

# zELDA : fitting Lyman – alpha line profiles using deep learning .



redshift Estimator using Line profiles  
of Distant Lyman-Alpha emitters

S . Gurung – Lopez, C . Byrohl and M . Gronke

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© Detect Galaxies at high redshift (LAEs)

- HETDEX (Hill et al. 2008 ) @  $1.9 < z < 3.5$  (spectroscopy)
- SILVERRUSH (Ouchi et al. 2018 ) ~2k LAEs @  $5 < z < 7$  (photometry)
- J-PLUS (Spinoso et al. 2020) ~14k LAEs @  $2 < z < 2.3$  (photometry)

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- Ly $\alpha$  as a star formation rate indicator (similar to H $\alpha$ ) : (Sobral et al. 2019)
- Tracer of Lyman continuum photons emitters : (Steidel et al. +18 , Izotov et al.+21)

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- Outflows : Bresolin et al. 2017, Tumlinson et al 2017

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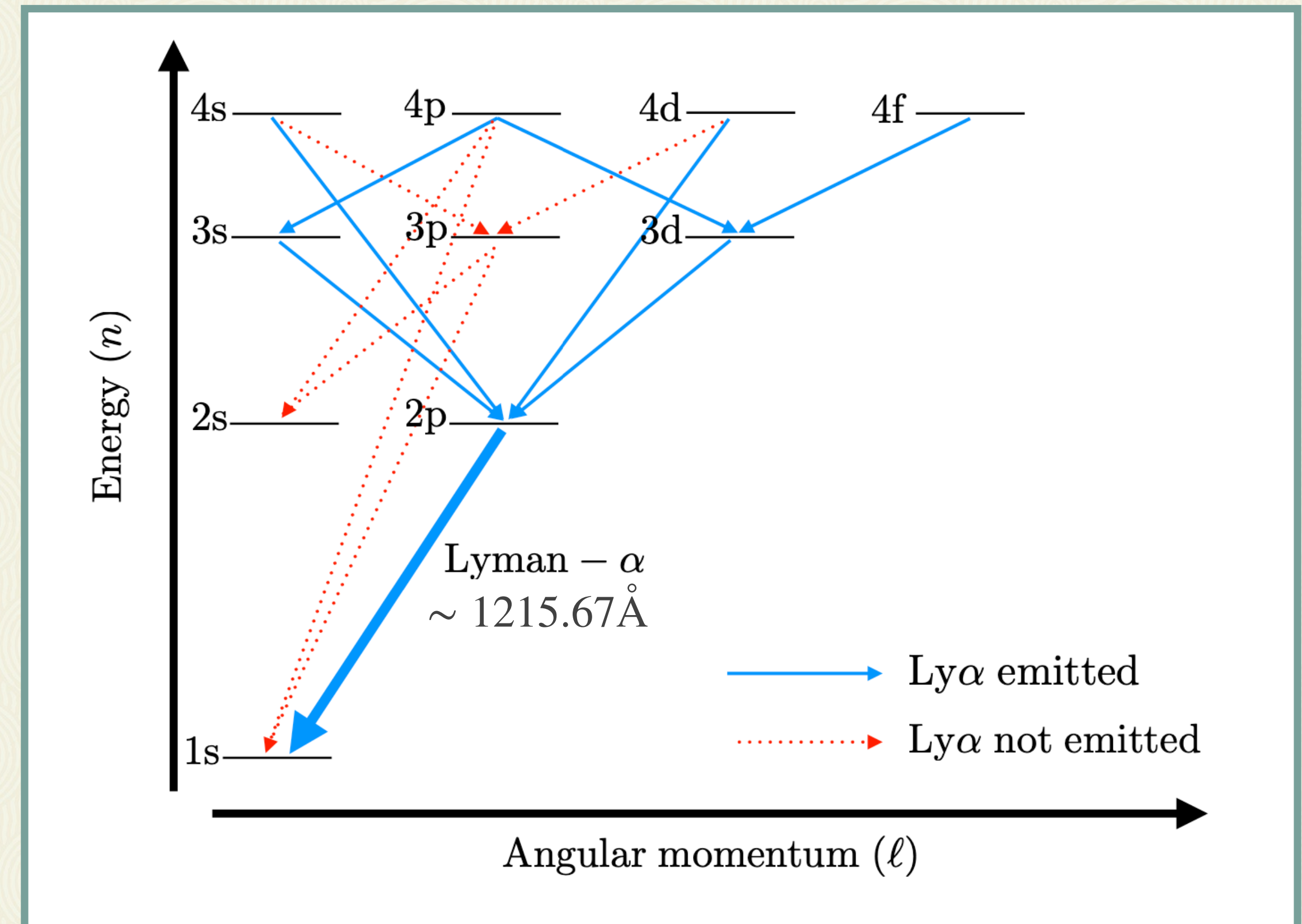
- Ly $\alpha$  halos : Leclercq et al. 2017, Erb et al. 2018

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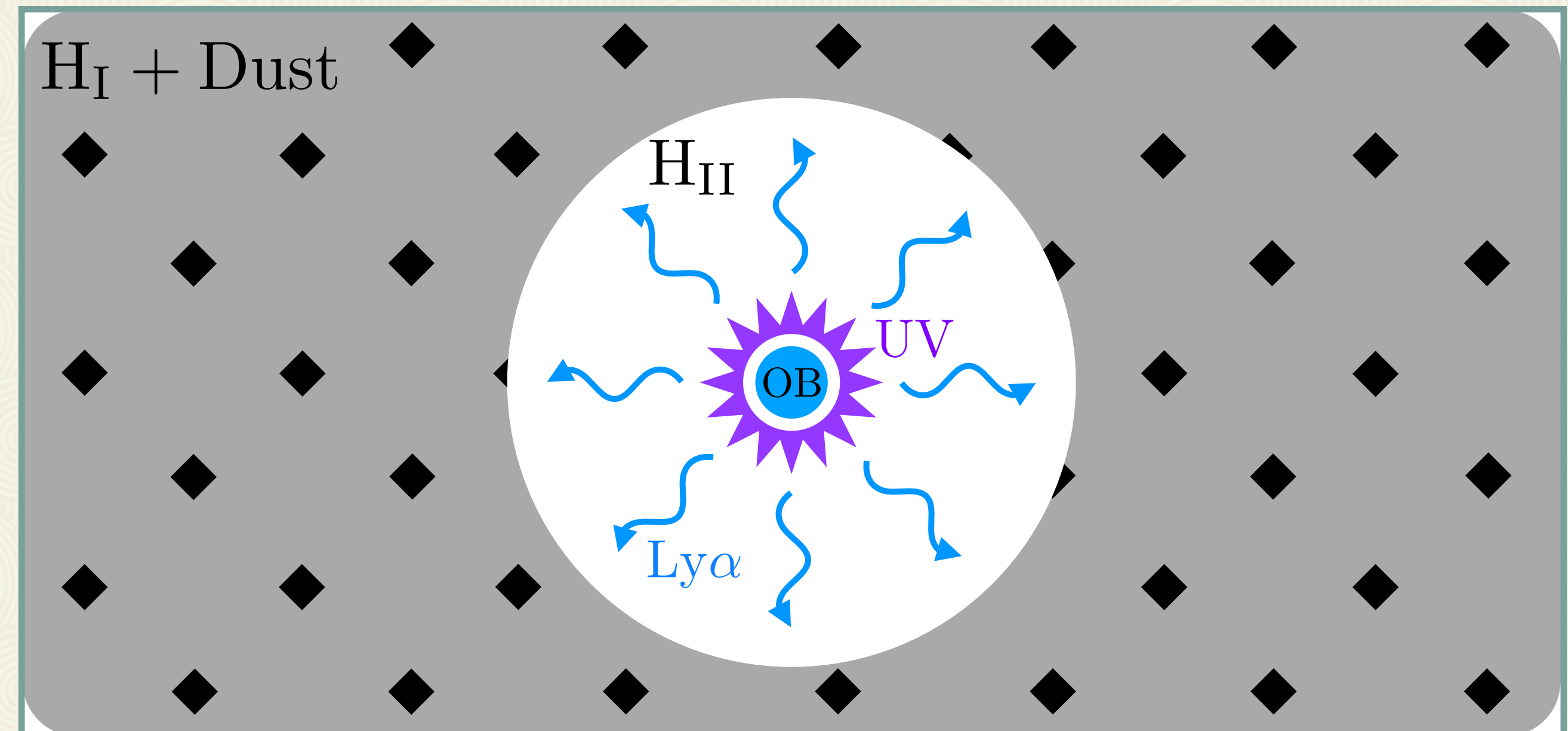
- Ly $\alpha$  halos : Leclercq et al. 2017, Erb et al. 2018

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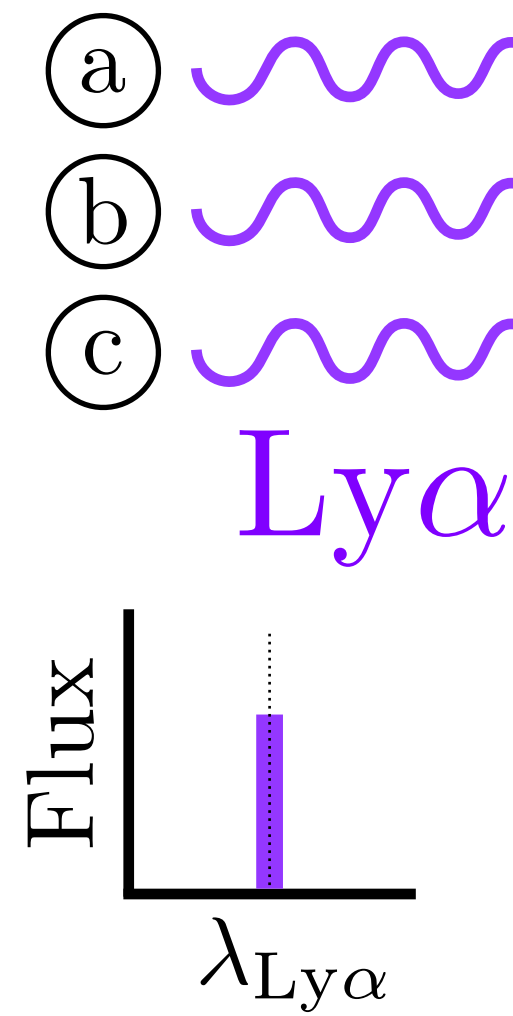
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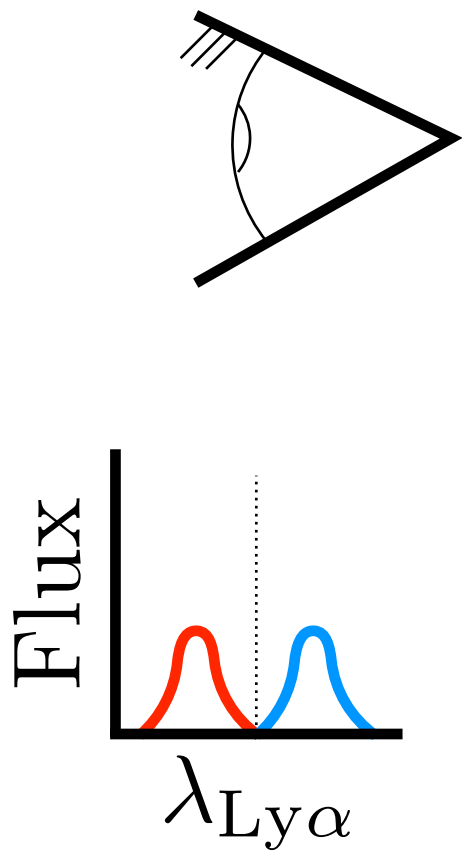
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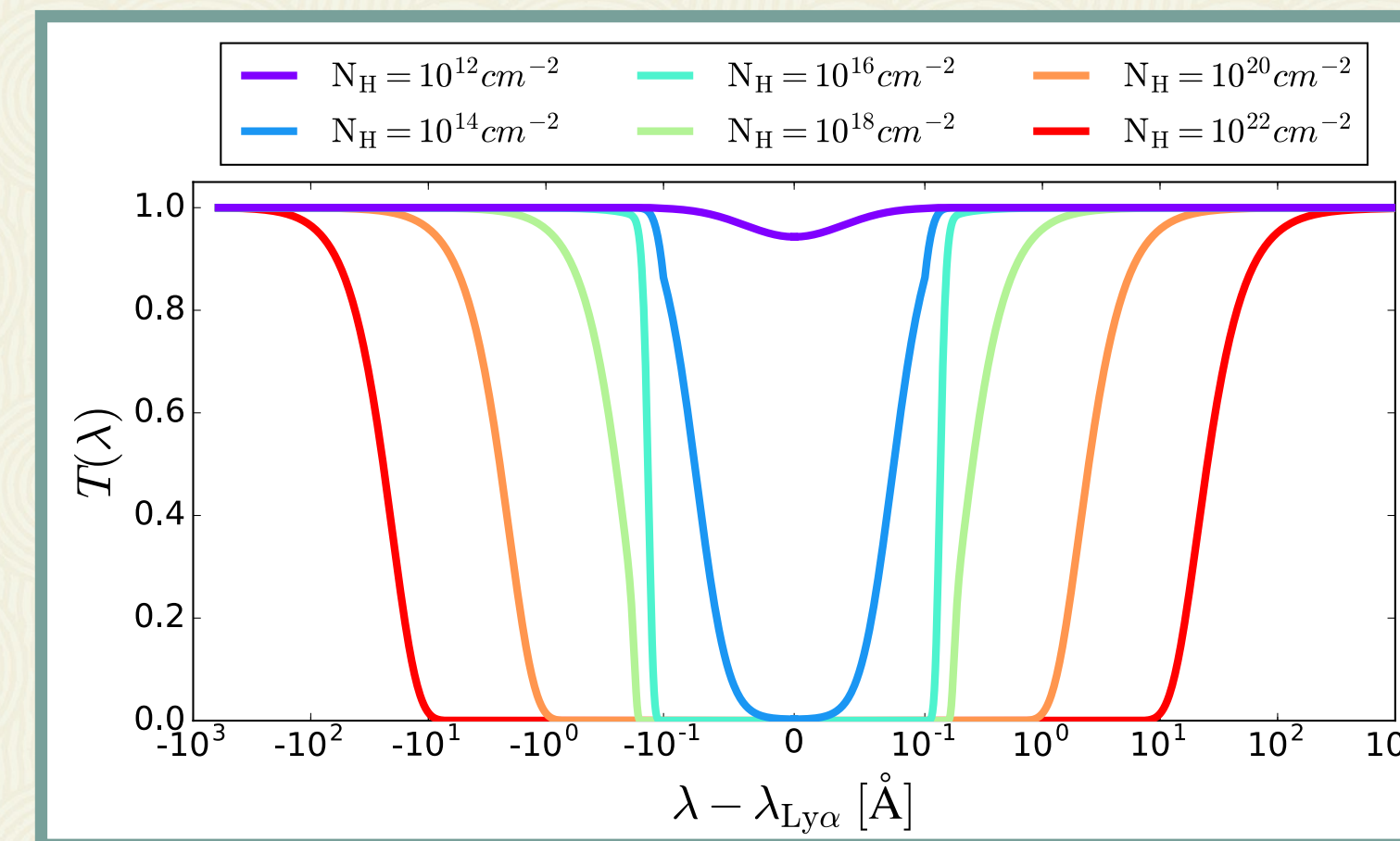
## Emitted



## Observed



- ◆ Dust interaction
- Hydrogen interaction



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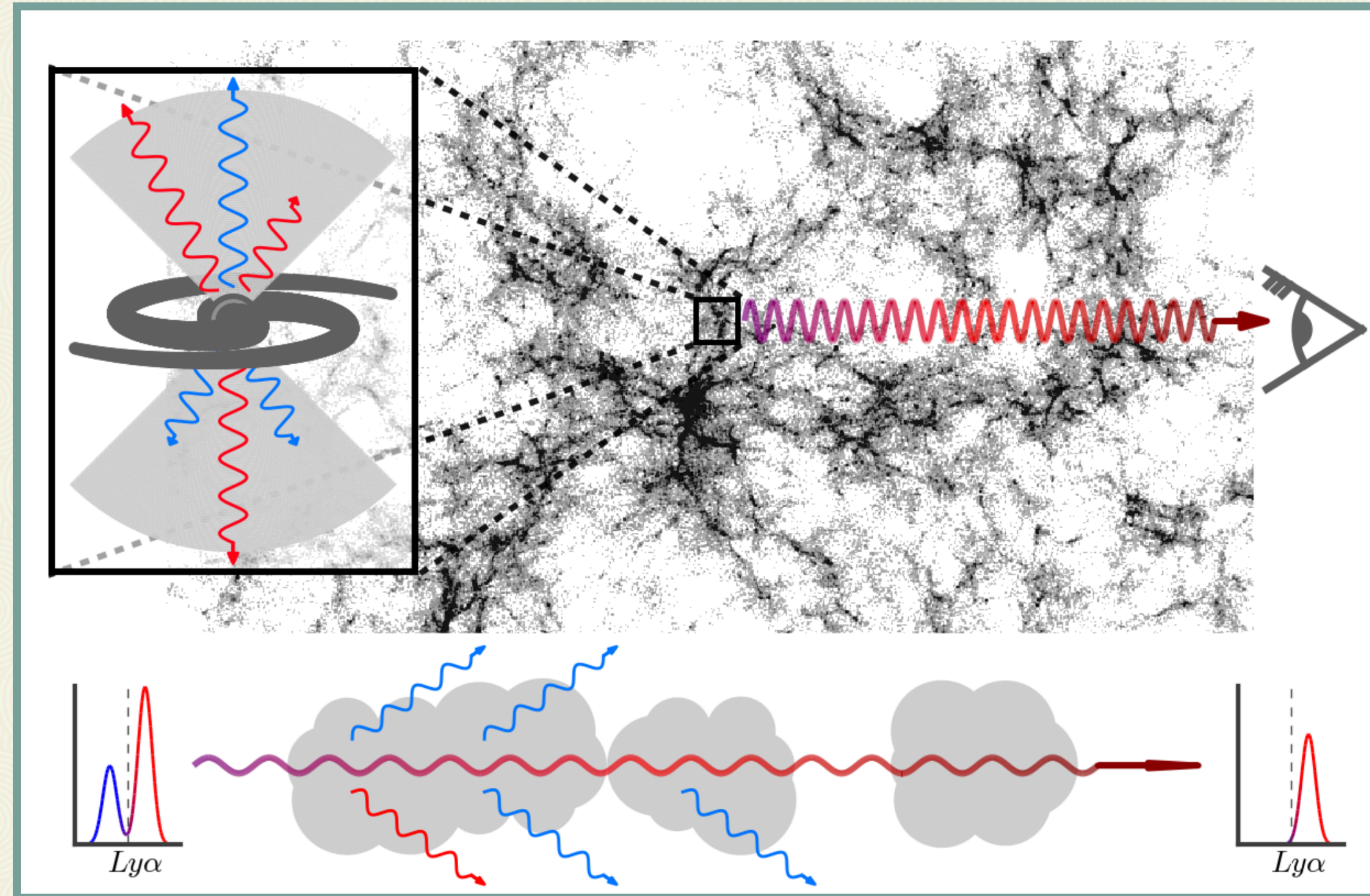
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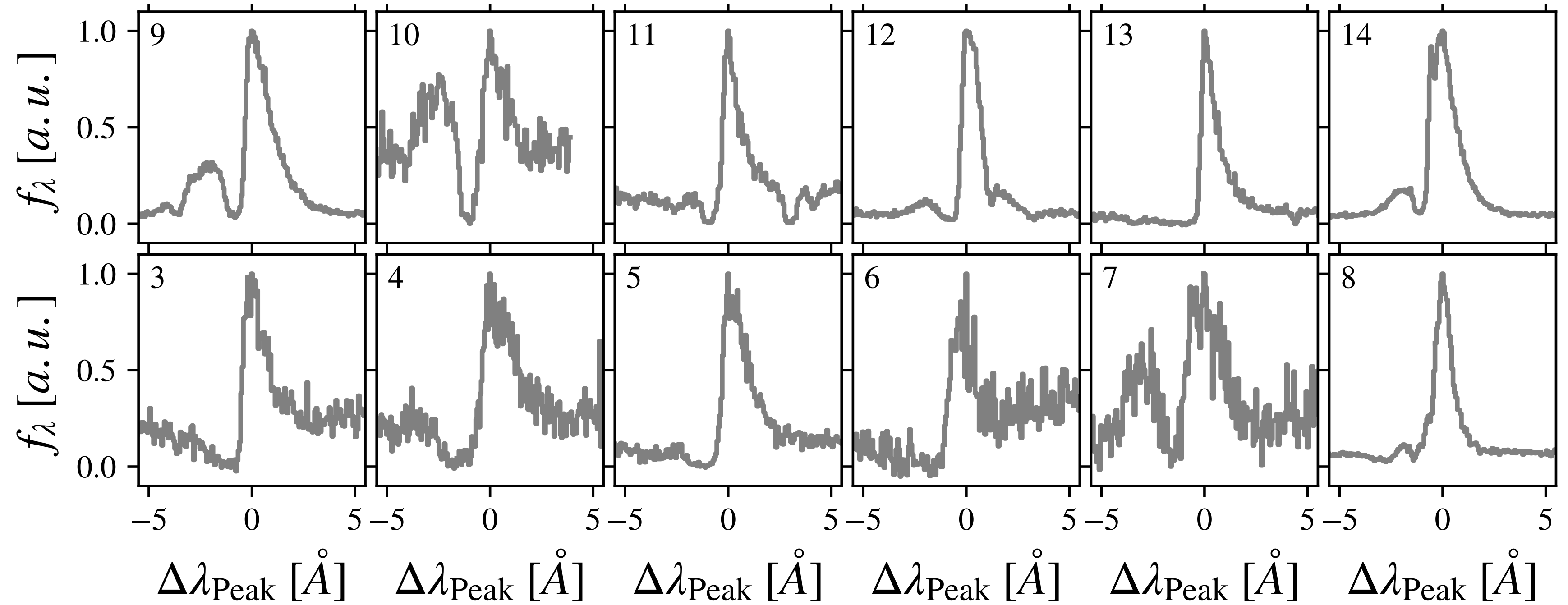




