BARCELONA

EXCELENCIA
MARIA
DE MAEZTU
2020-2023

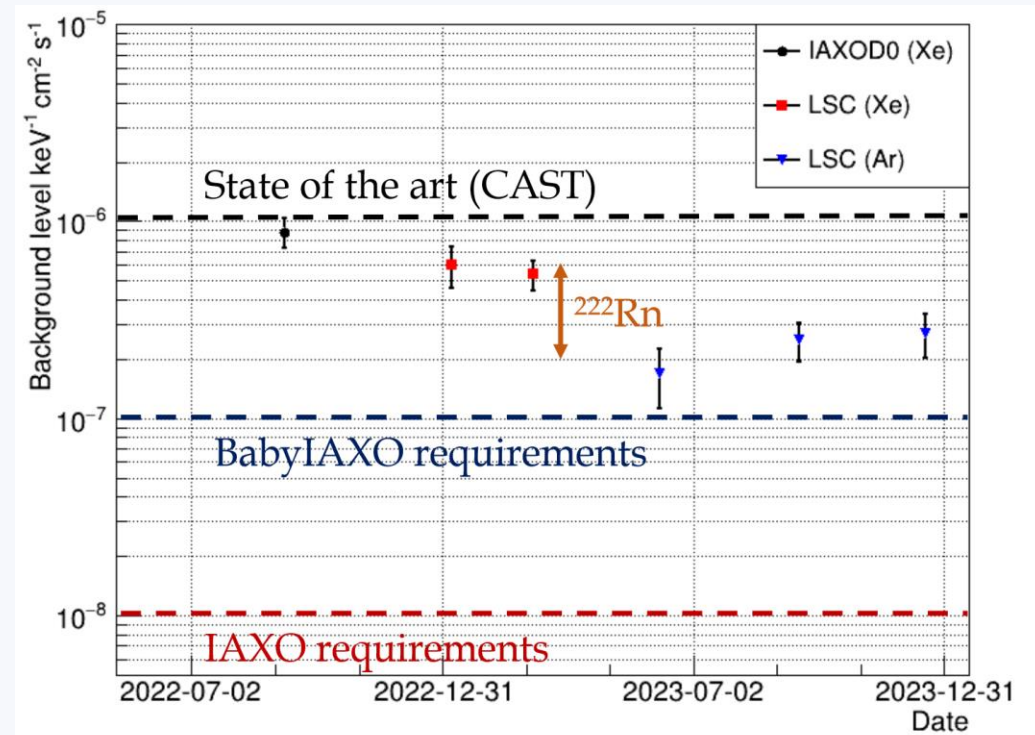
Goals & Milestones

TREX-DM

- GEM installed (one side)
- First ^{37}Ar calibration
- Cathode changed
- AlphaCMM (non-radiopure) commissioning ongoing

IAXO-D

- Reduced background level at IAXO-D1 (LSC)



- Radiopure electronics soon

THANKS!