<http://lasd.lyman-alpha.com>

Runnholm et al.+20

97 Ly $\alpha$  line profiles obtained by the Cosmic Origins Spectrograph on HST  
 GO 11522 and 12027 (PI: Green, Salzer et al. 2001)  
 GO 11727 and 13017 (PI: Heckman et al. 2011,2015)  
 GO 12269 (PI: Scarlata, Songaila et al. 2018)  
 GO 12583 (PI: Hayes, Hayes et al. 2014)  
 GO 12928 (PI: Henry, Henry et al. 2015),  
 GO 13293 and 14080 (PI: Jaskot, Jaskot & Oey 2014)  
 GO 14201 (PI: Malhotra, Yang et al. 2017)  
 GO 13744 (PI: Thuan Izotov et al. 2016, 2018)



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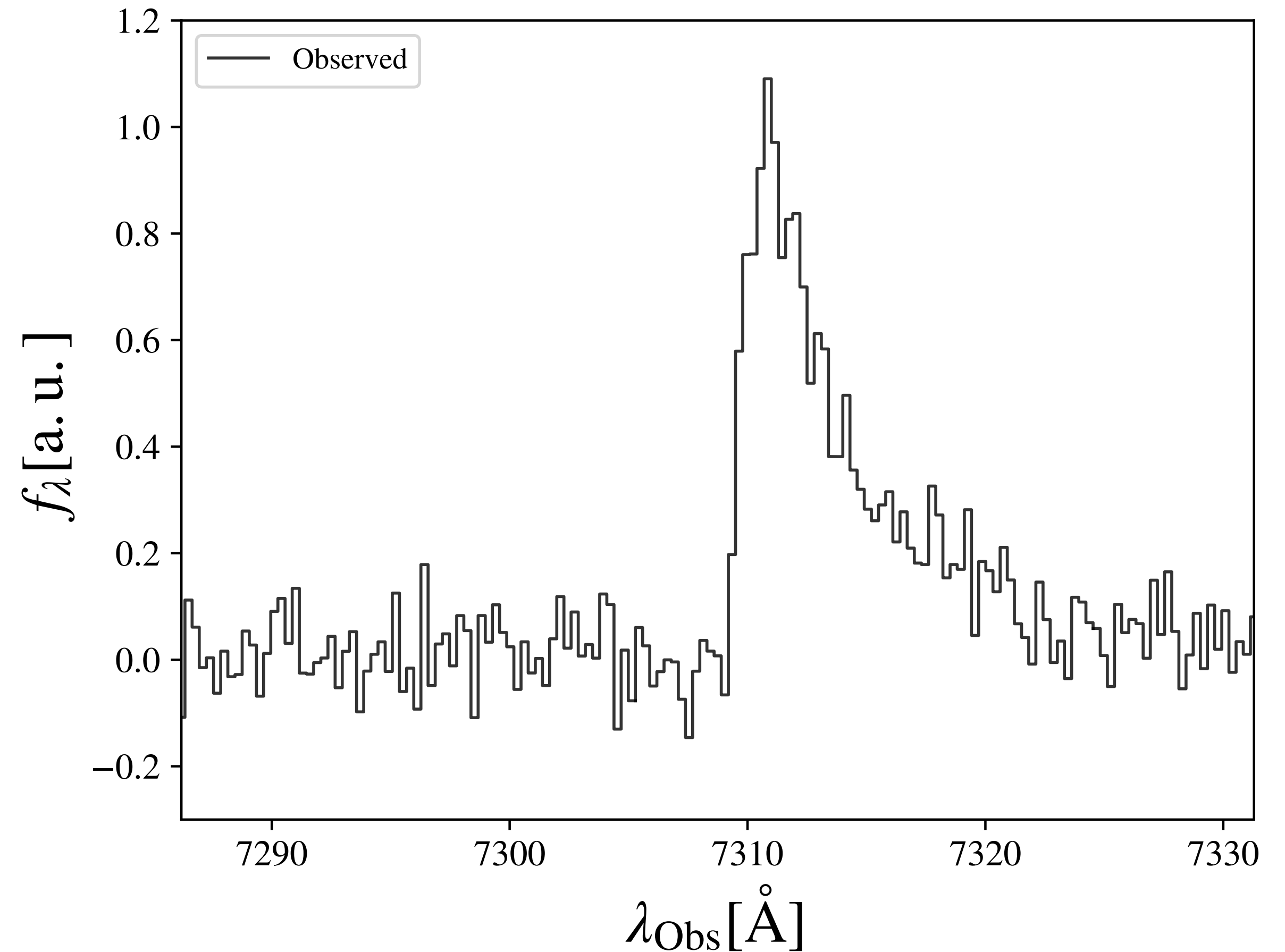
- Ly $\alpha$  halos : Leclercq et al. 2017, Erb et al. 2018

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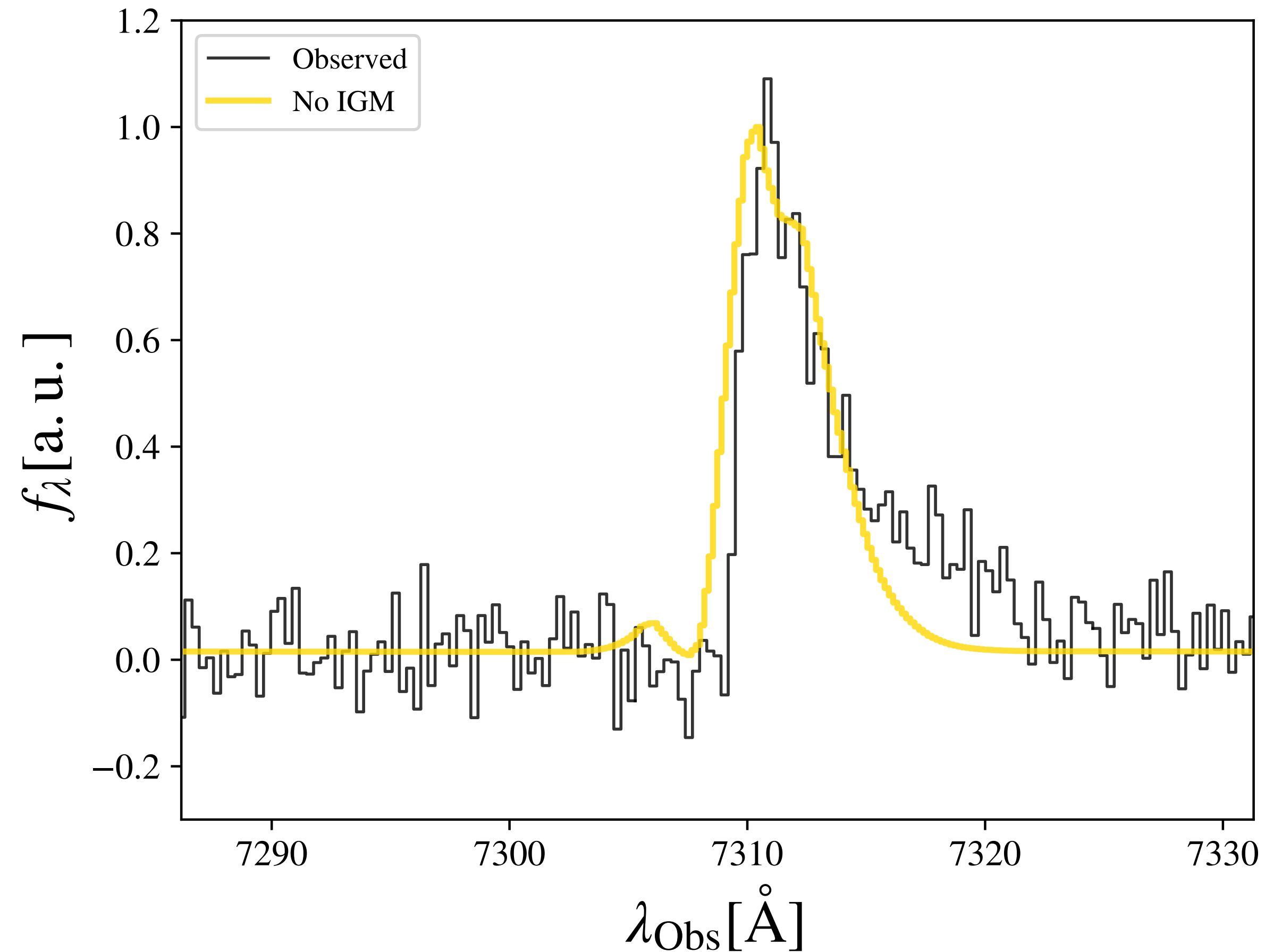
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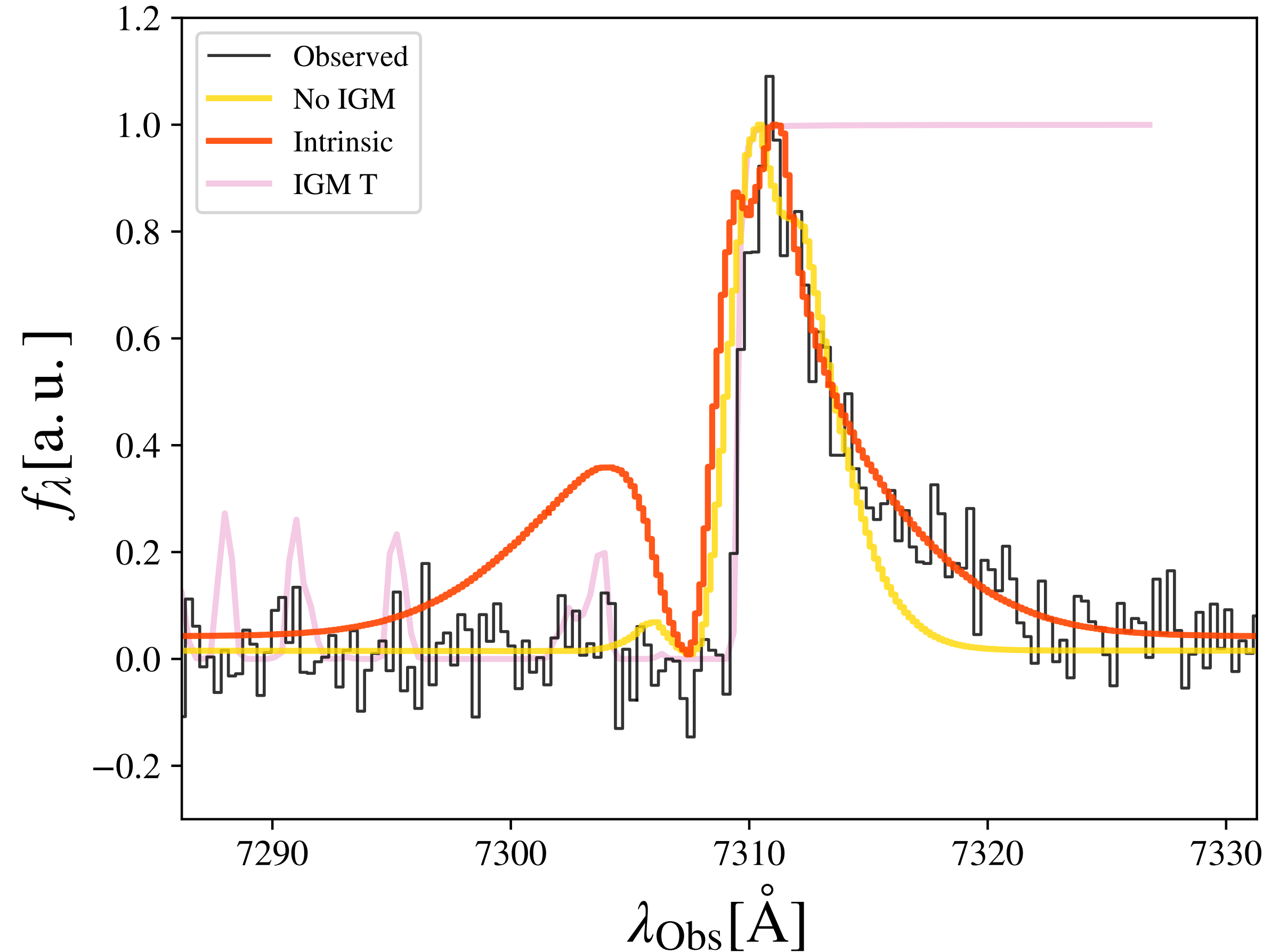
- Ly $\alpha$  halos : Leclercq et al. 2017, Erb et al. 2018

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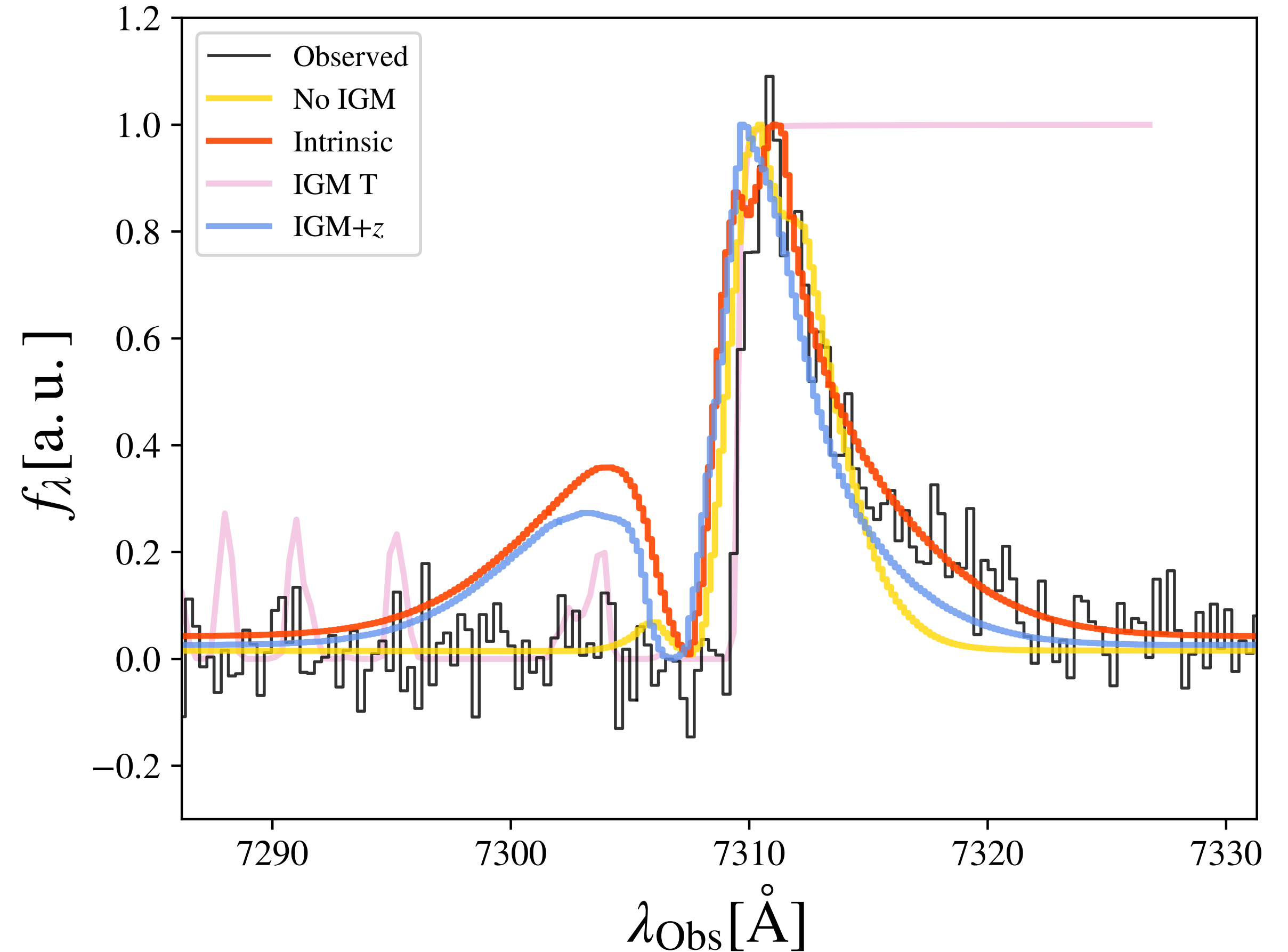
- Ly $\alpha$  halos : Leclercq et al. 2017, Erb et al. 2018

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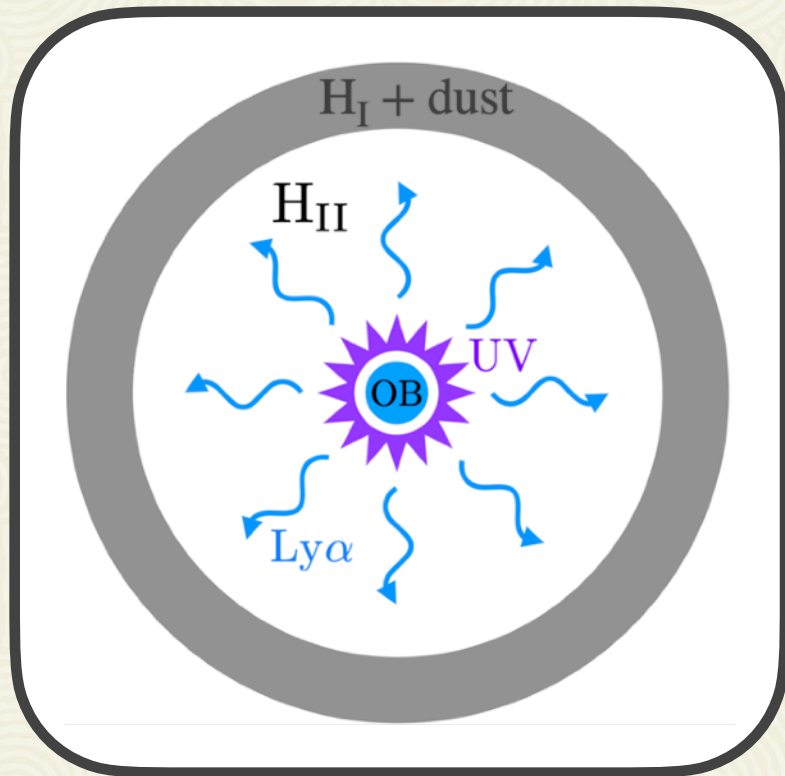
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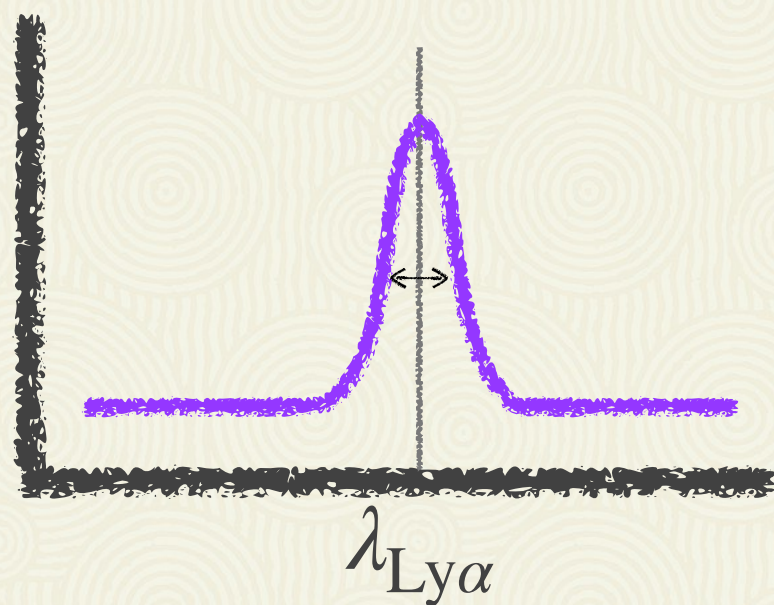


## Gas properties



- Gas Geometry
- Column Density
- Expansion Velocity
- Dust content

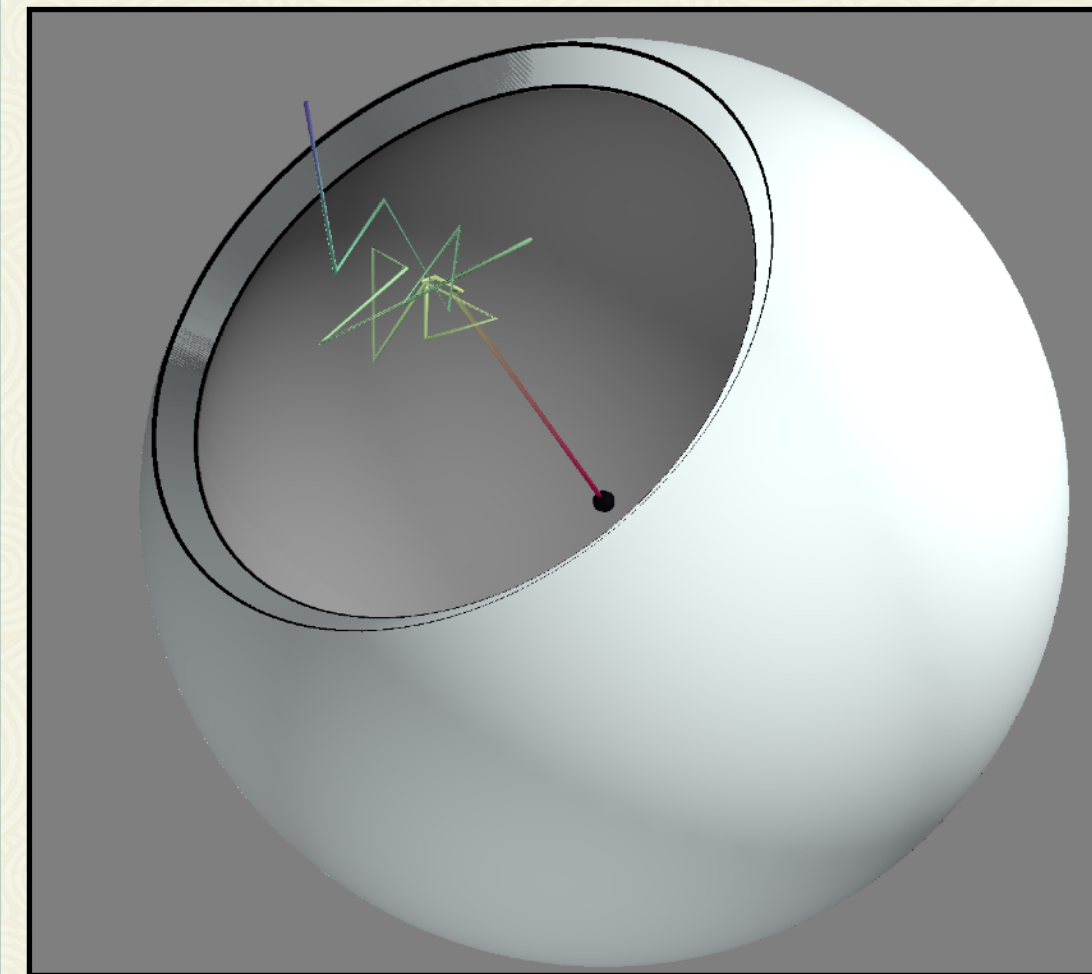
## Injected Spectrum



- Line width
- Equivalent width

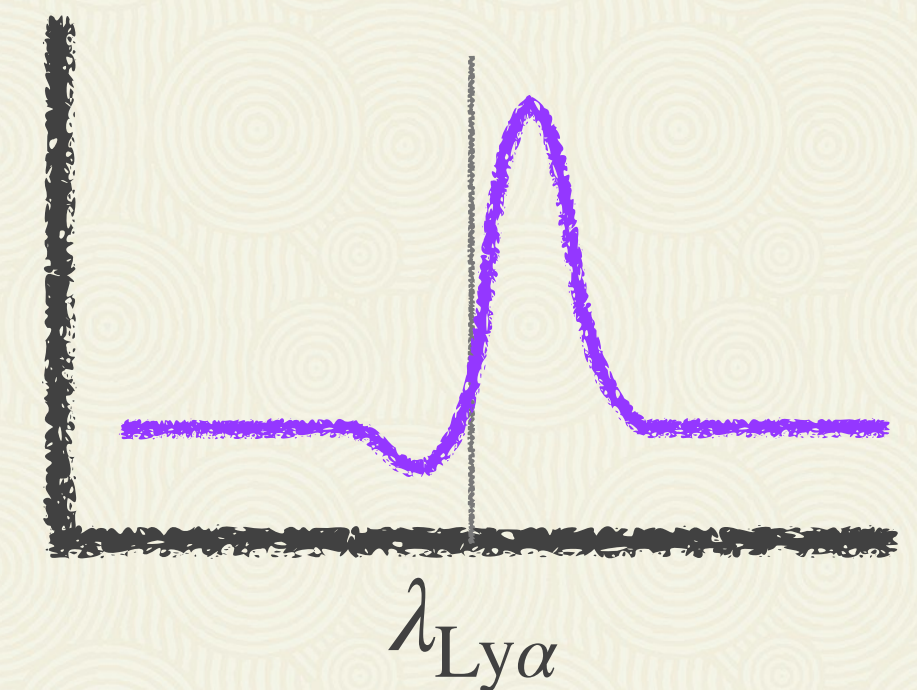
$N_H$   
 $V_{exp}$   
 $\tau_a$   
 $W_{in}$   
 $EW_{in}$   
 Model parameters

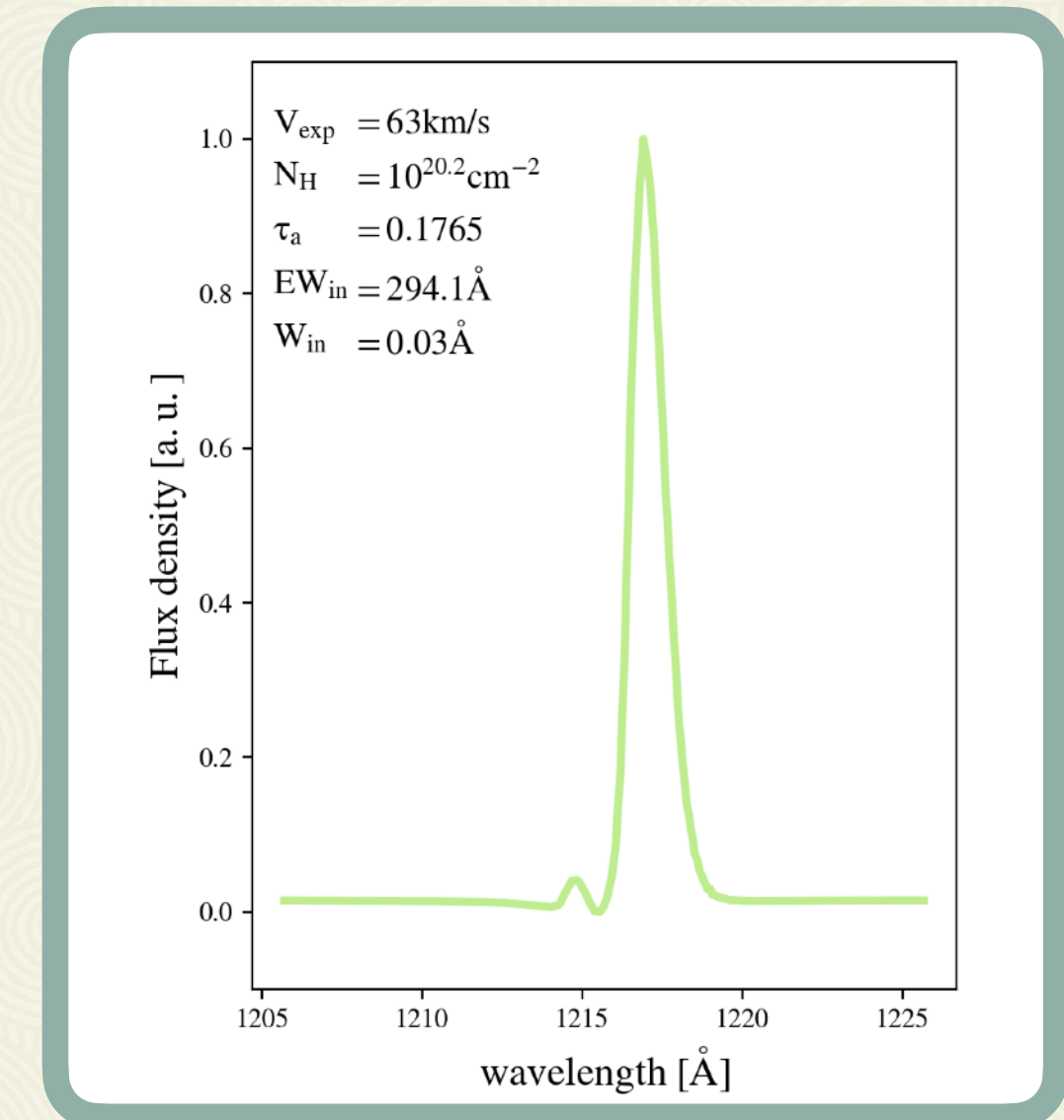
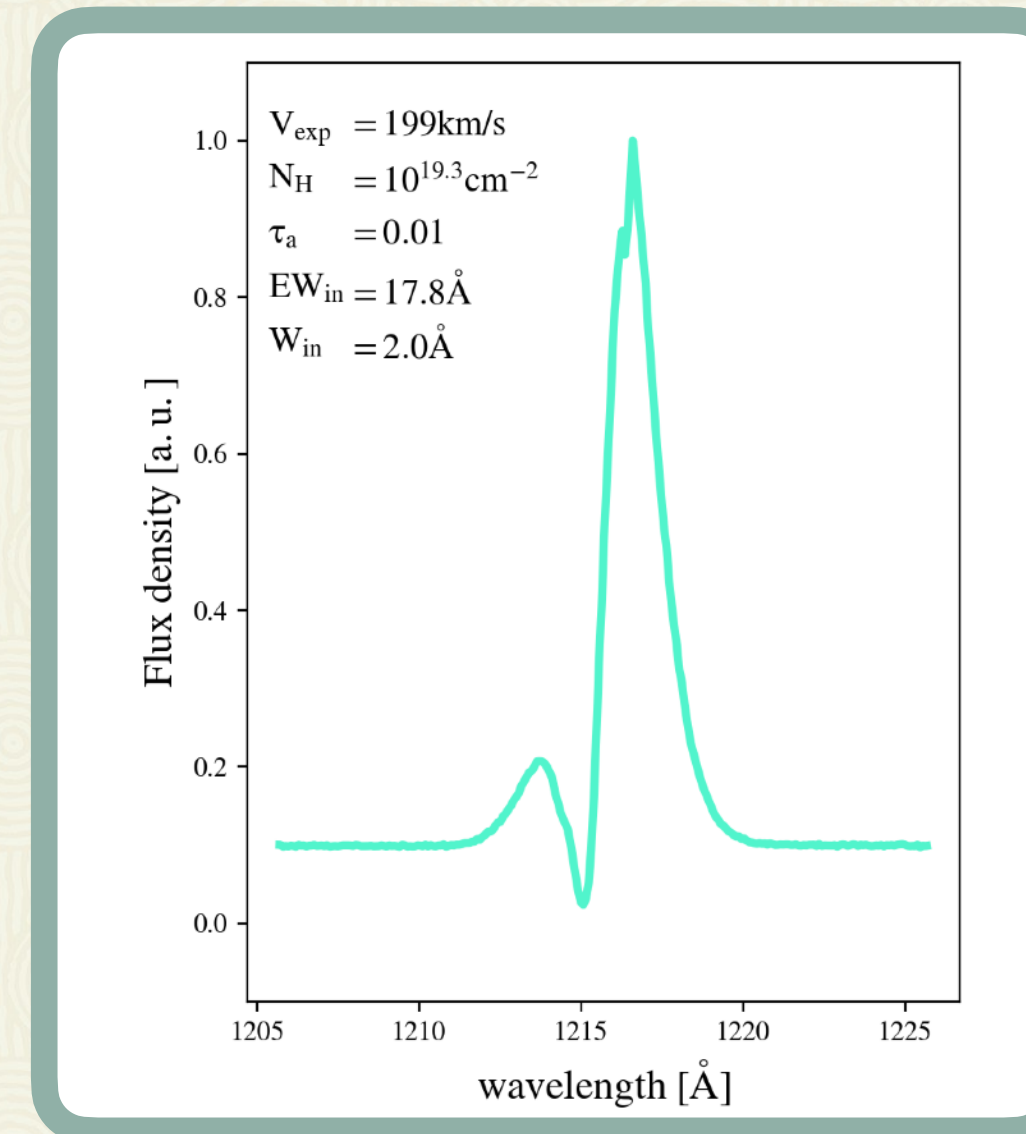
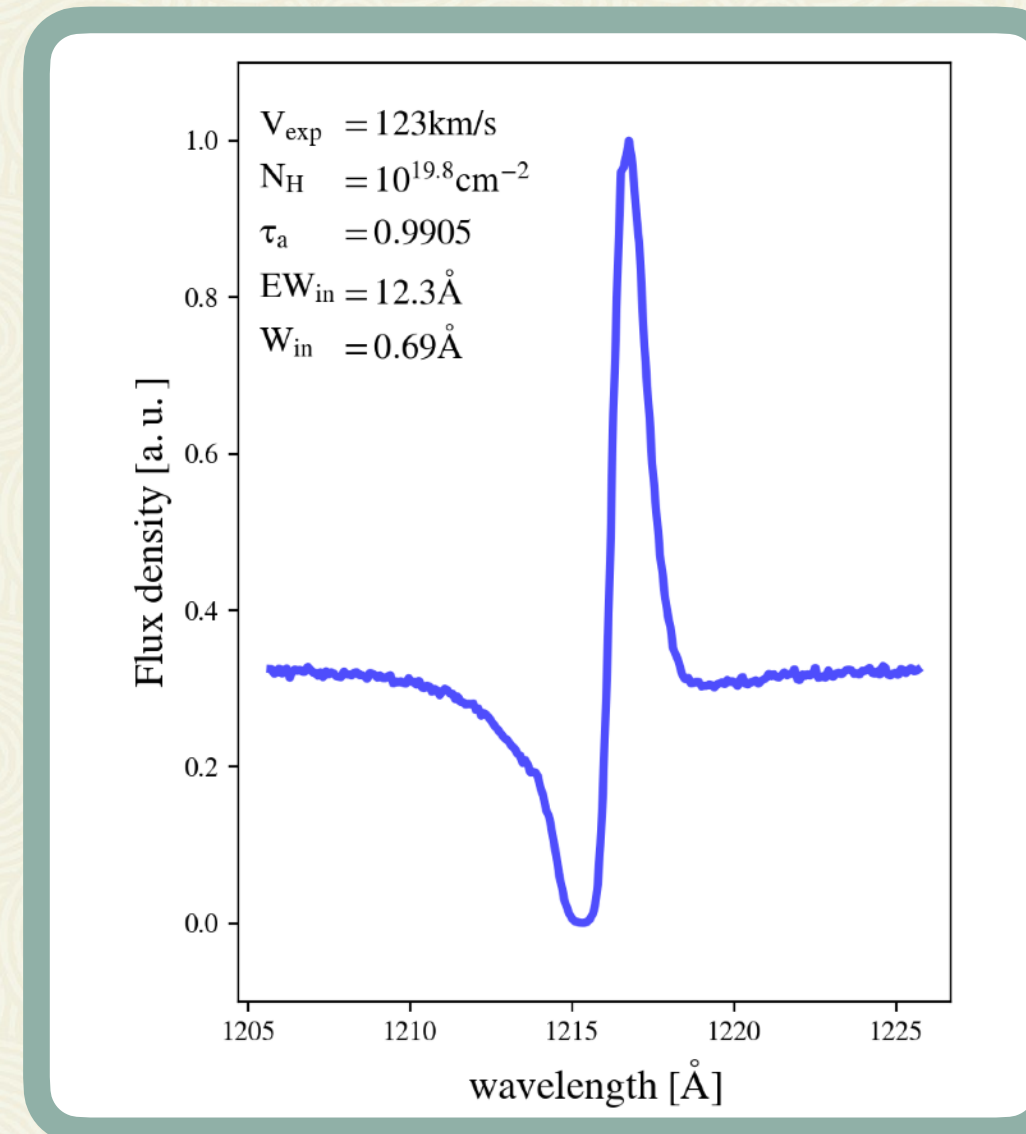
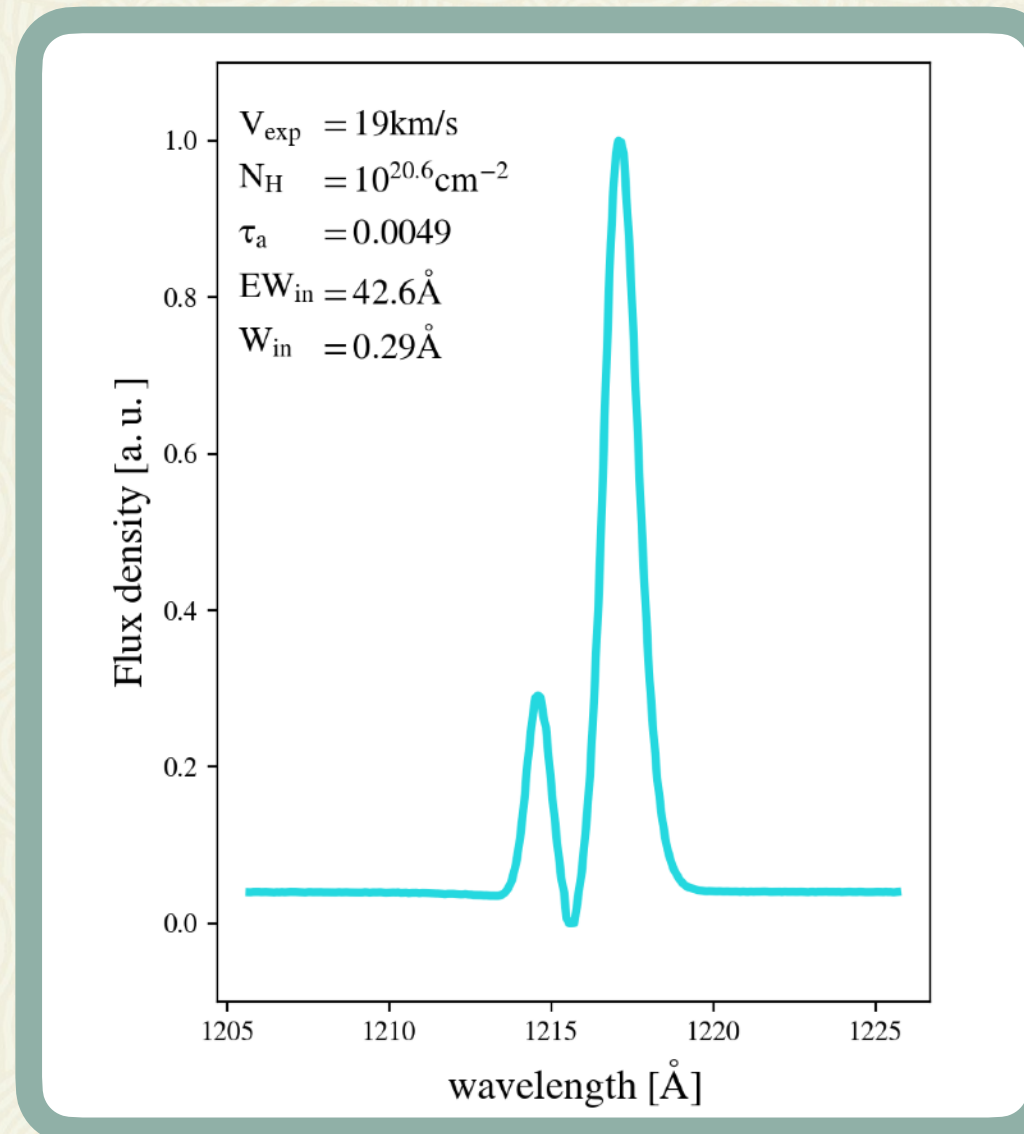
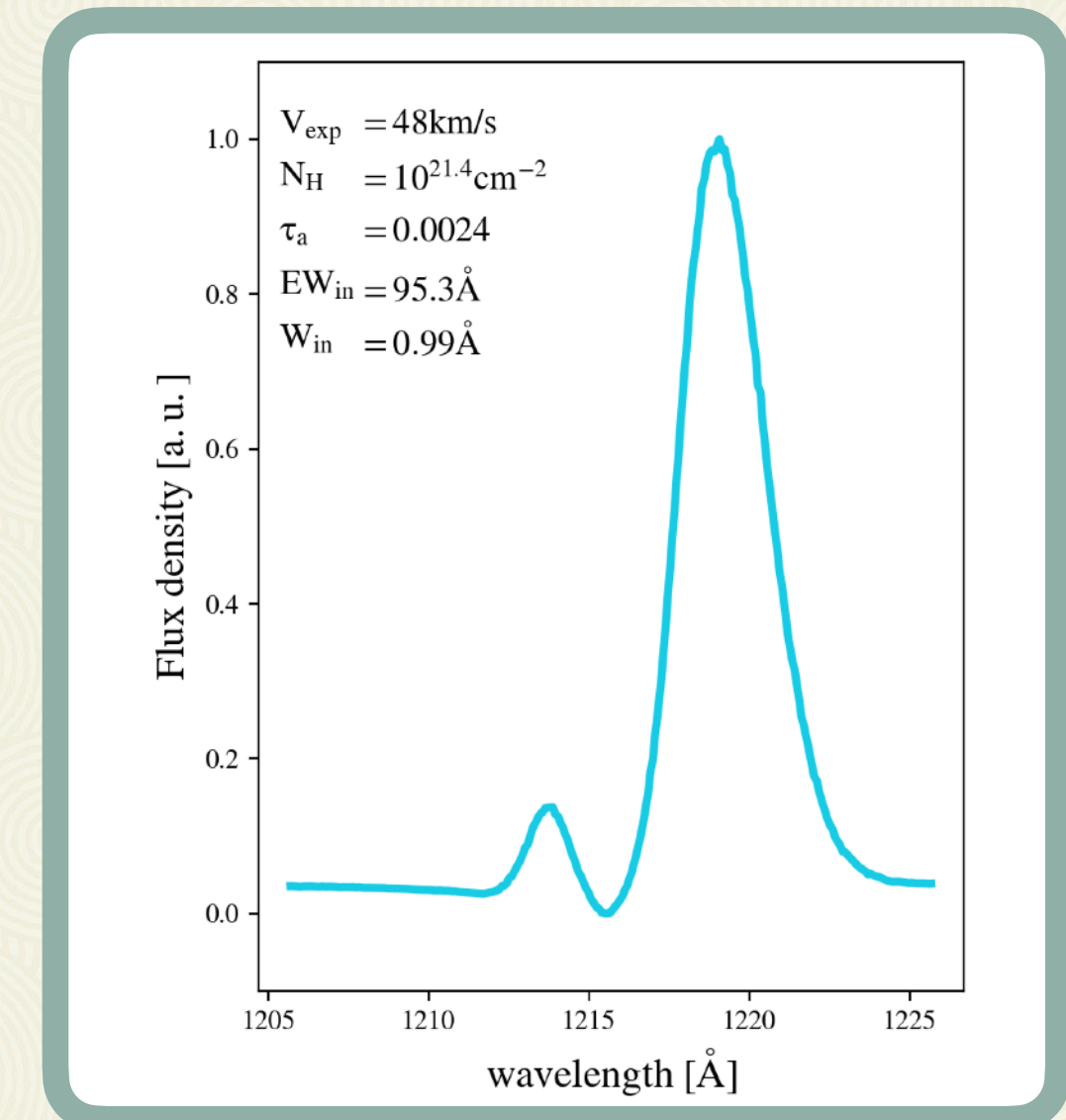
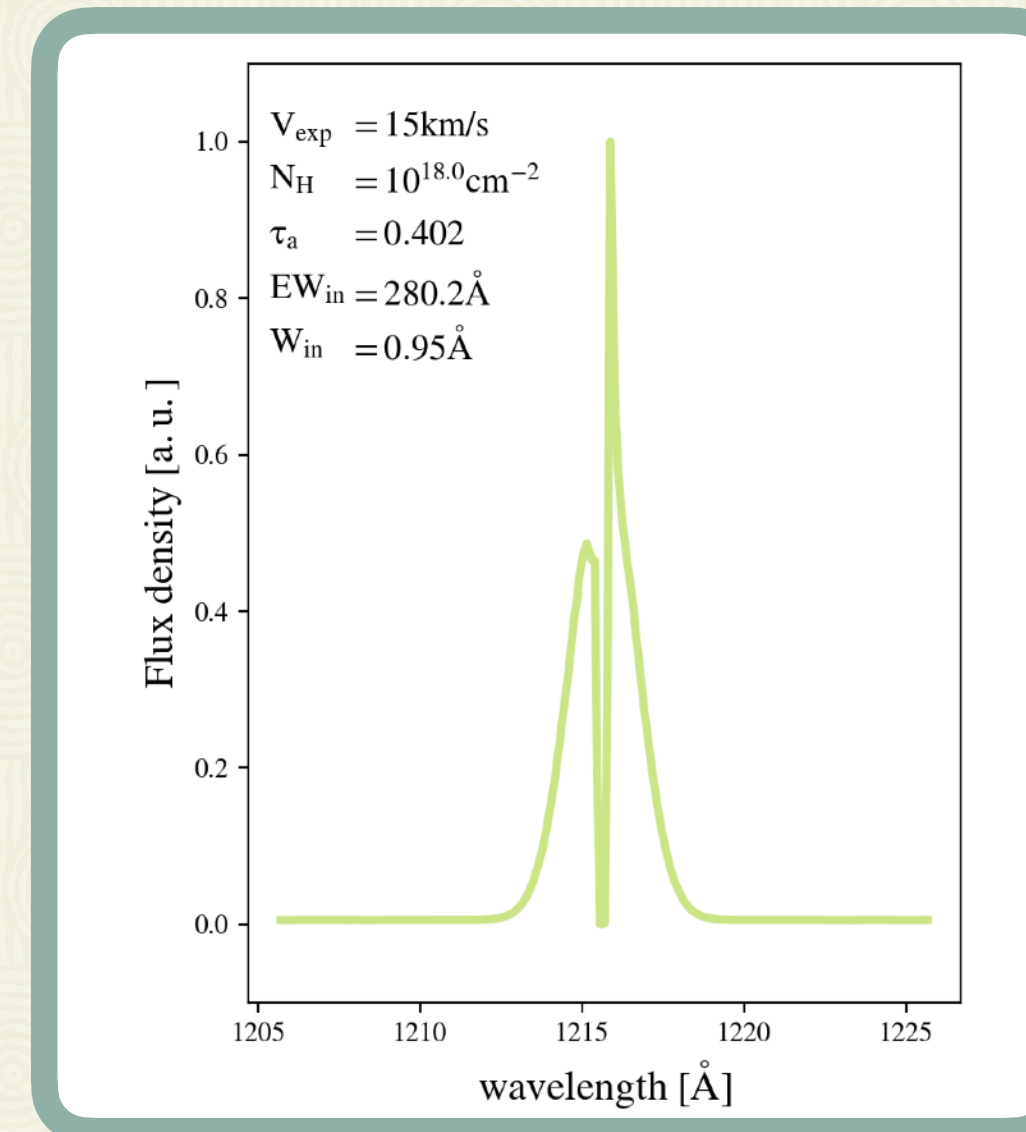
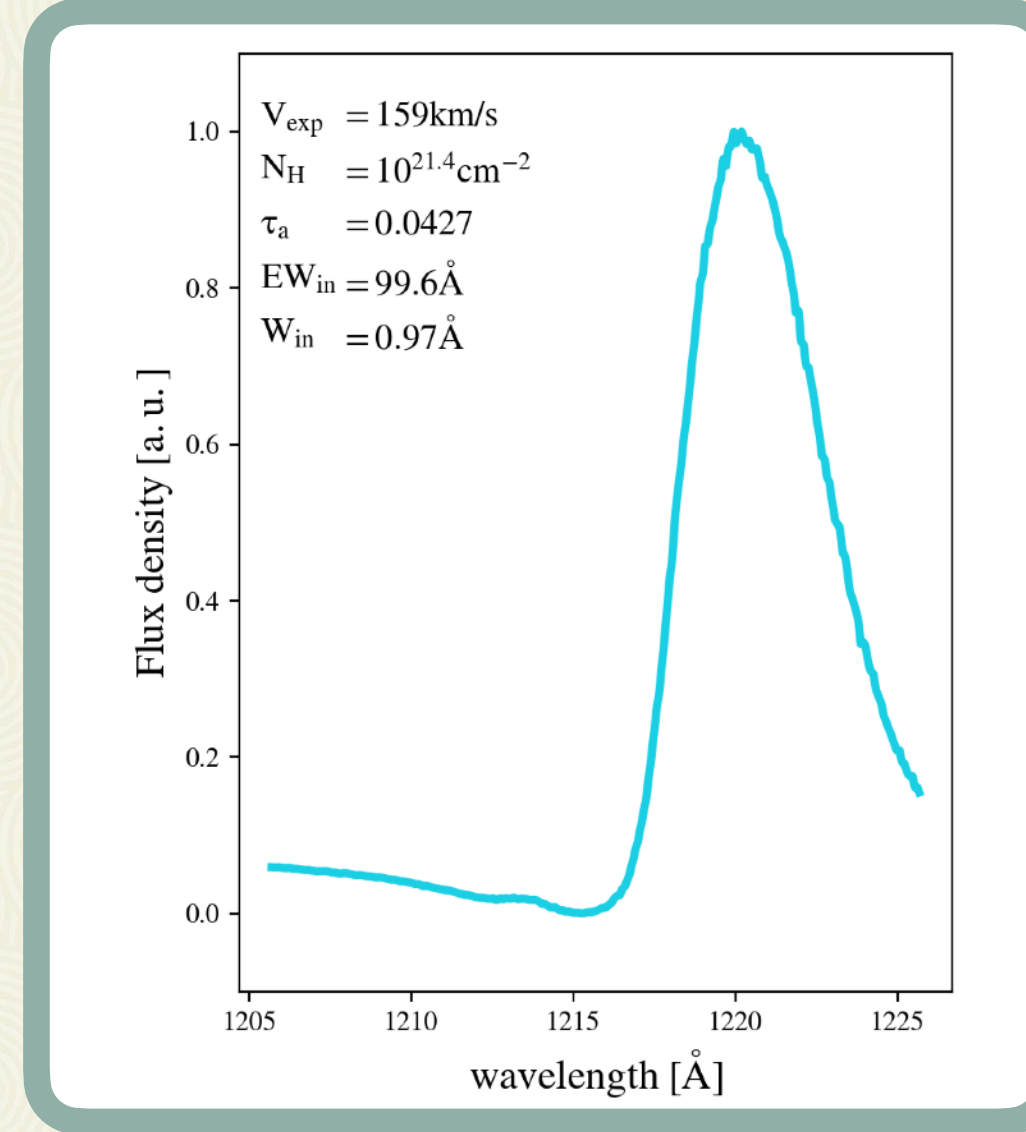
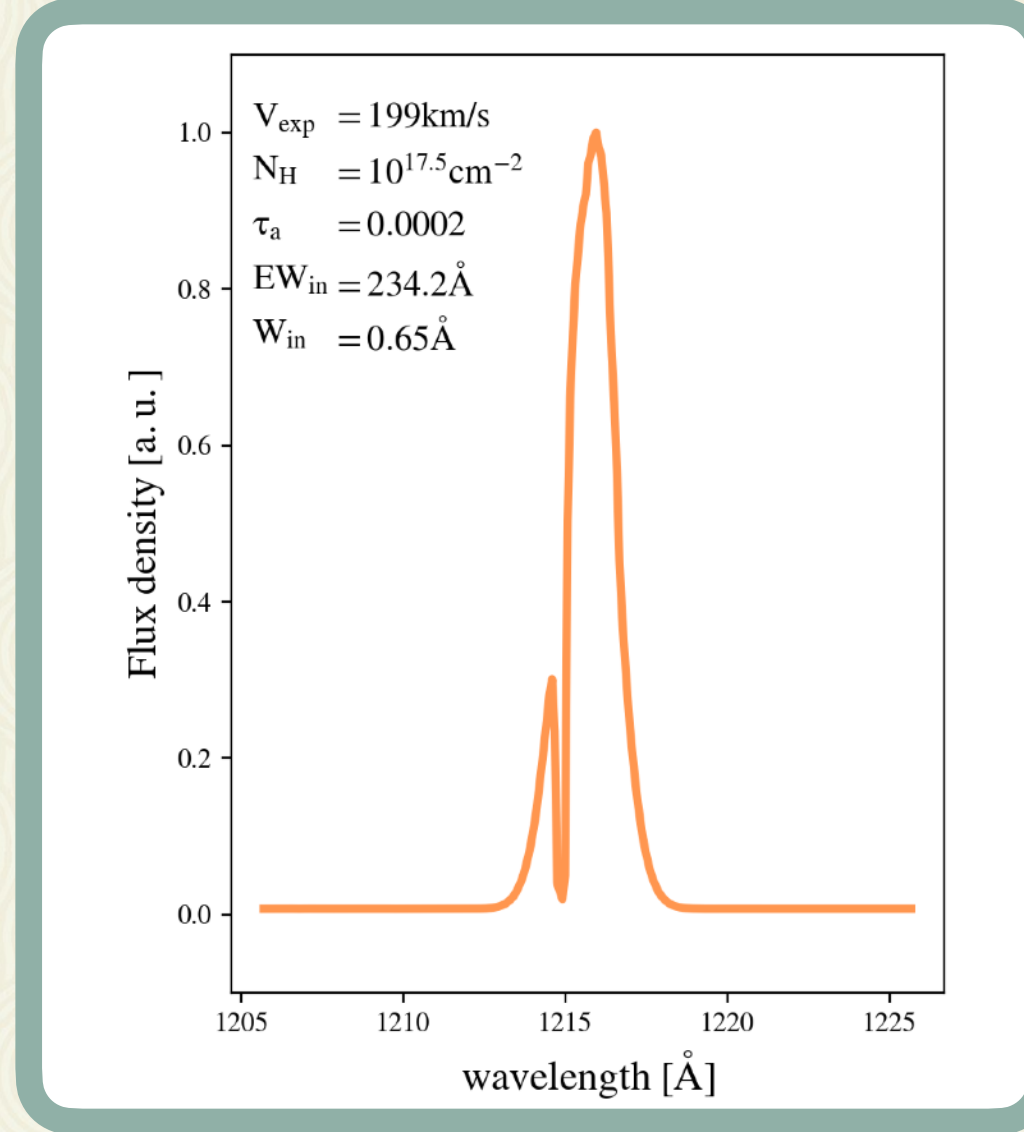
## Monte Carlo Radiative Transfer Code