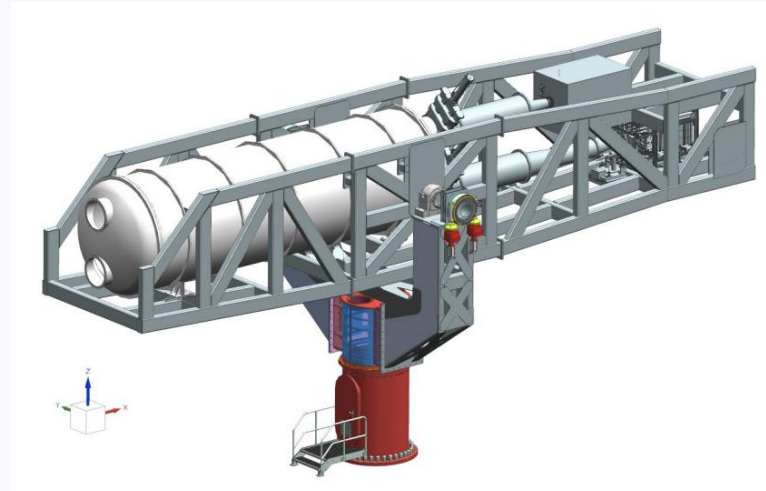


BACK-UP



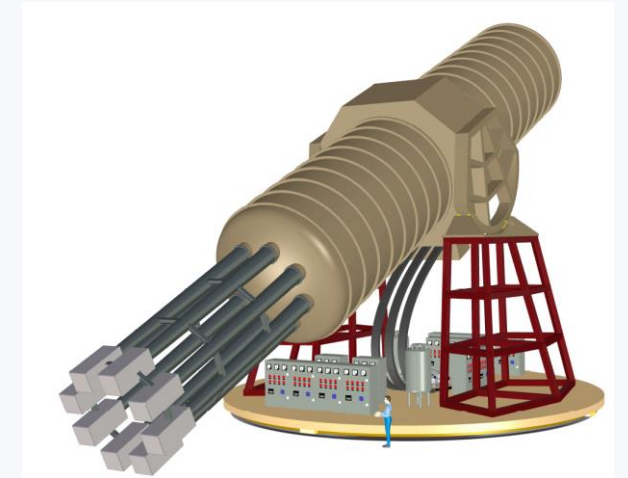
CAST

Tracking Time	2 x 1.5h / 24h
Magnetic Field	9 T
Length (B)	9.3 m
# Magnet Bores	1*
$f_M [T^2m^4]$	21



BabyIAXO

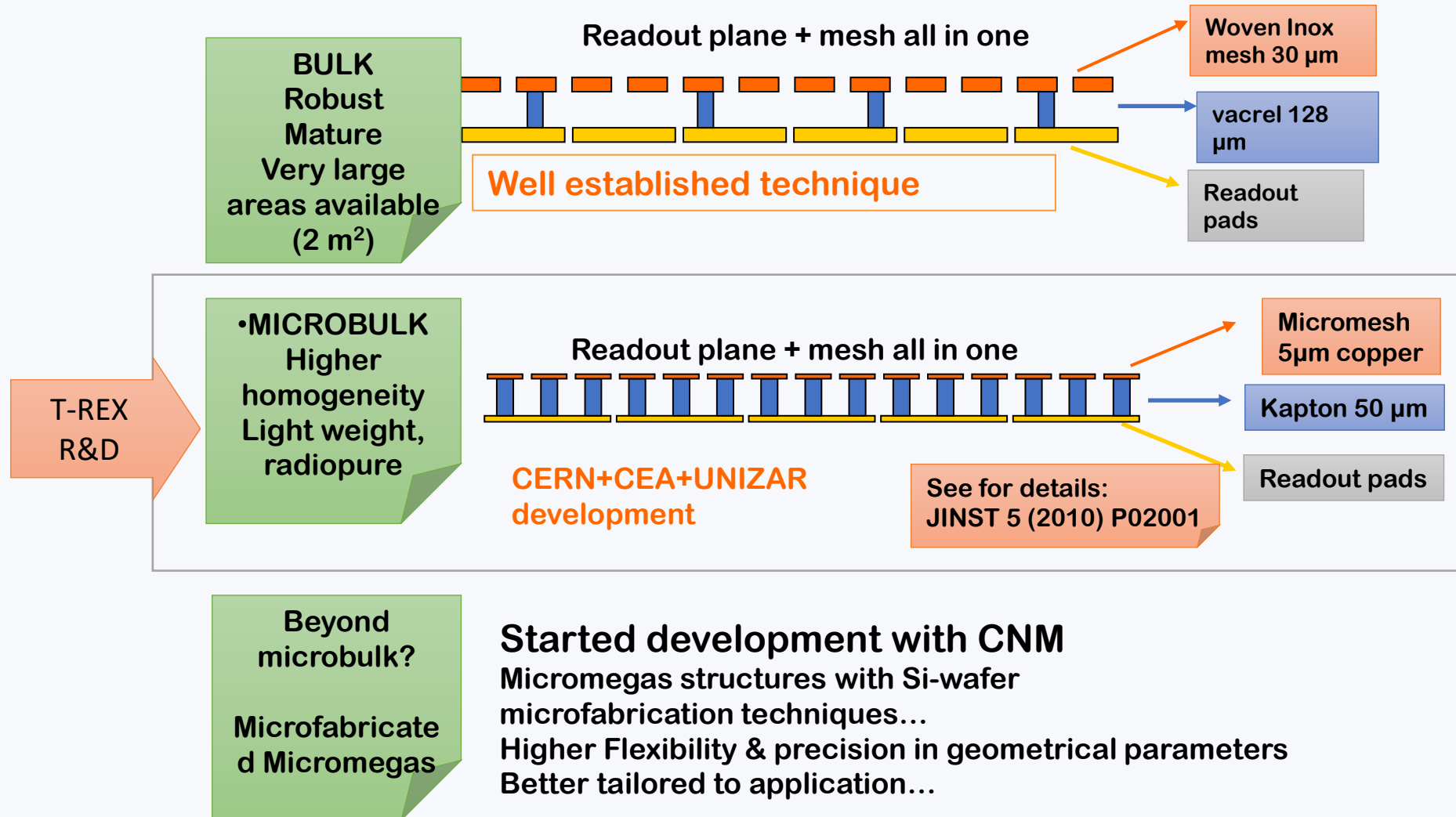
Tracking Time	12h / 24h
Magnetic Field	~ 2 T
Length (B)	9.3 m
# Magnet Bores	2
$f_M [T^2m^4]$	~ 230