LyaRT  
Orsi et al. +12

## Ly-alpha line profile

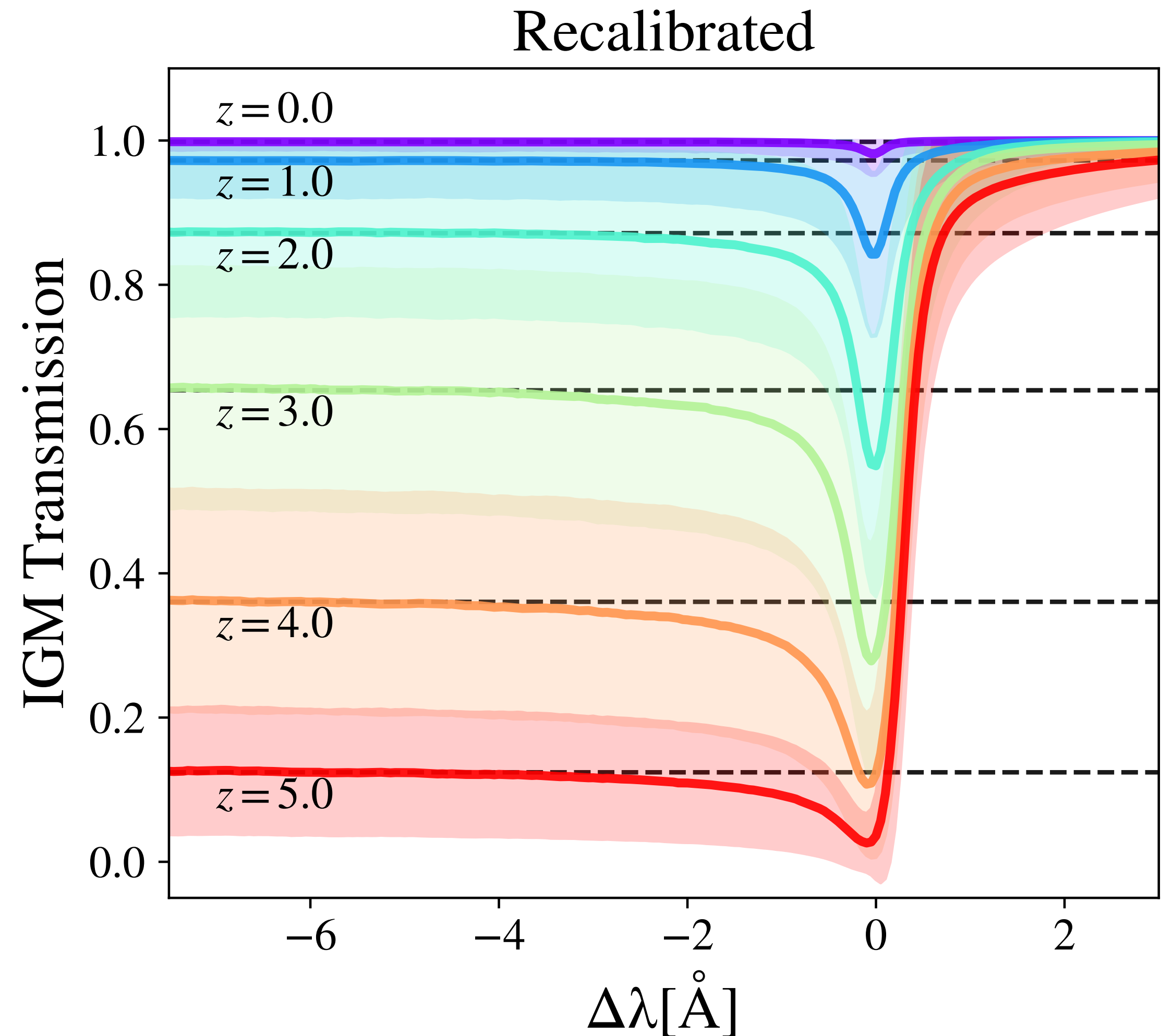




IGM transmission curves from  
C.Byrohl & M. Gronke 2020  
(arXiv:2006.10041):

IllustrisTNG100 simulation : tracing the optical depth in the IGM between the Ly $\alpha$  emitting galaxies and the observer for chosen lines of sight.

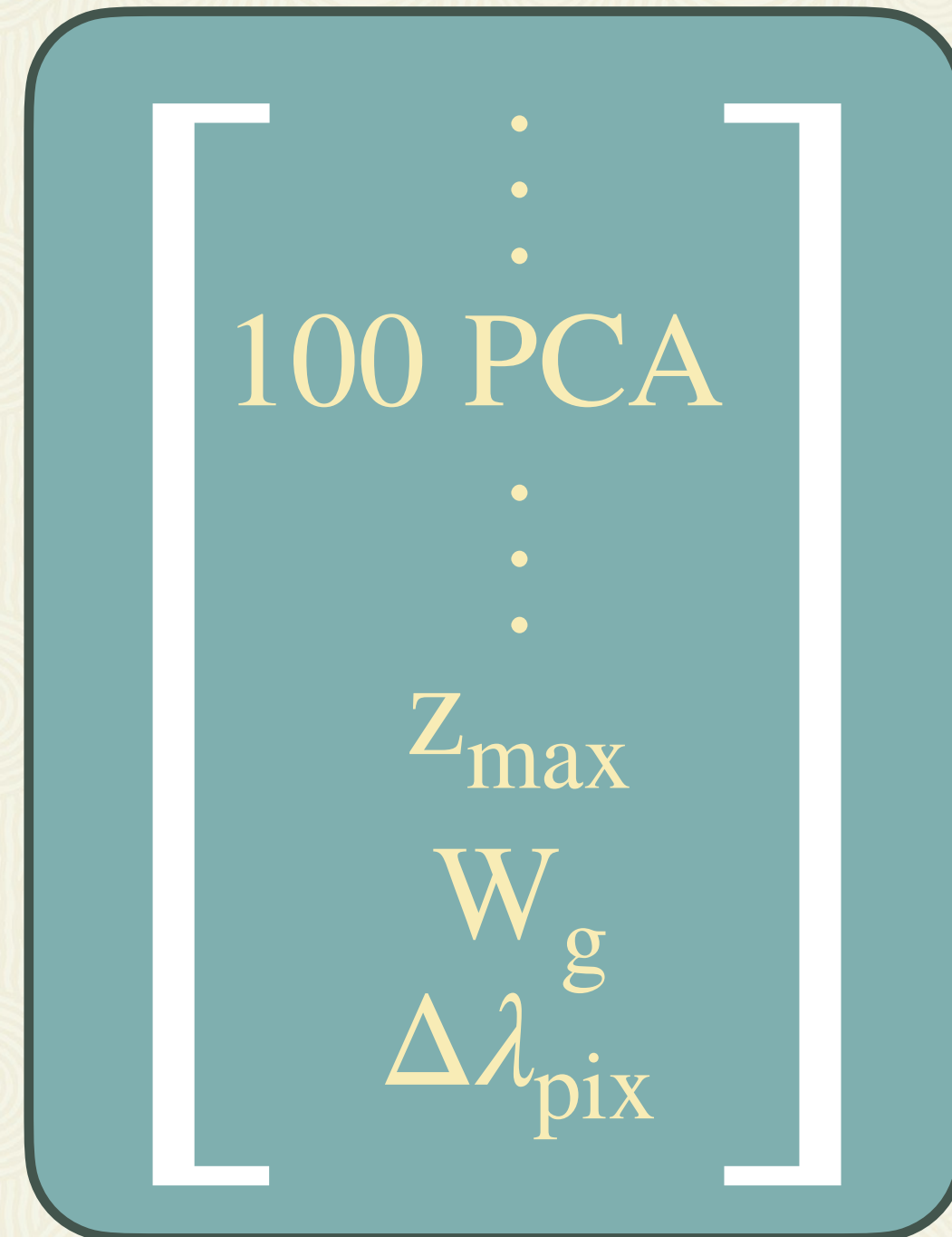
doi:10.5281/zenodo.3832098, <https://doi.org/10.5281/zenodo.3832098>