IAXO

Tracking Time	12h / 24h
Magnetic Field	~ 2.5 T
Length (B)	21 m
# Magnet Bores	8
$f_M [T^2m^4]$	~ 6000

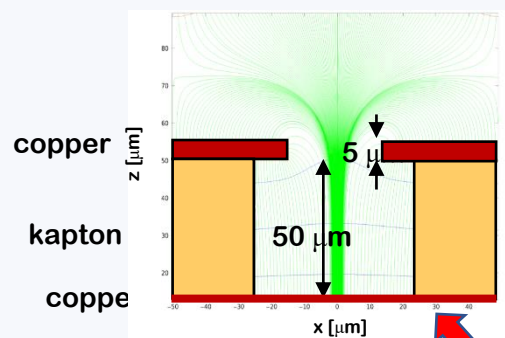
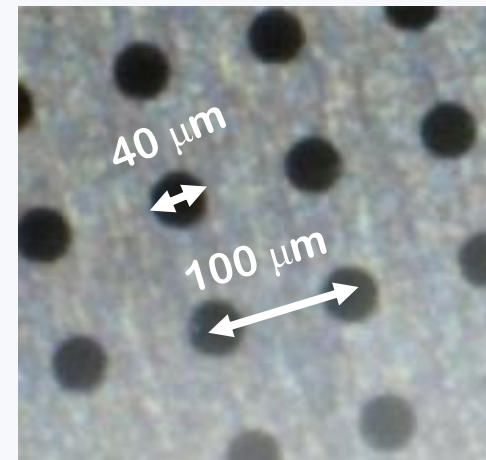
Micromegas readouts