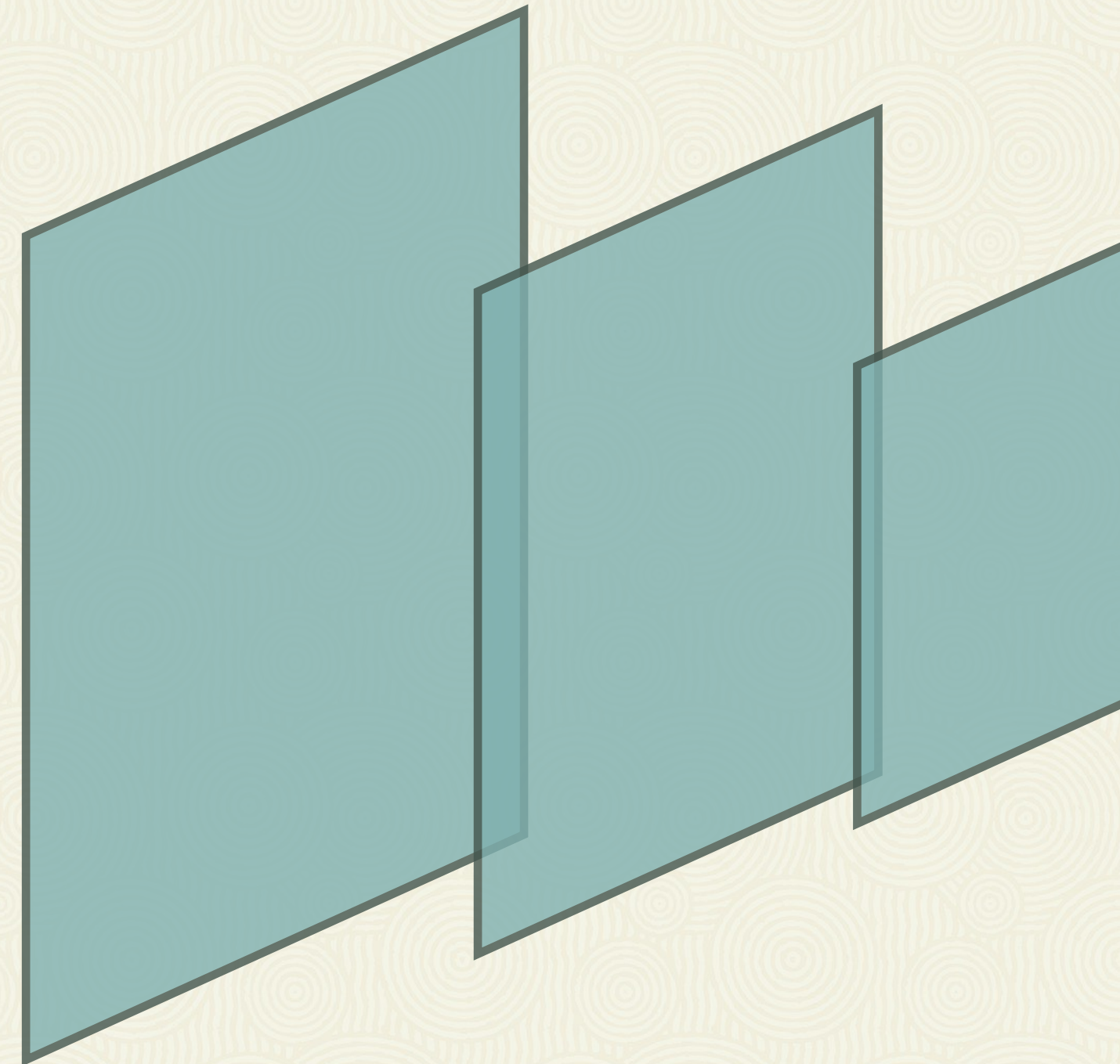
Input

Hidden layers

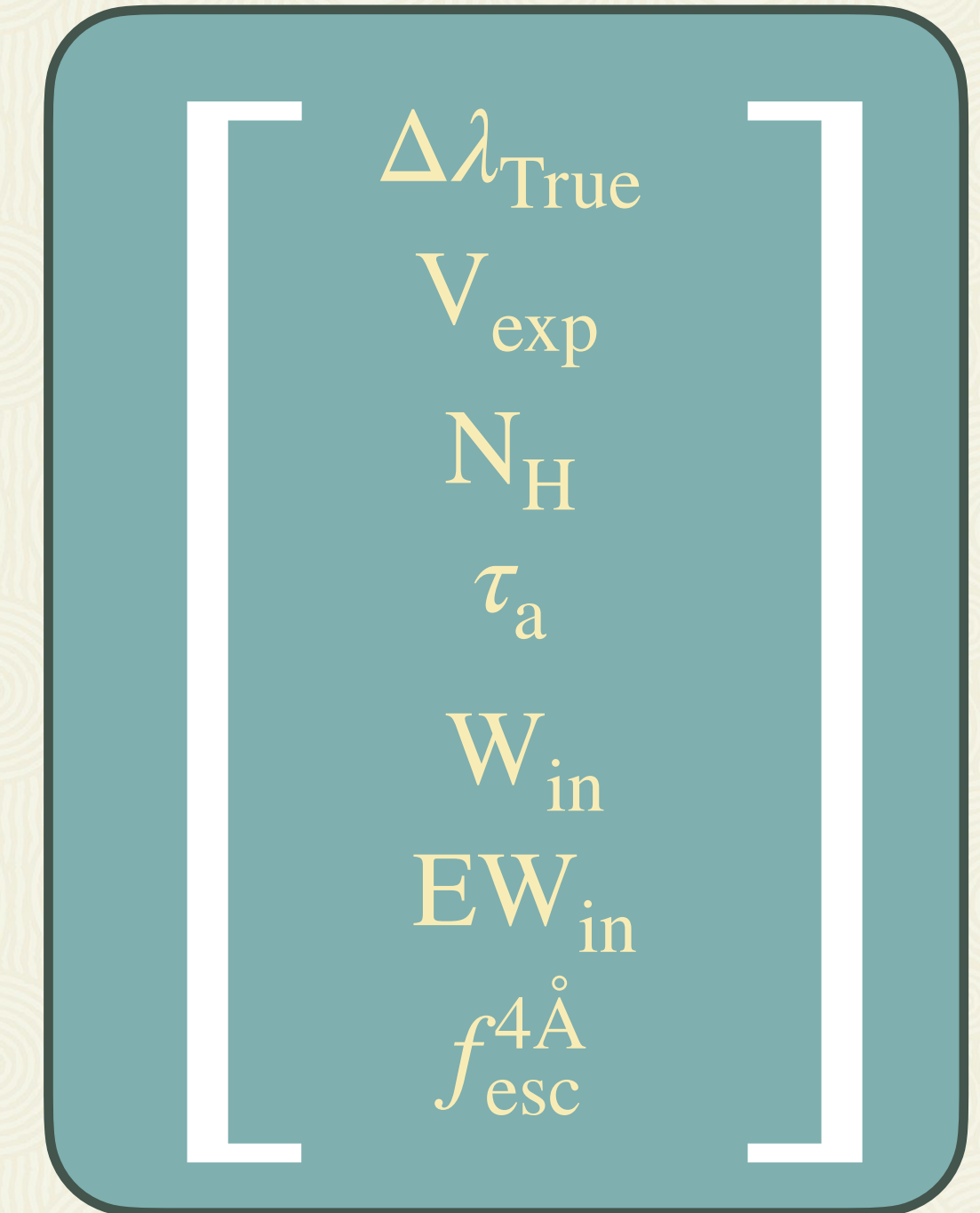
Output



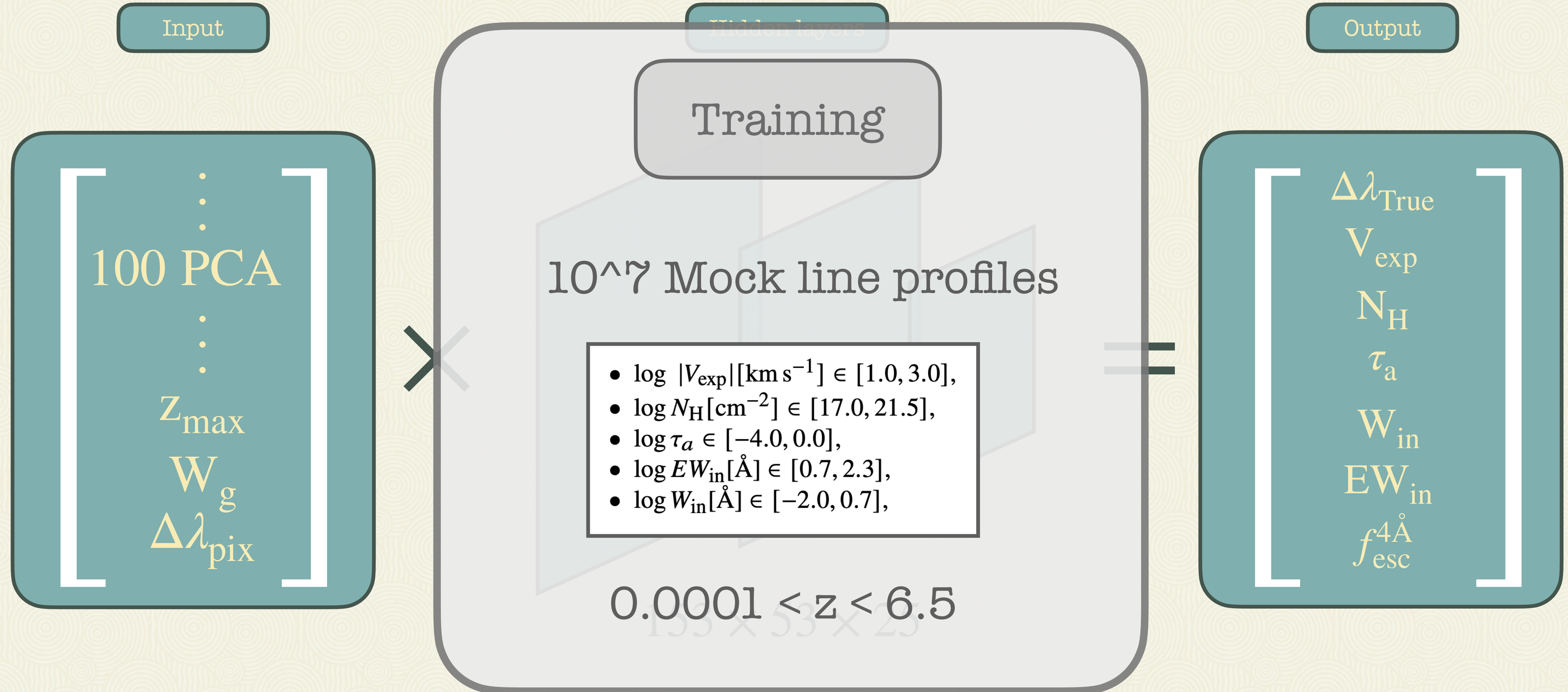
$\times$



$=$



$153 \times 53 \times 25$



## 3 NN models

## IGM+z:

- Input includes redshift.
- Trained with redshift accurate IGM transmission curves.

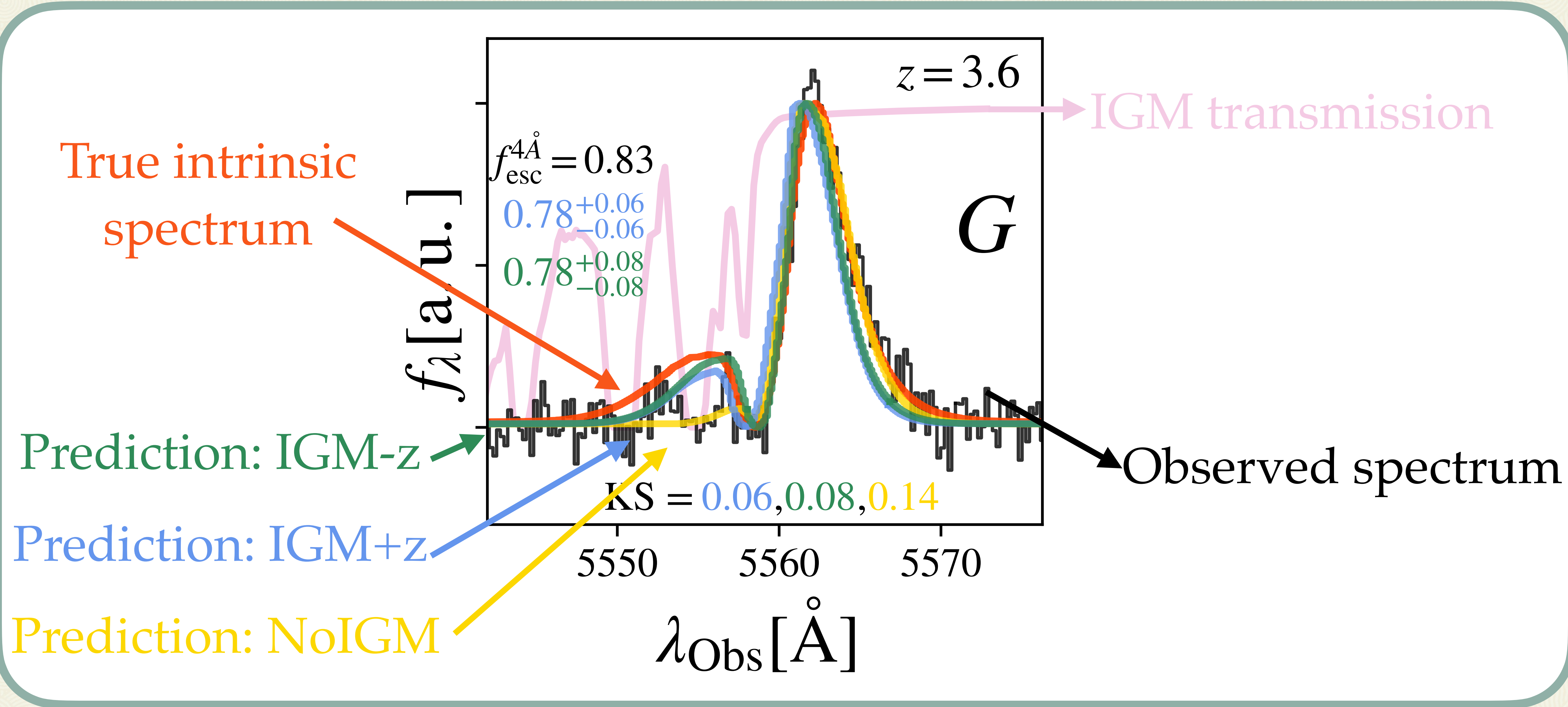
## IGM-z:

- Input does not include redshift.
- Trained with redshift shuffled IGM transmission curves.

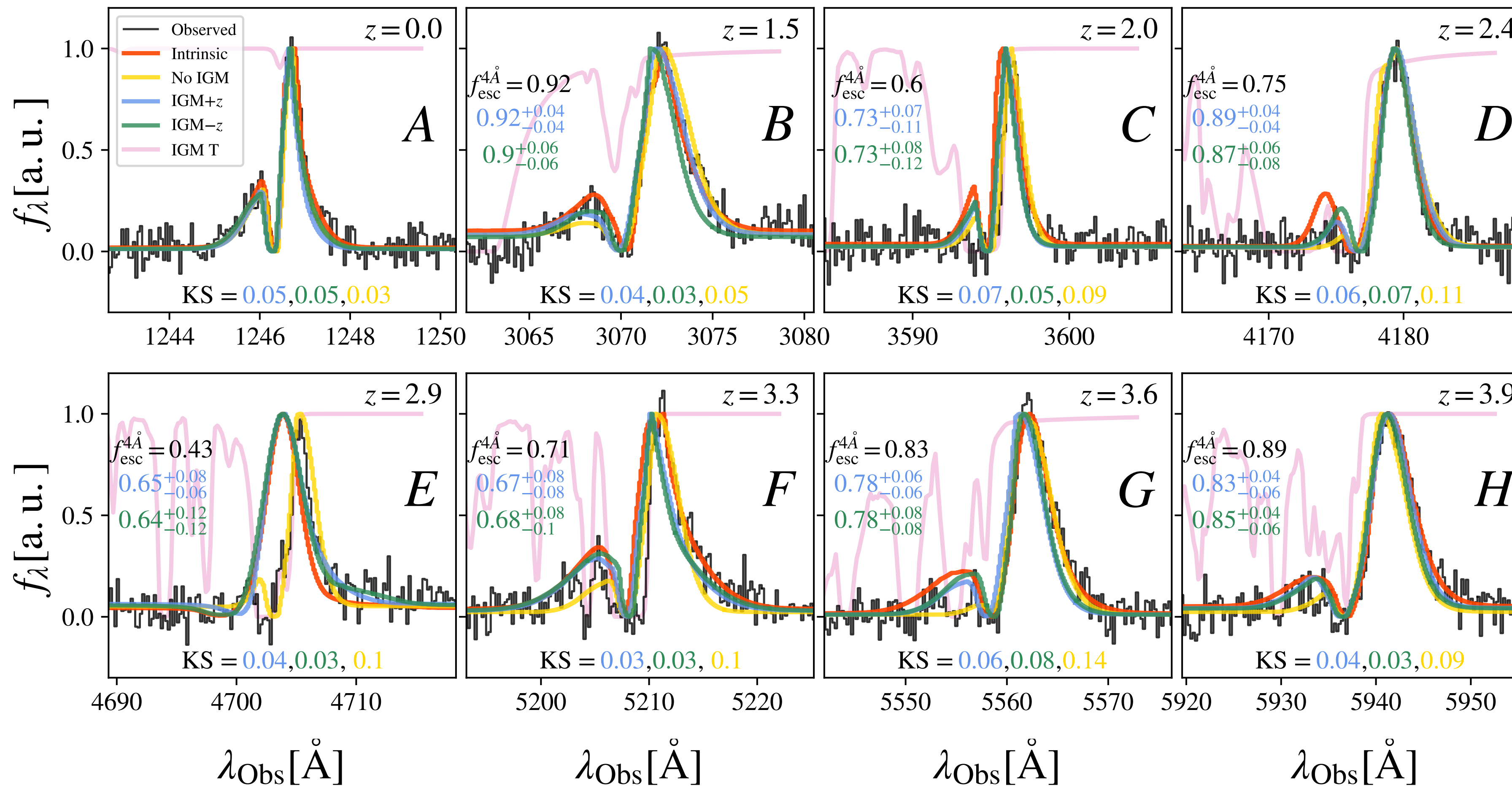
## NoIGM:

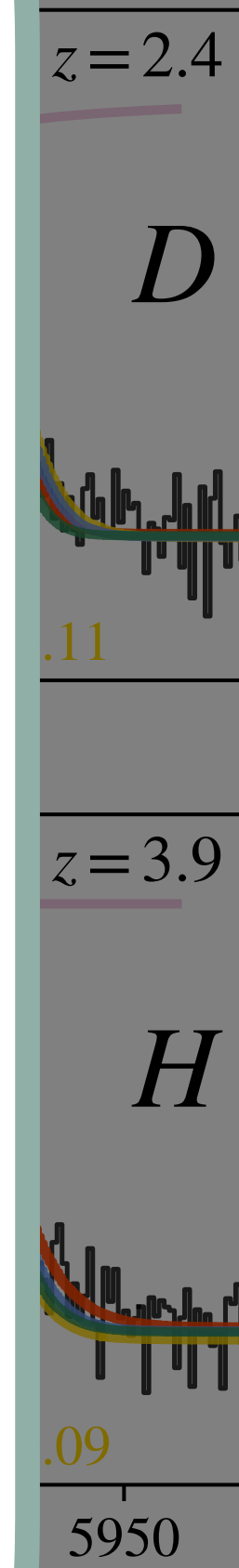
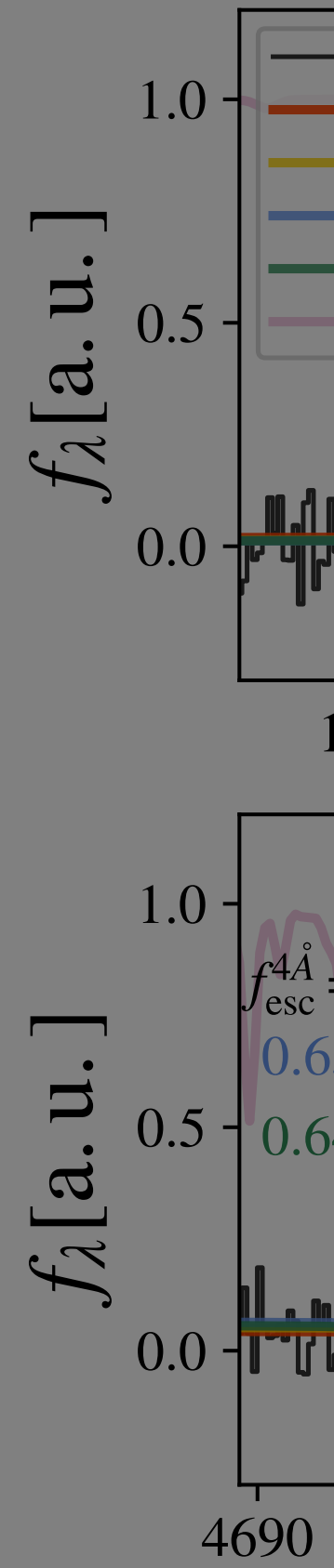
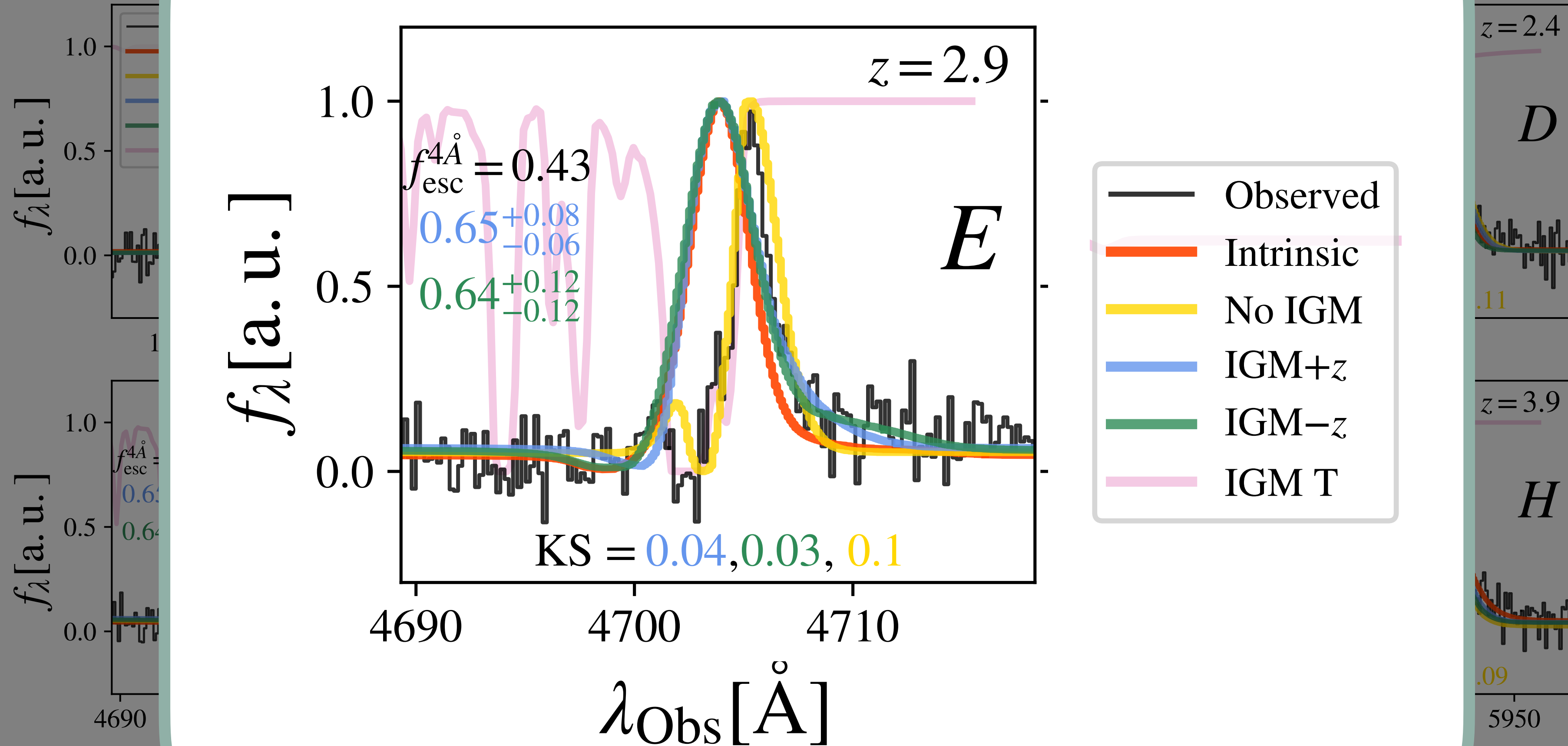
- Input includes redshift.
- Trained without IGM transmission curves.