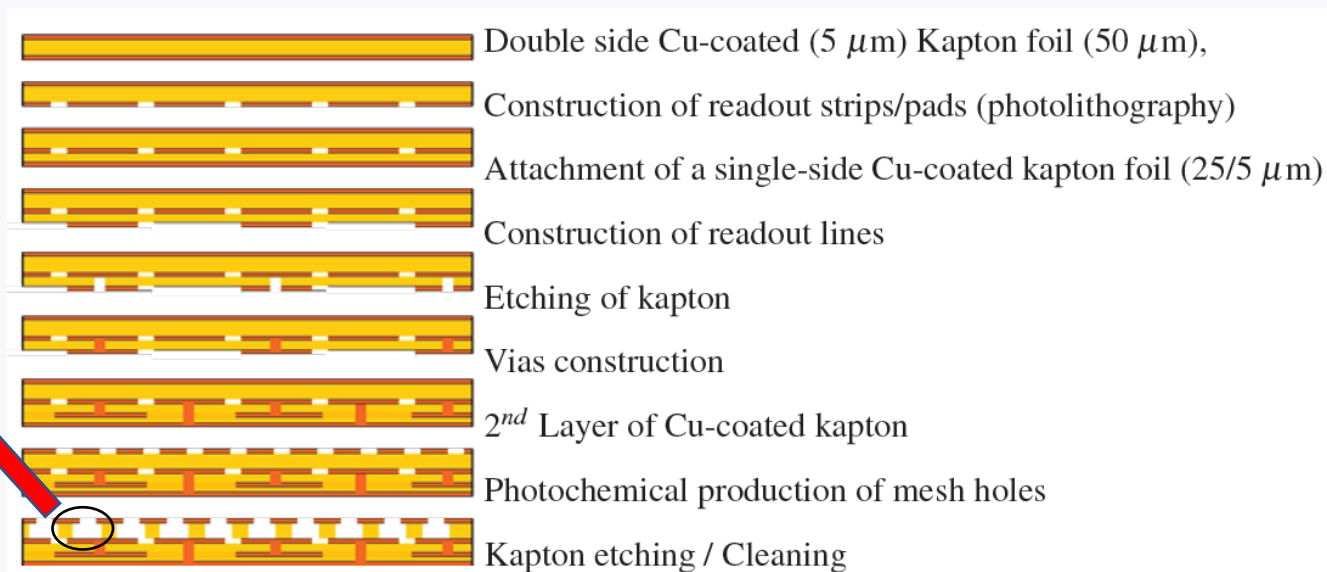
Microbulk Micromegas

- Made out of copper & polyimide (kapton)
 - potentially very radiopure
- High gap homogeneity
 - good energy resolution
 - Stability/homegeneity in response

Manufactured at Rui de Oliveira's workshop at CERN

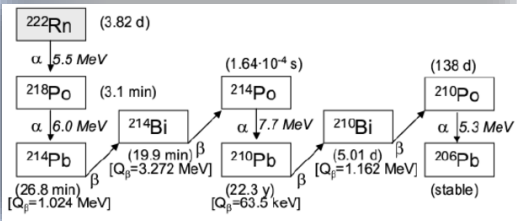


See for details:
JINST 5 (2010) P02001



Background status: Rn issue

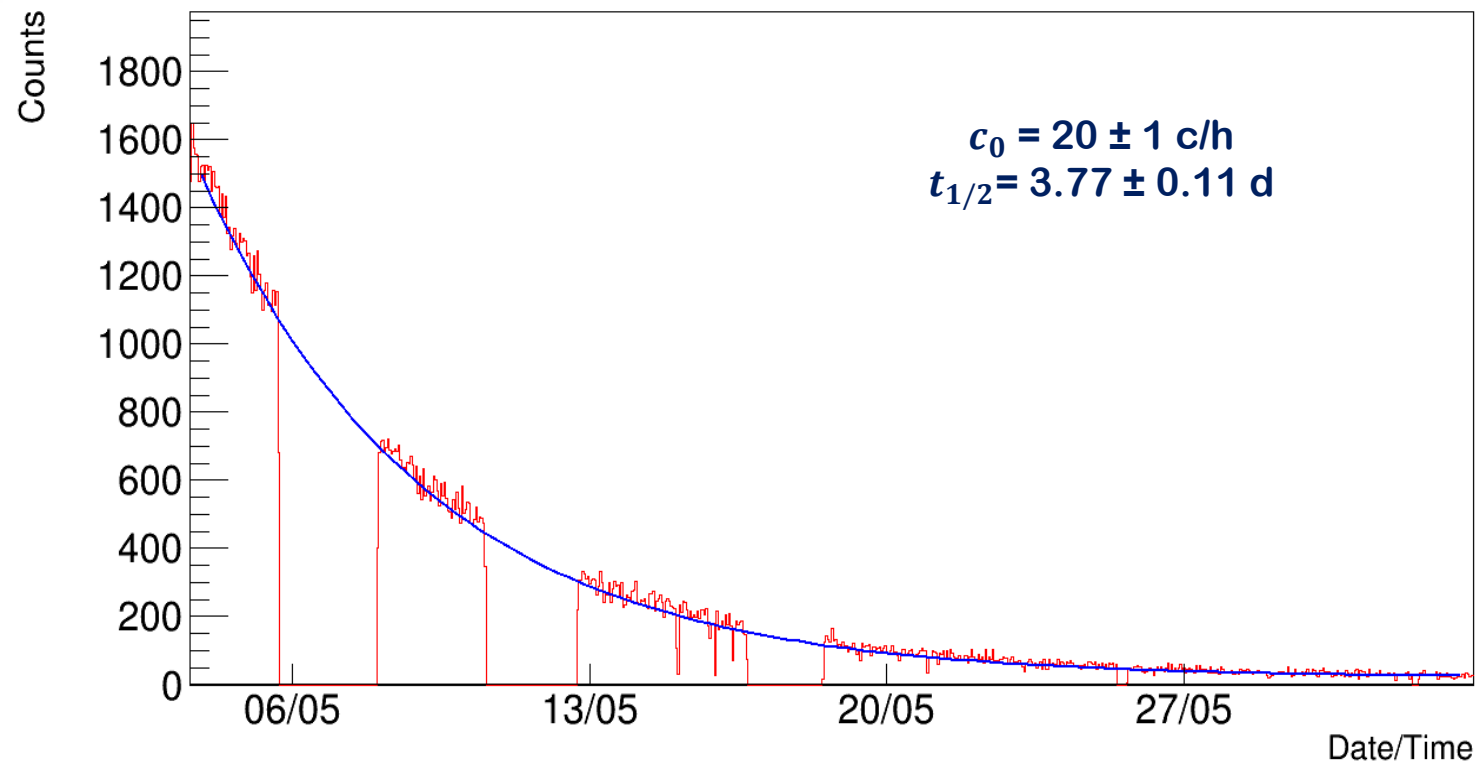
- LE and HE runs during 1 month in seal mode
- They supported the hypothesis of Rn + constant component: decrease in alpha rate and background rate in 0-50 keV after cuts



Low-gain runs (alphas)

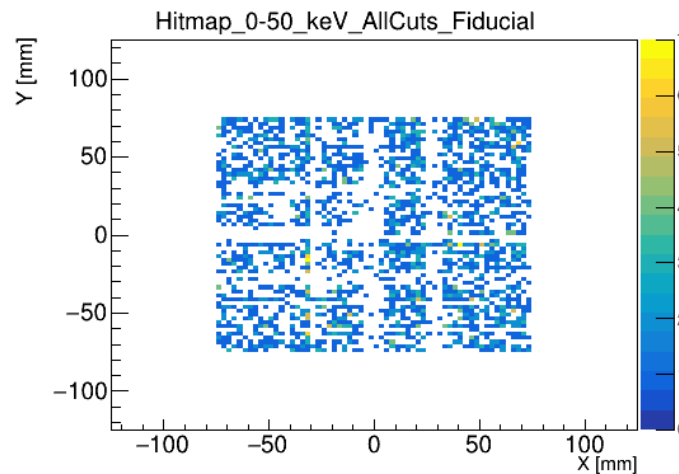
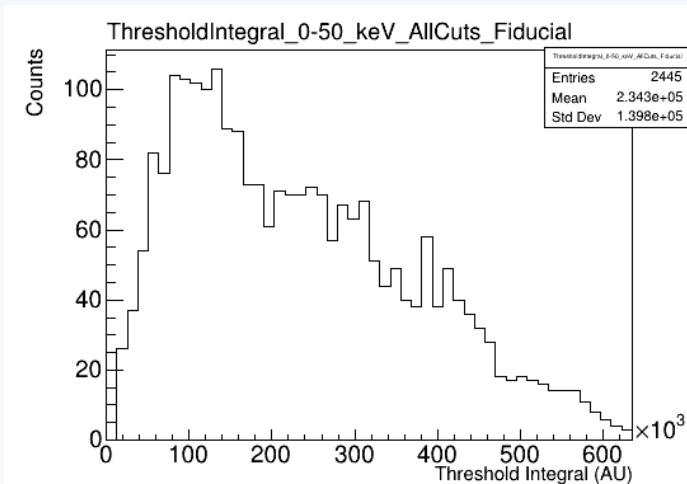
$$c_0 + c_1 e^{-c_3 t}$$