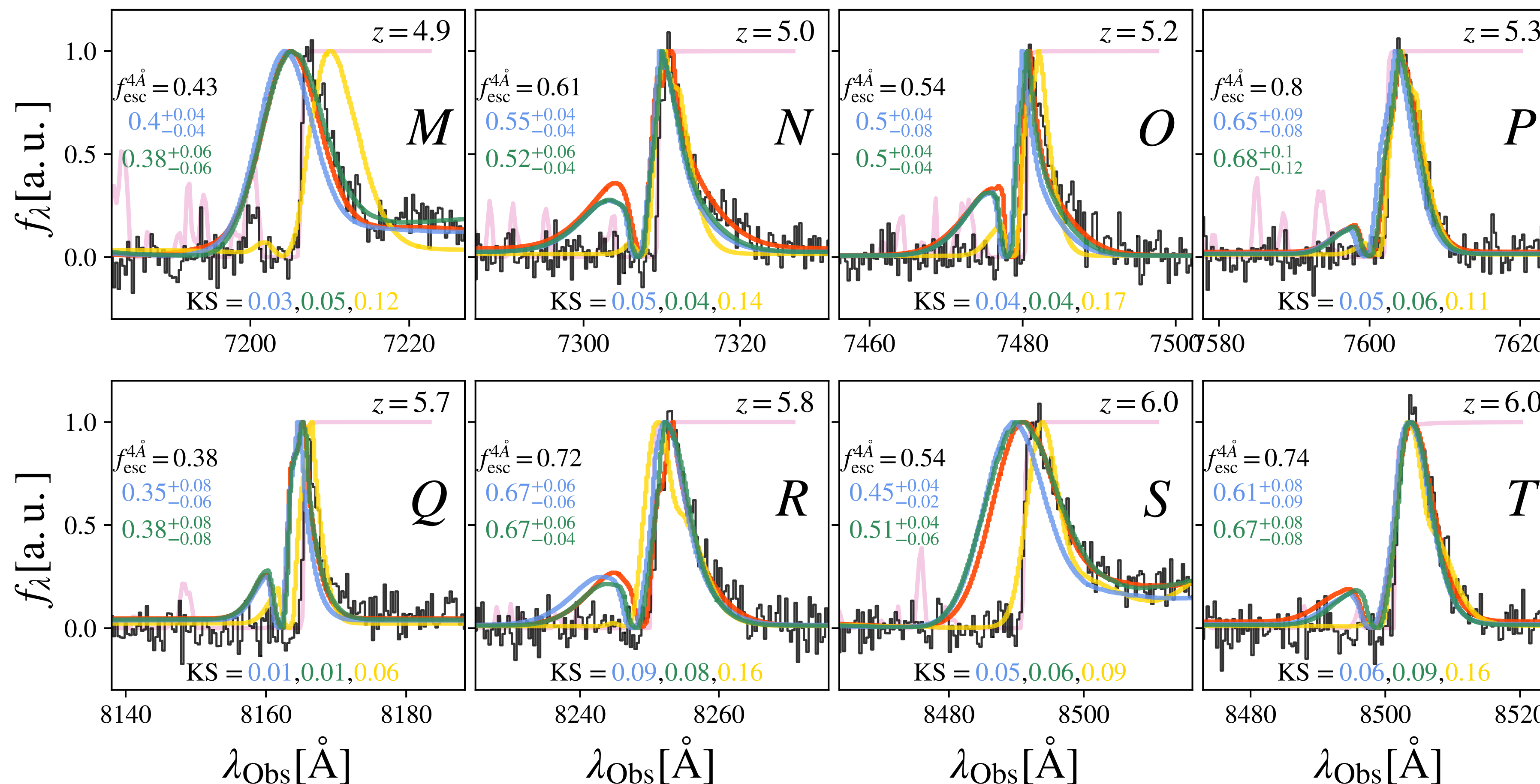


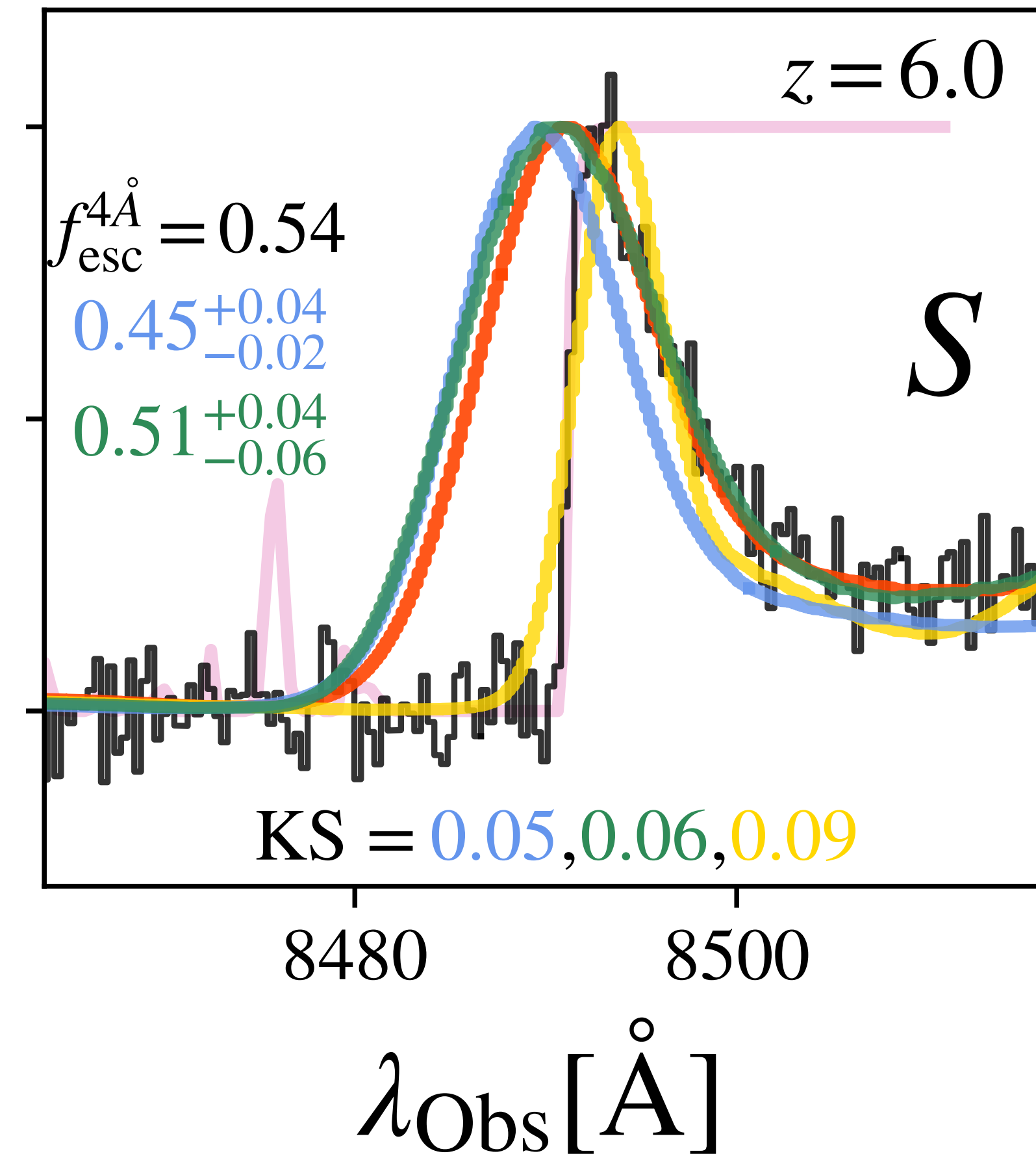
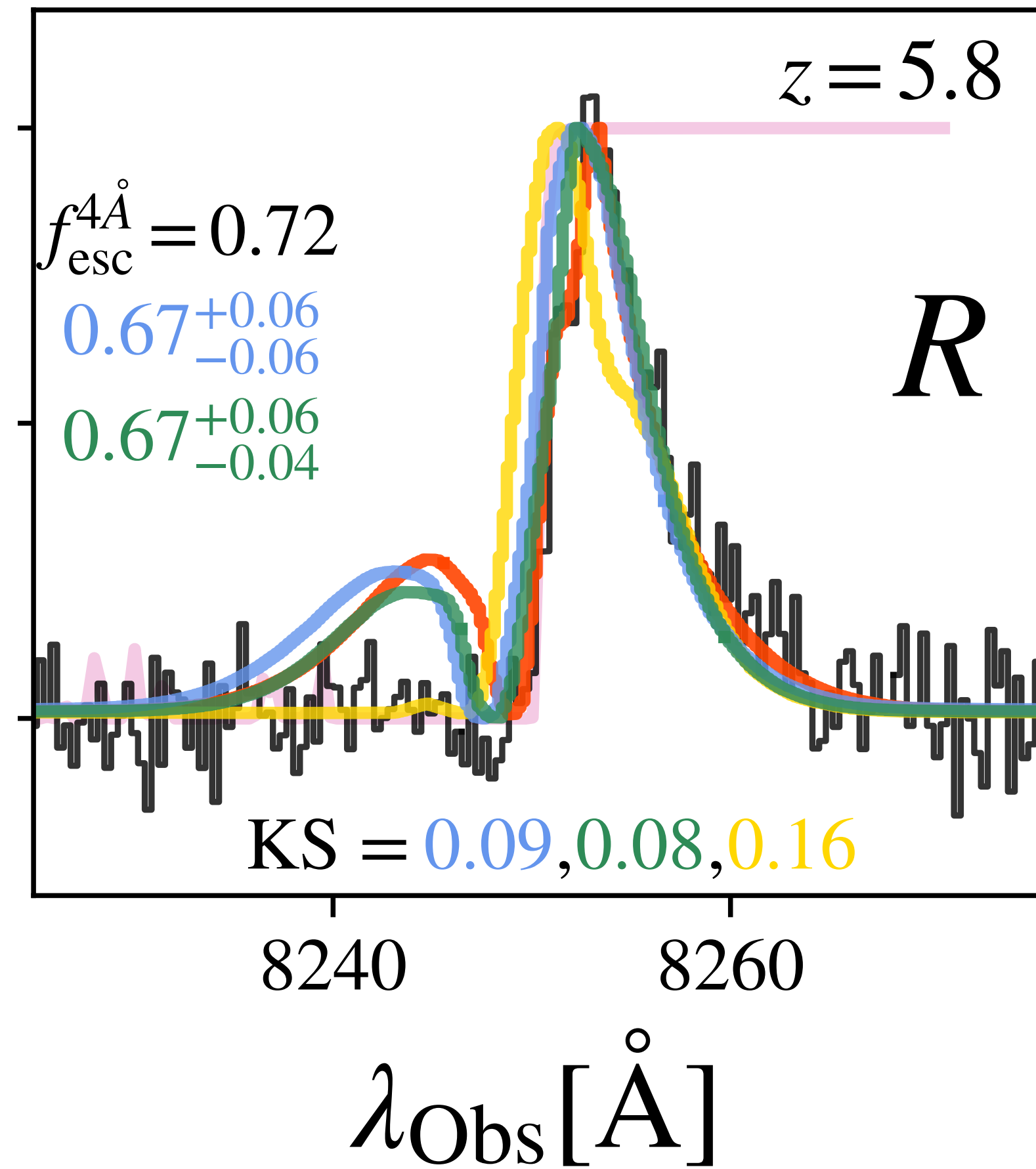
## Successful reconstruction

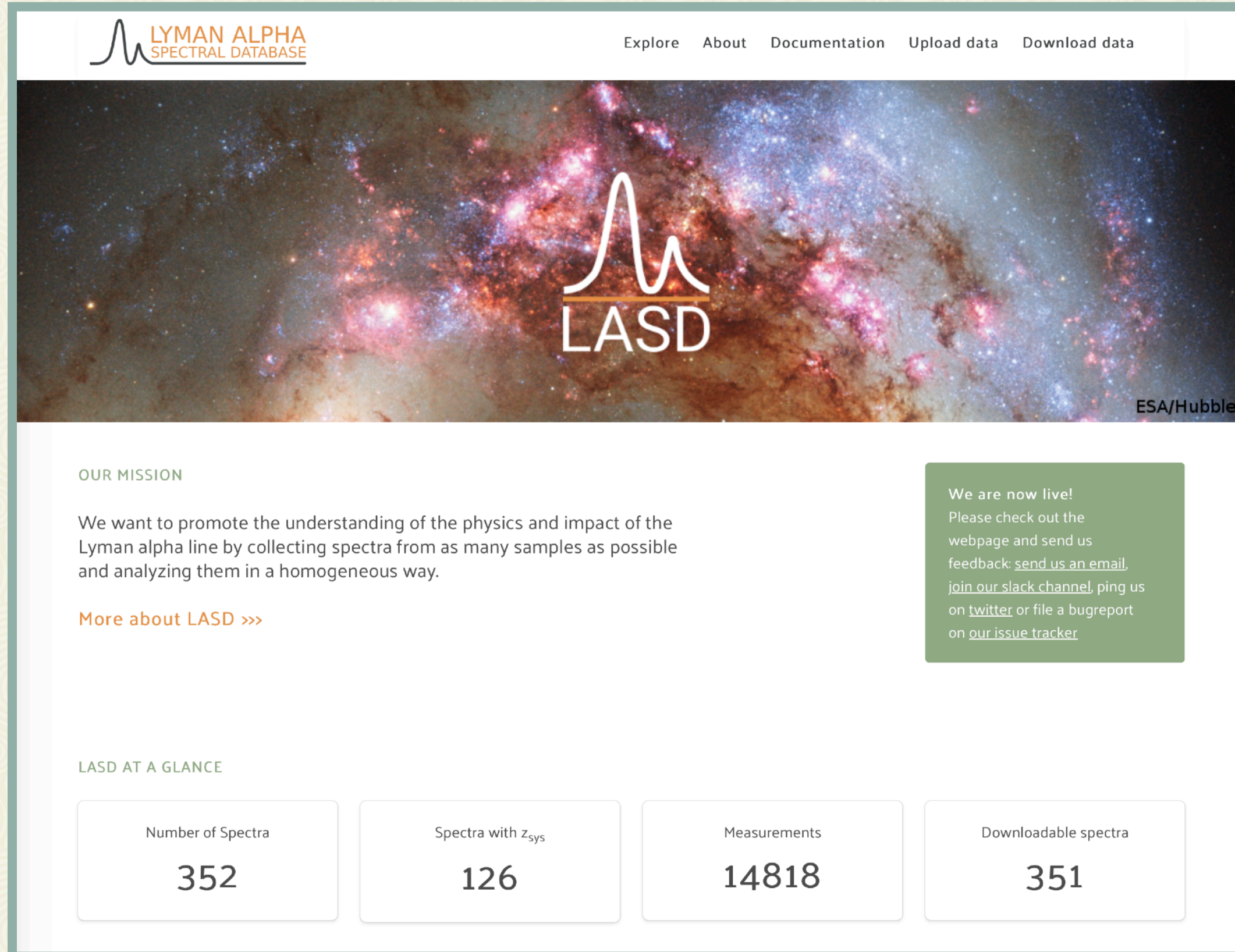




## Successful reconstruction







LYMAN ALPHA SPECTRAL DATABASE

Explore About Documentation Upload data Download data

OUR MISSION

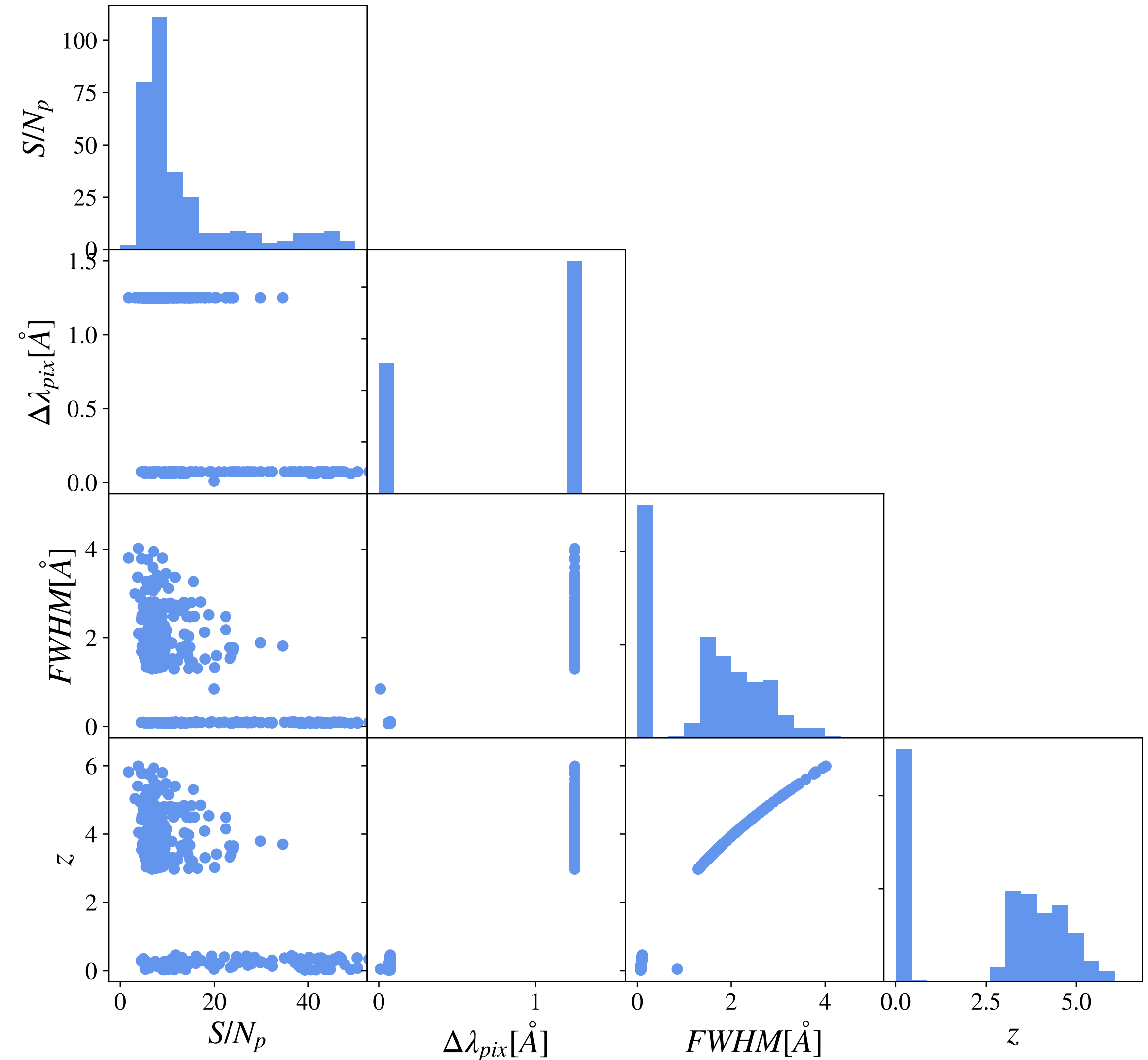
We want to promote the understanding of the physics and impact of the Lyman alpha line by collecting spectra from as many samples as possible and analyzing them in a homogeneous way.