Timestamp_HECuts

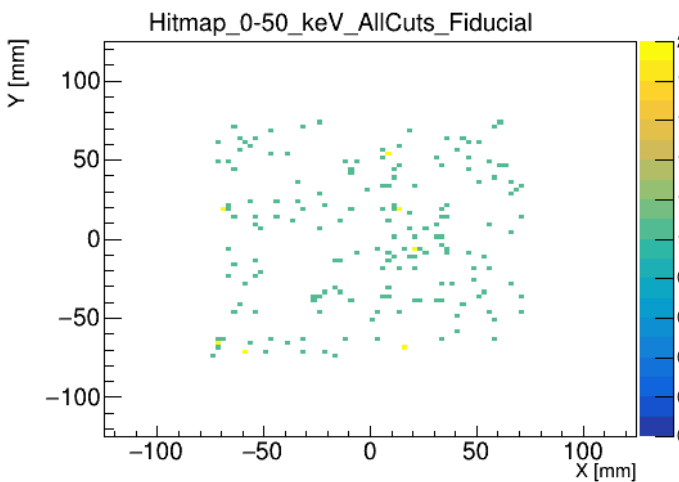
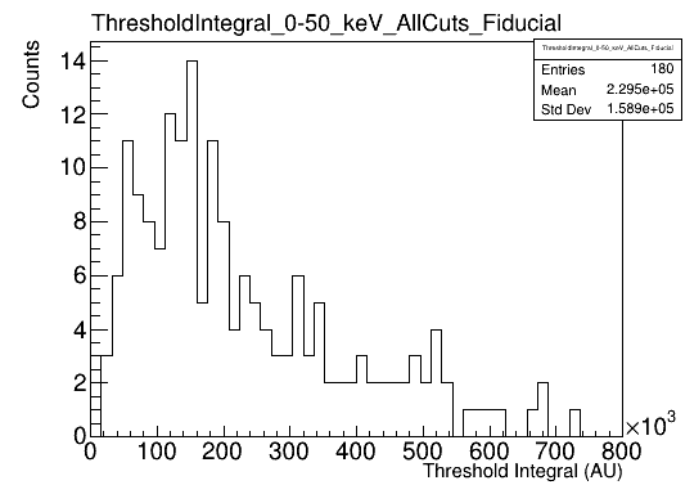


Background status: Rn issue

Background @ low energies



Before Rn decay
0-50 keV \rightarrow 600 dru

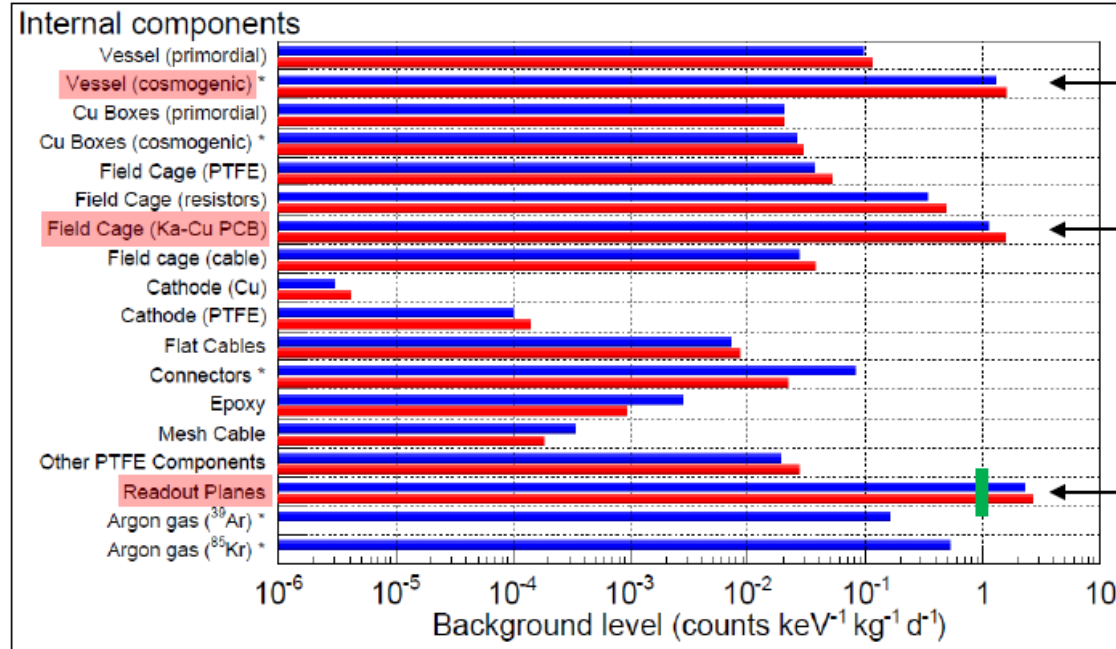


After Rn decay
0-50 keV \rightarrow 100 dru

Background status: Rn issue

- Internally emanated Radon is the main source of background (removing it takes us from ~600 dpu down to ~100 dpu in the 0-50 keV range)
- A lot of effort put into removing it from the system:
 - Trying with several commercial filters
 - Testing 5Å molecular sieves (we found out they do trap Rn, but emanate more than Agilent filters, best commercial filters we have)
 - Testing a custom-made O₂+H₂O filter developed by the University of Birmingham with low-emanation materials (ongoing collaboration with NEWS-G)
 - Testing activated carbon filters
 - **Open-loop operation bypassing the filters and the recirculation pump**
- Rn progeny surface contamination may well be responsible for the rest of background not accounted for in our background model
 - A program to identify alpha surface contaminations + its mitigation is ongoing

Challenges: Background level reduction



Next step once the background is dominated by the vessel activation.

New design on going.

Factor ~ 3 of contamination reduction respect to the previous Micromegas.

Main challenges in radio-purity of materials:

- Search of clean commercial materials. Large screening programs
- Synthesize clean materials.
- Control of processes in companies
- Storage in controlled environments