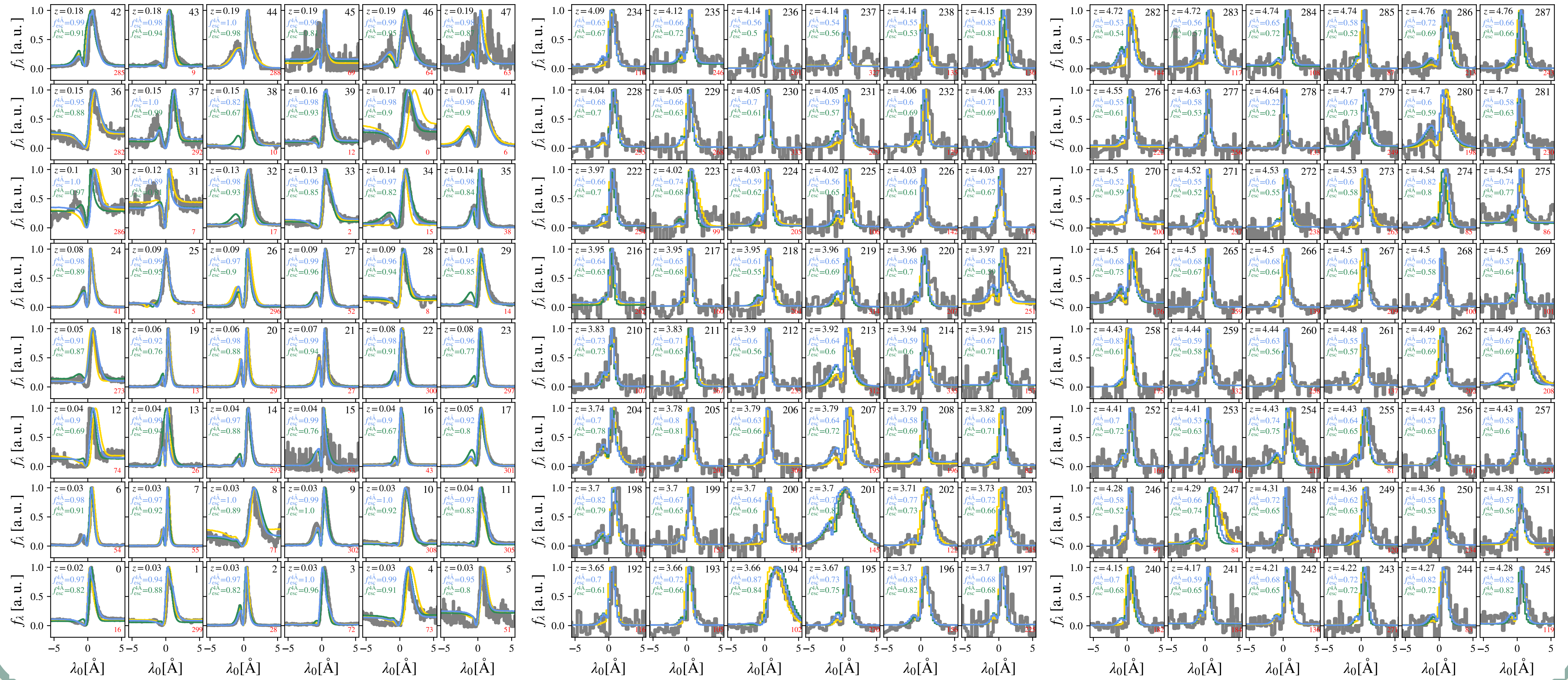
[More about LASD >>>](#)

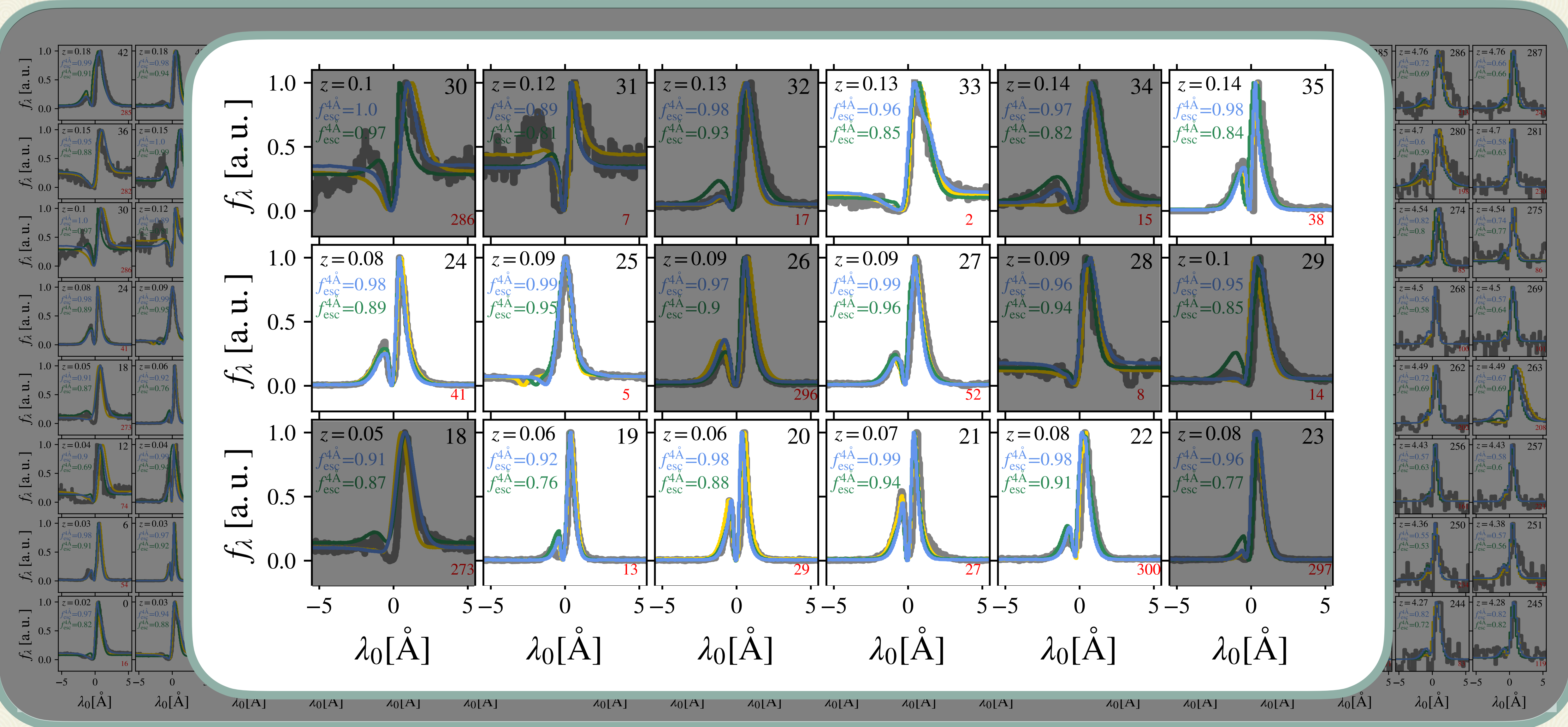
We are now live!  
Please check out the webpage and send us feedback: [send us an email](#), [join our slack channel](#), ping us on [twitter](#) or file a bugreport on [our issue tracker](#).

LASD AT A GLANCE

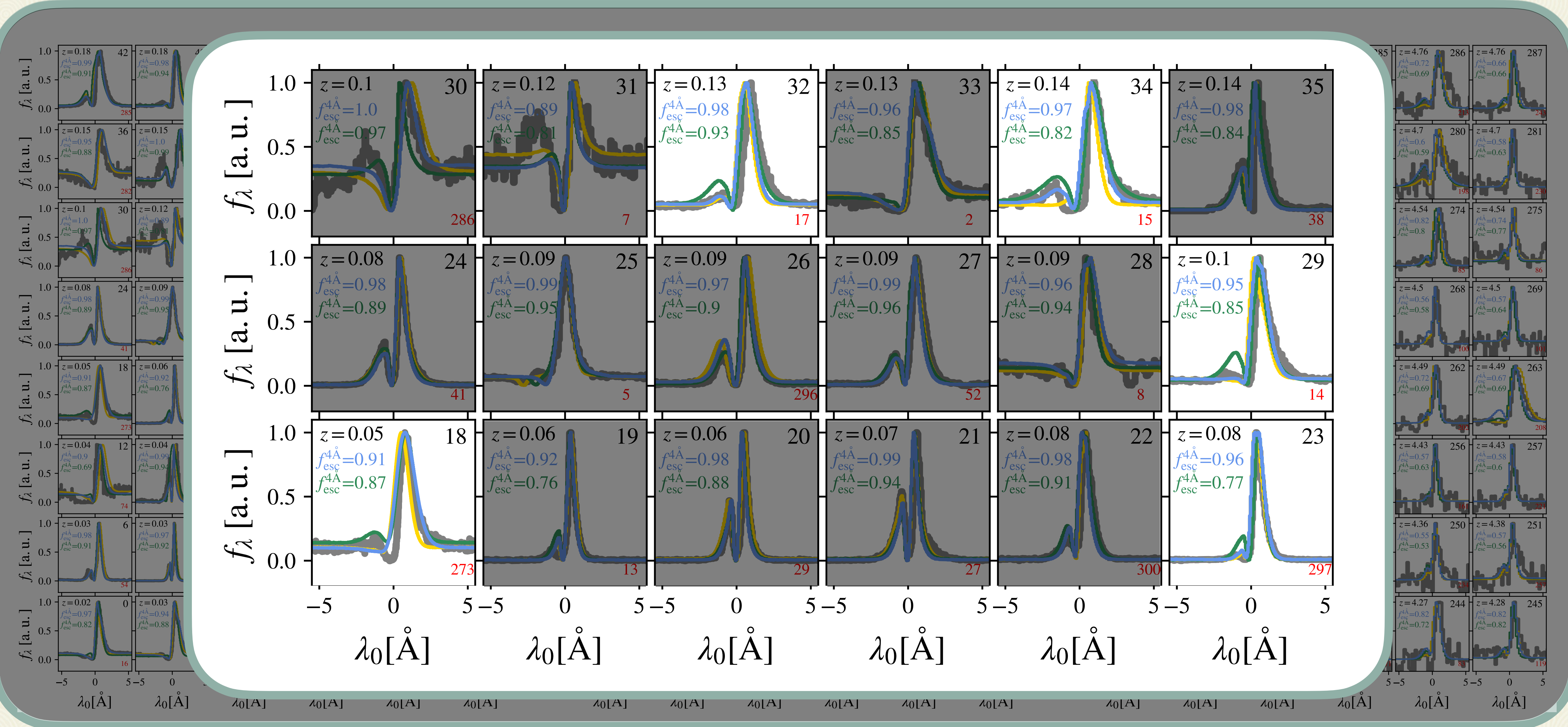
Number of Spectra	Spectra with $z_{sys}$	Measurements	Downloadable spectra
352	126	14818	351

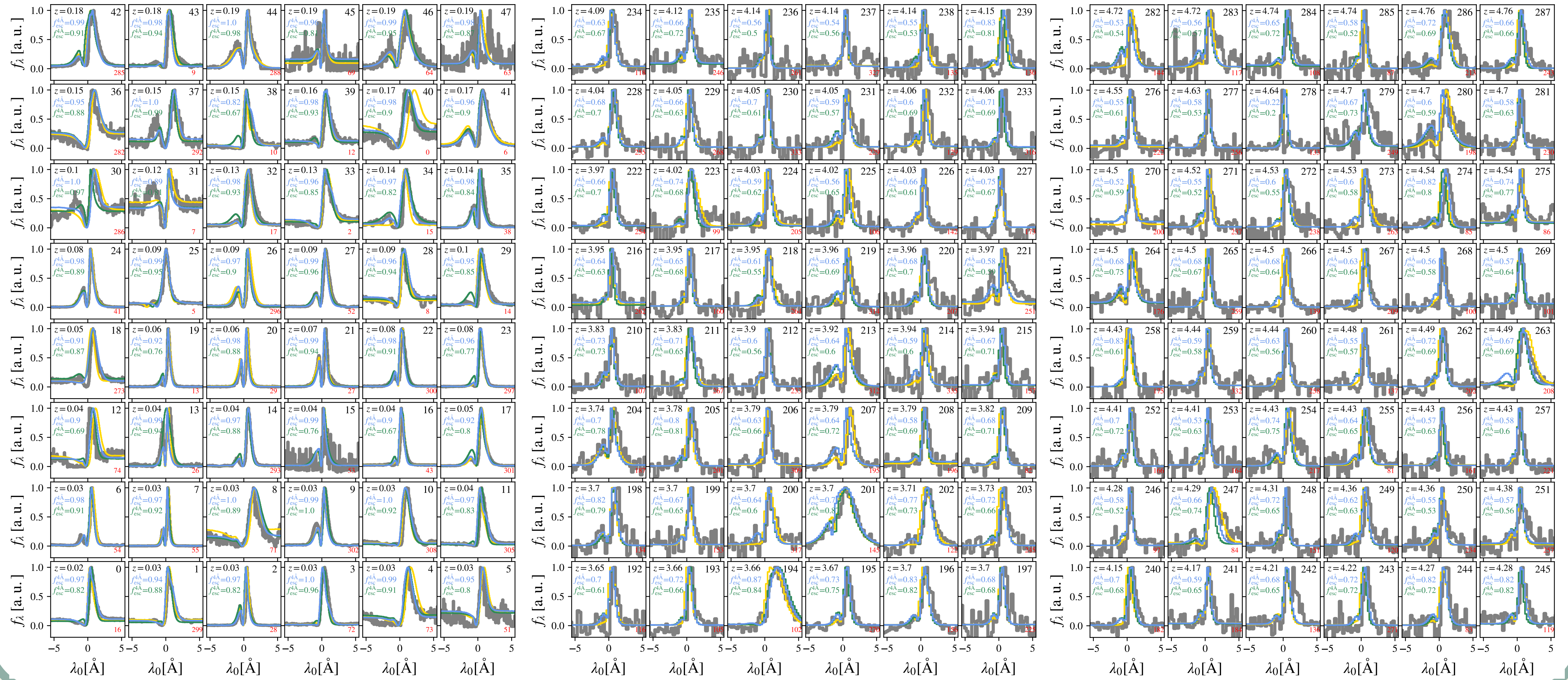








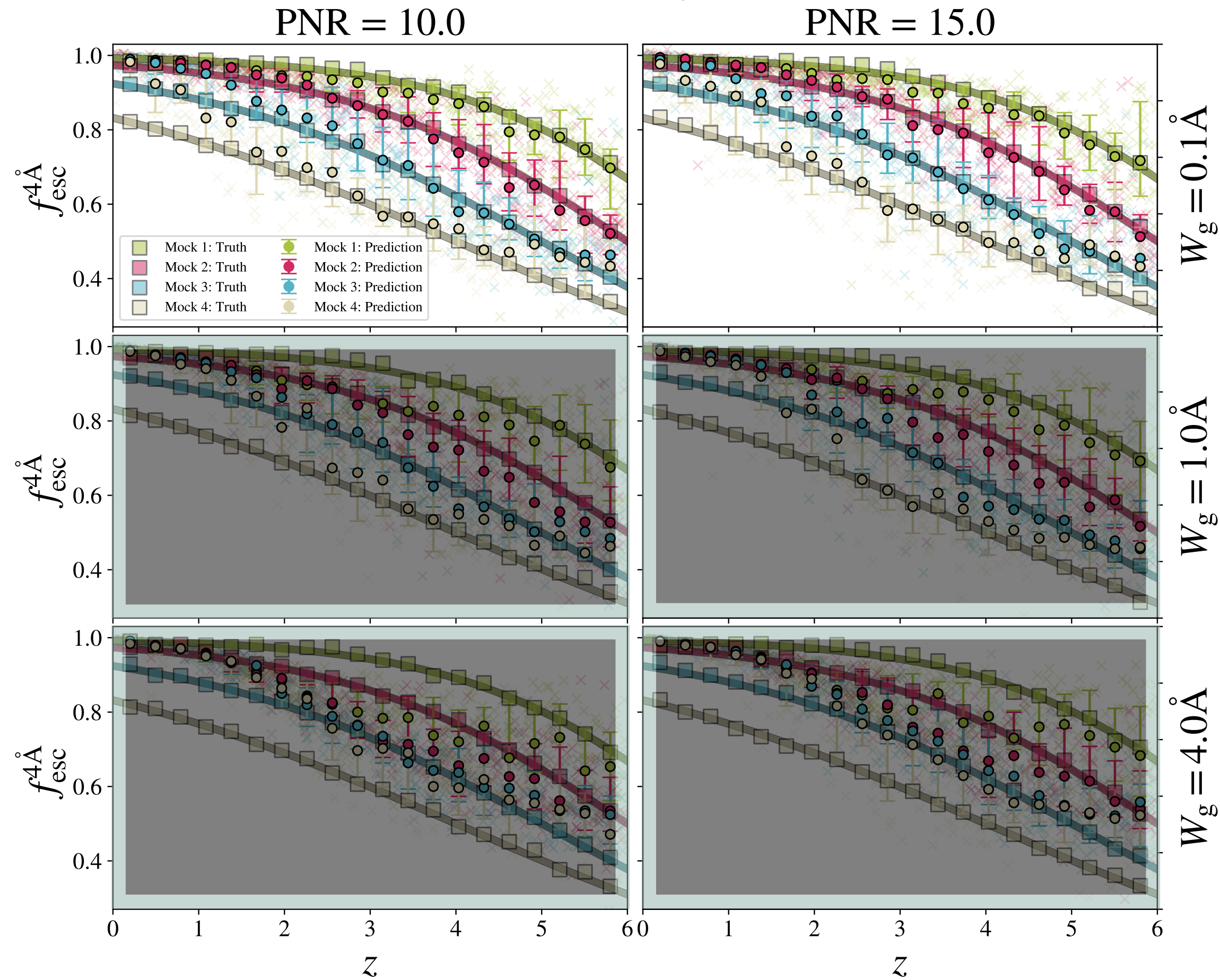




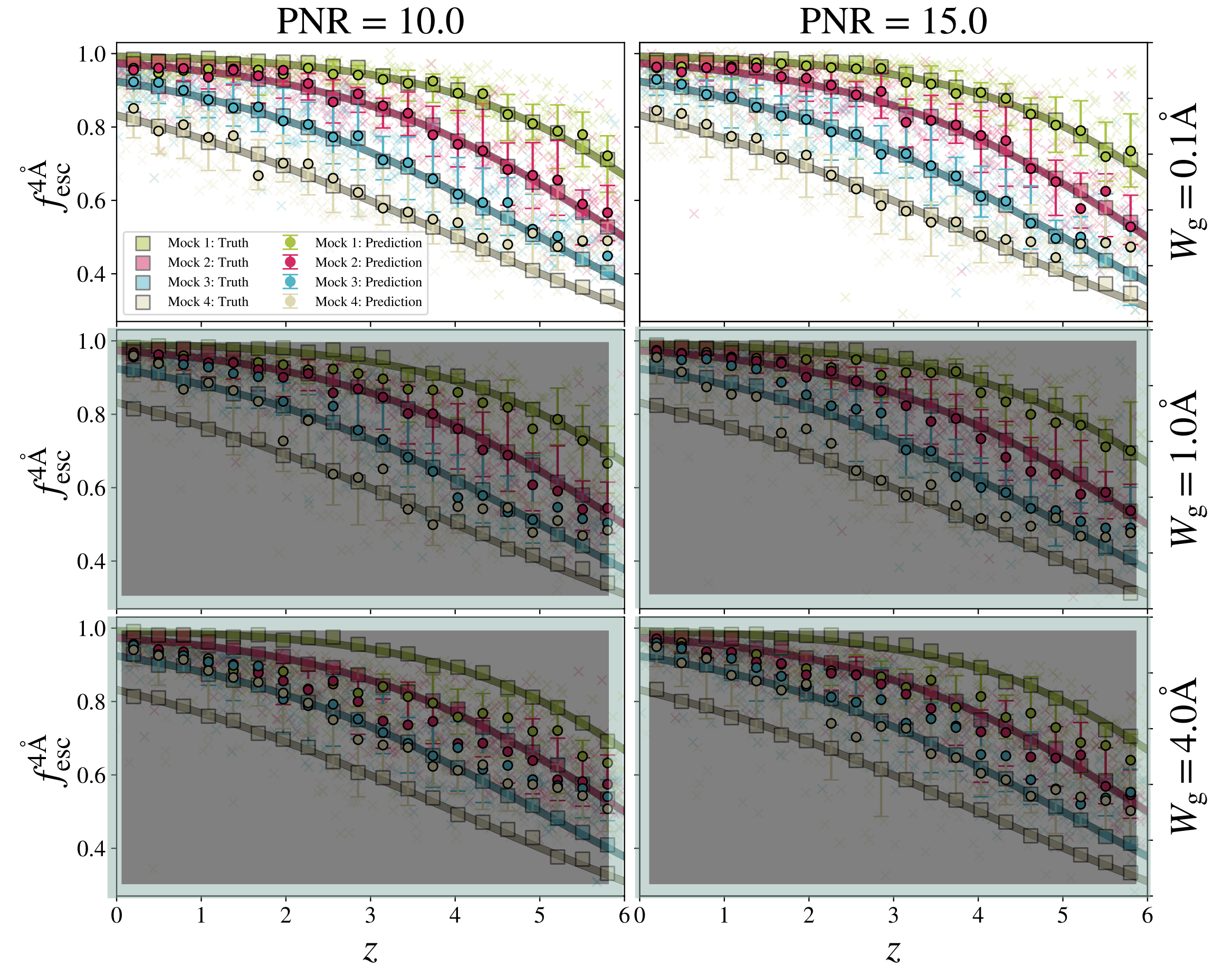
$z \in [0.0, 7.0]$

	$\log V_{\text{exp}}$	$\log N_{\text{H}}$	$\log \tau_{\text{a}}$	$\log EW_{\text{in}}$	$\log W_{\text{in}}$	$z$	$f_{\text{esc}}^{4\text{\AA}}$
$\log V_{\text{exp}}$		$0.07^{+0.01}_{-0.03}$	$0.13^{+0.04}_{-0.04}$	$0.01^{+0.03}_{-0.06}$	$0.15^{+0.03}_{-0.03}$	$-0.2^{+0.03}_{-0.02}$	$0.11^{+0.03}_{-0.0}$
$\log N_{\text{H}}$	$0.2^{+0.03}_{-0.02}$		$-0.07^{+0.02}_{-0.04}$	$-0.0^{+0.05}_{-0.03}$	$0.23^{+0.03}_{-0.05}$	$-0.19^{+0.01}_{-0.02}$	$0.32^{+0.02}_{-0.02}$
$\log \tau_{\text{a}}$	$0.12^{+0.03}_{-0.03}$	$-0.02^{+0.03}_{-0.03}$		$-0.28^{+0.03}_{-0.02}$	$-0.01^{+0.05}_{-0.03}$	$-0.22^{+0.03}_{-0.02}$	$0.03^{+0.02}_{-0.02}$
$\log EW_{\text{in}}$	$0.01^{+0.01}_{-0.04}$	$-0.05^{+0.04}_{-0.01}$	$-0.4^{+0.03}_{-0.05}$		$0.18^{+0.02}_{-0.02}$	$0.14^{+0.03}_{-0.02}$	$-0.03^{+0.03}_{-0.02}$
$\log W_{\text{in}}$	$-0.06^{+0.01}_{-0.06}$	$0.15^{+0.01}_{-0.03}$	$-0.16^{+0.01}_{-0.02}$	$0.23^{+0.04}_{-0.03}$		$-0.26^{+0.02}_{-0.01}$	$0.18^{+0.03}_{-0.03}$
$z$	$-0.1^{+0.01}_{-0.02}$	$-0.15^{+0.03}_{-0.01}$	$-0.35^{+0.04}_{-0.03}$	$0.13^{+0.02}_{-0.01}$	$-0.02^{+0.01}_{-0.03}$		$-0.66^{+0.02}_{-0.01}$
$f_{\text{esc}}^{4\text{\AA}}$	$0.16^{+0.02}_{-0.03}$	$0.24^{+0.04}_{-0.02}$	$0.14^{+0.03}_{-0.02}$	$0.07^{+0.02}_{-0.03}$	$0.14^{+0.03}_{-0.02}$	$-0.8^{+0.02}_{-0.02}$	