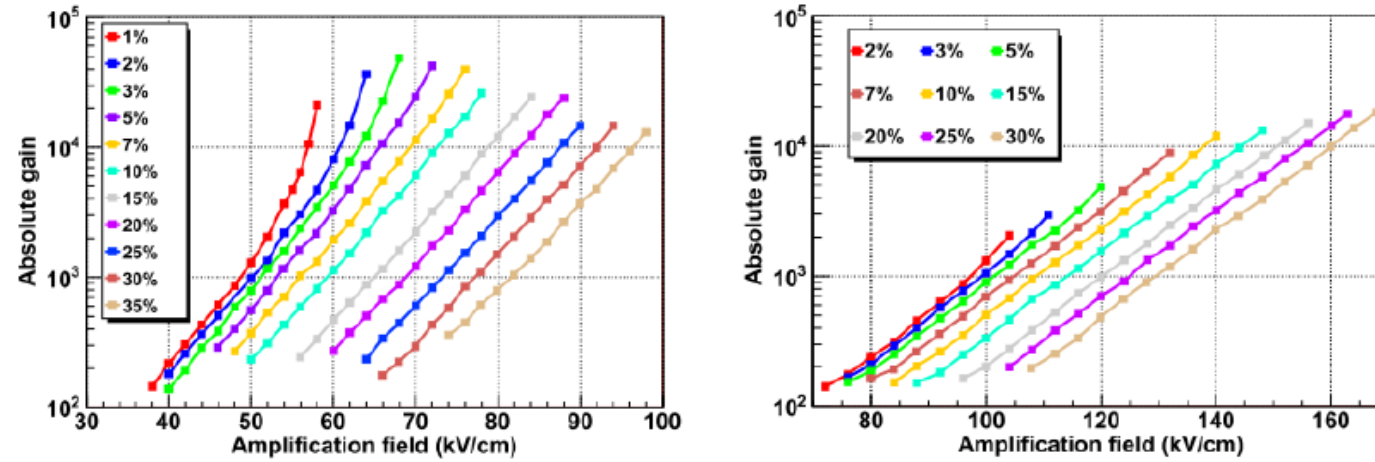


Figure 6. Dependence of the absolute gain with the amplification field for two microbulk detectors with gaps of 50 (left) and 25 μm (right) in argon-isobutane mixtures. The maximum gain of each curve was obtained just before the spark limit. The percentage of each series corresponds to the isobutane concentration.

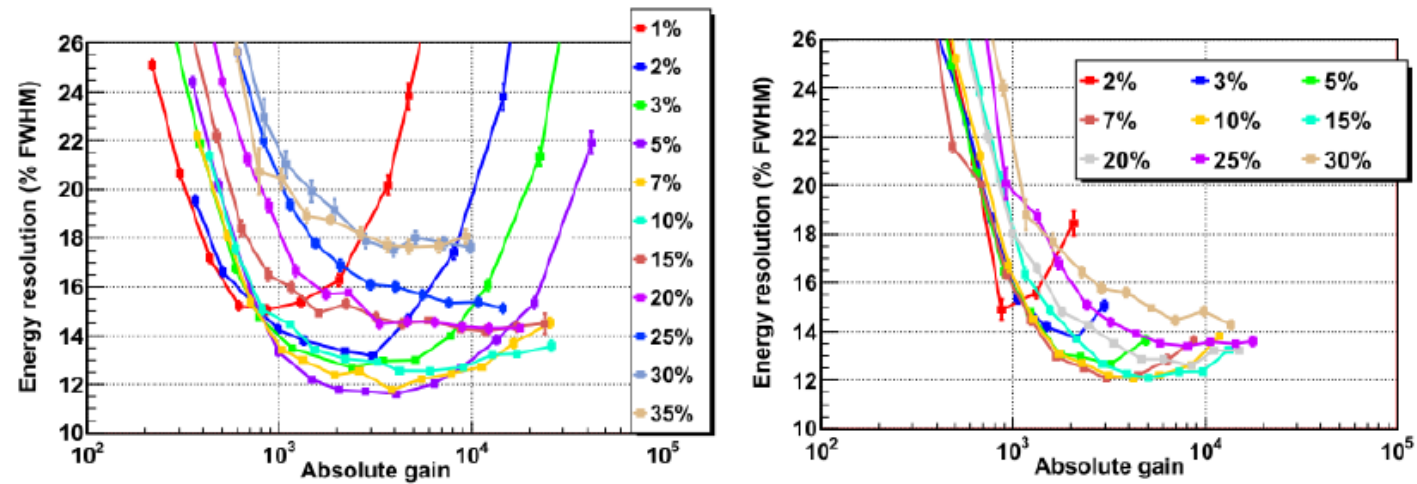


Figure 7. Dependence of the energy resolution with the absolute gain for two detectors of 50 (left) and 25 μm -thickness-gap (right) in argon-isobutane mixtures. The maximum gain of each curve was obtained just before the spark limit. The percentage of each series corresponds to the isobutane concentration.