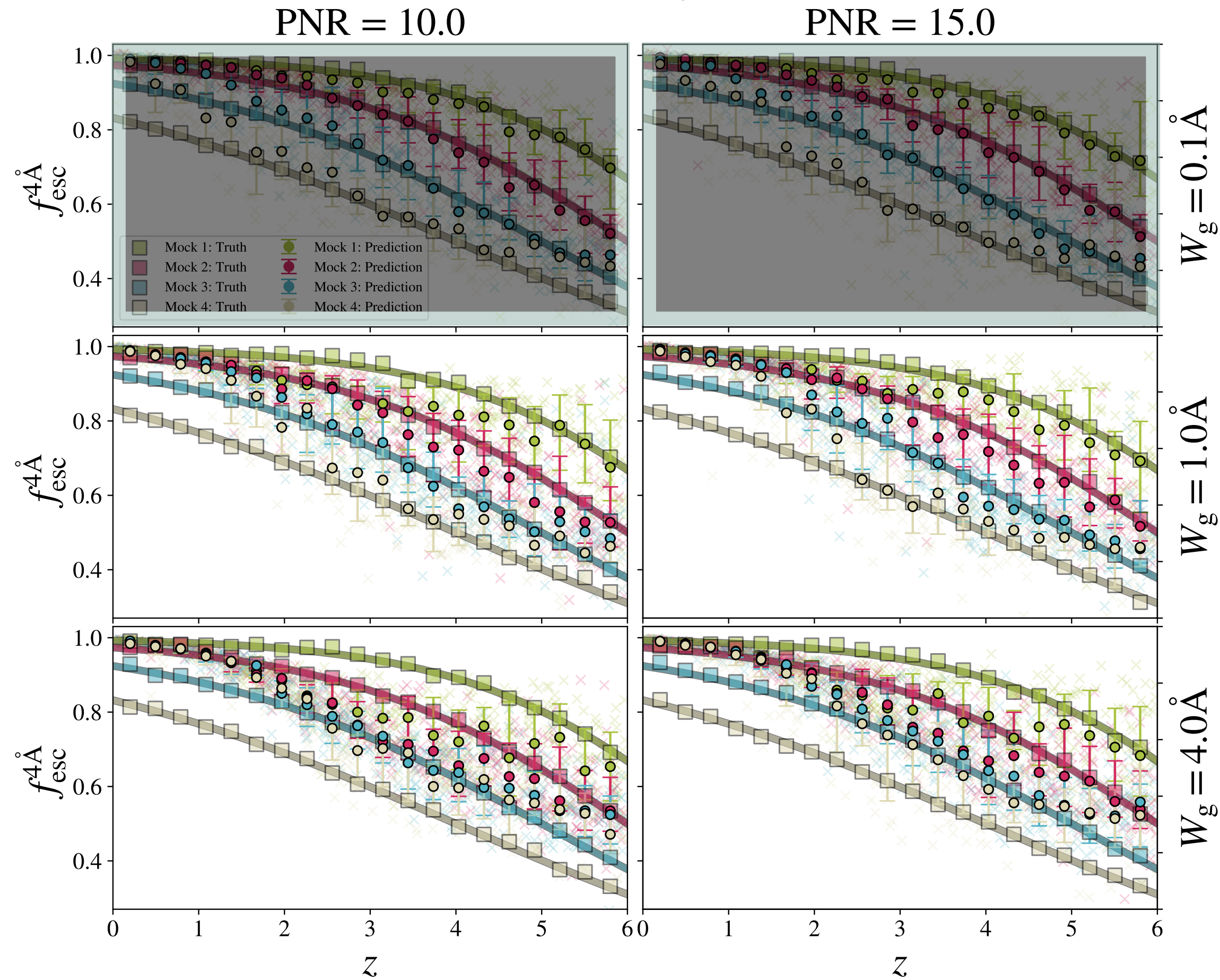
## IGM+z



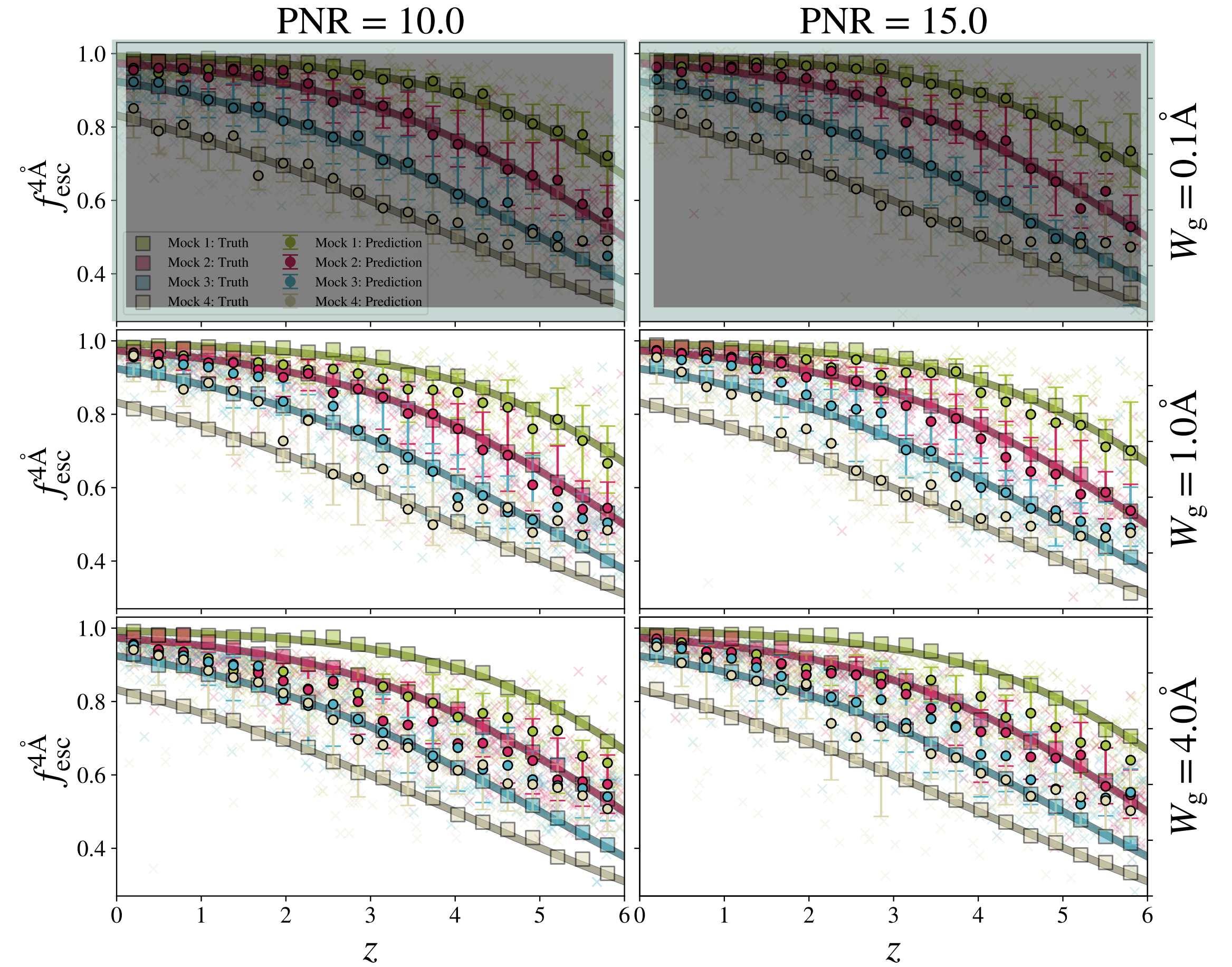
## IGM-z

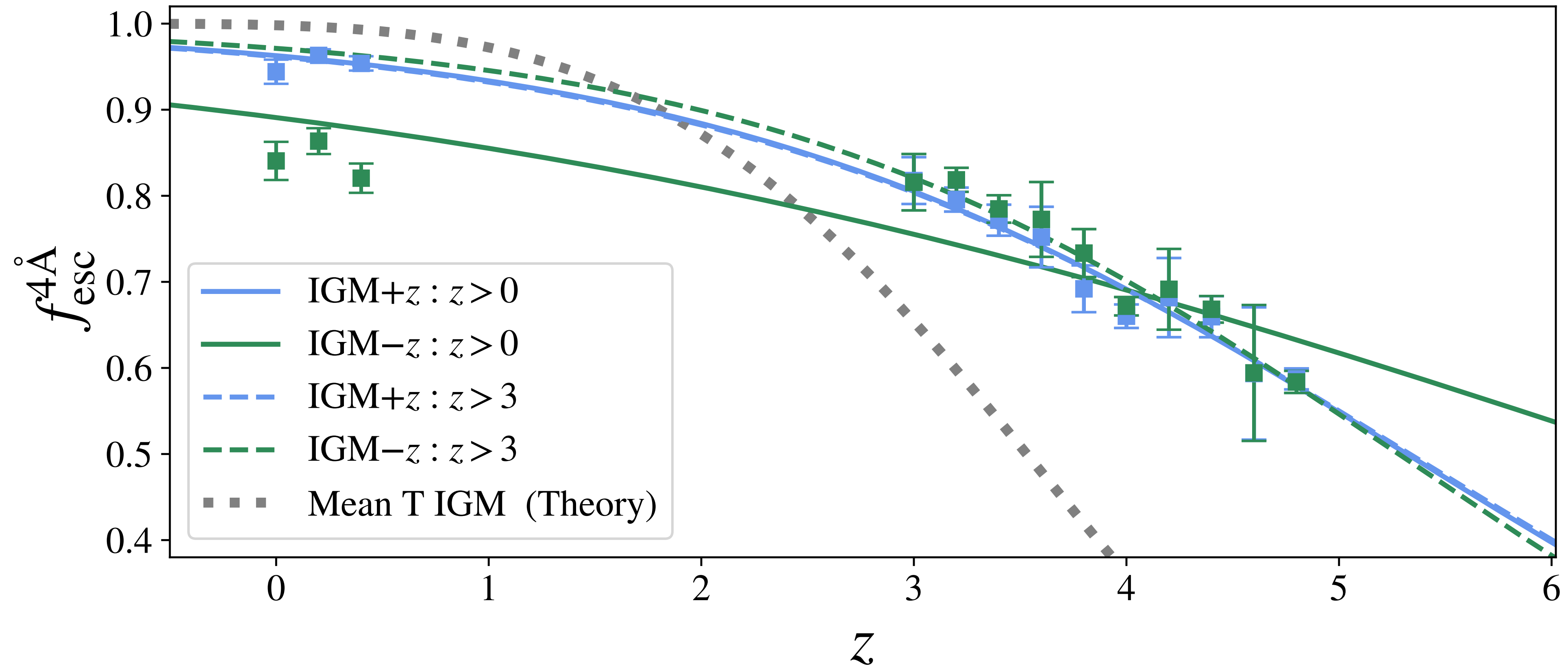


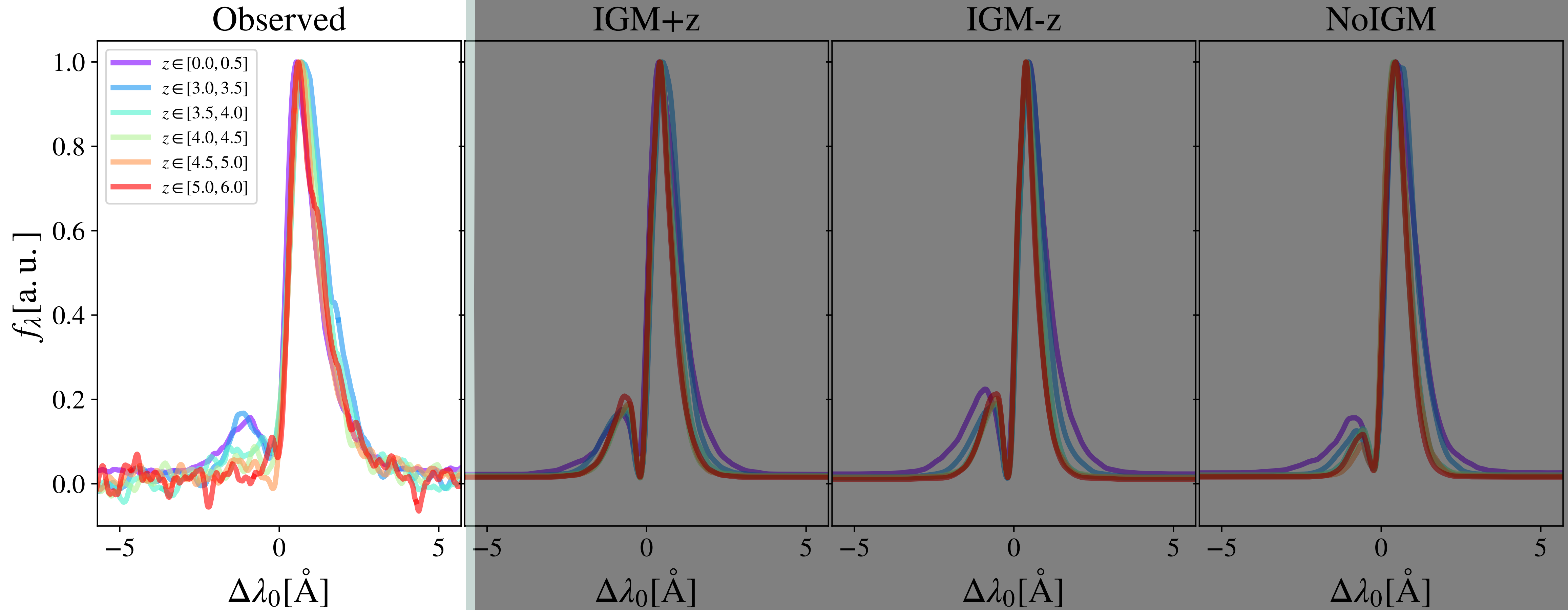
## IGM+ $z$

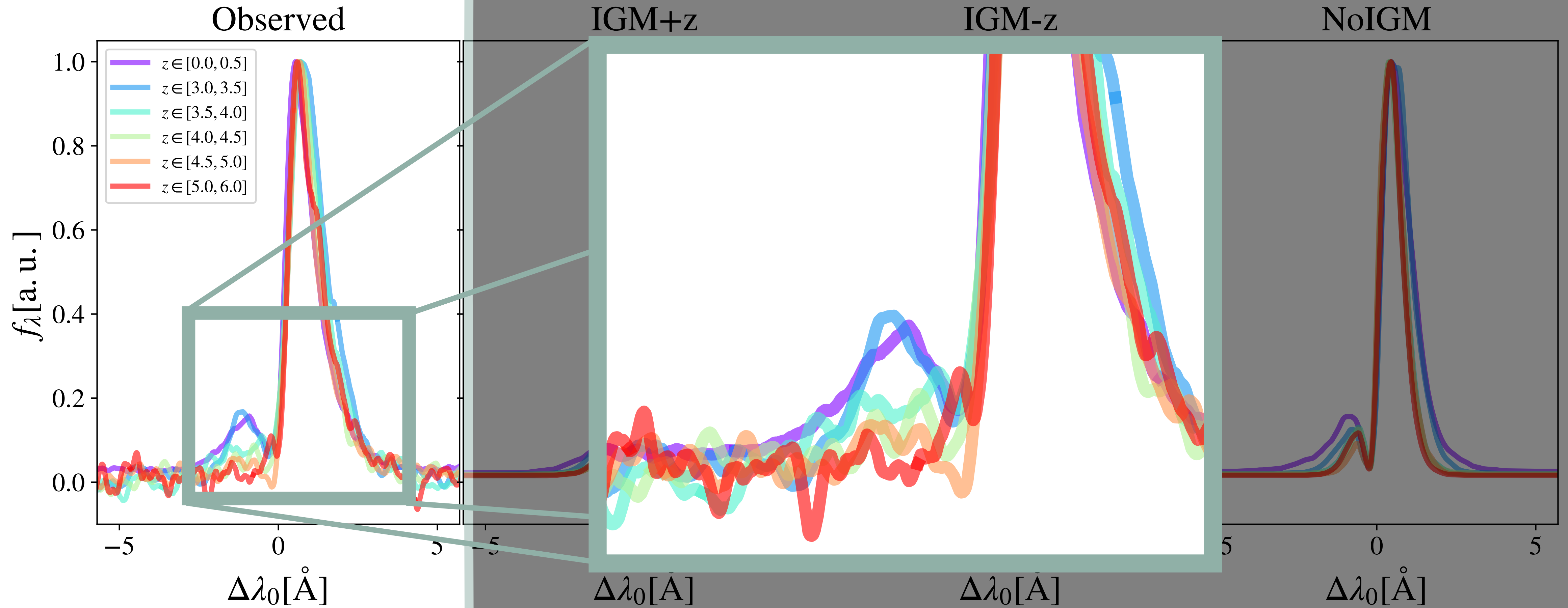


## IGM- $z$

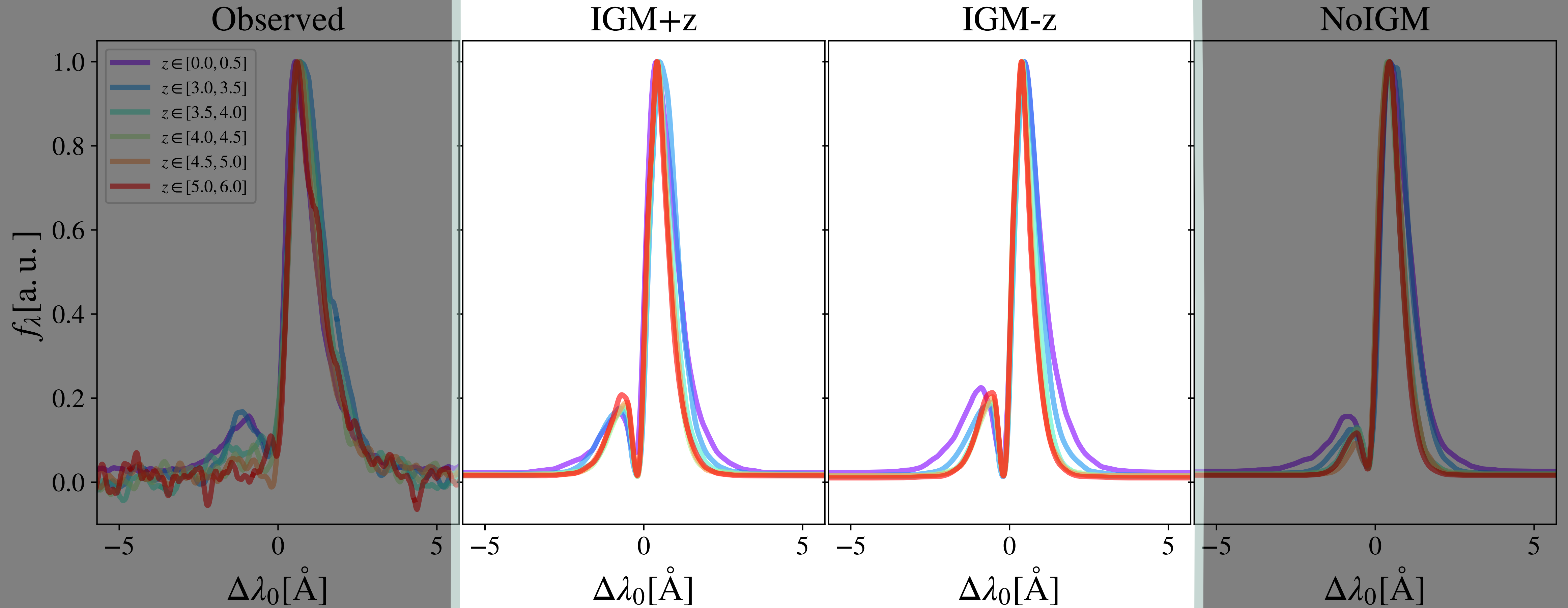


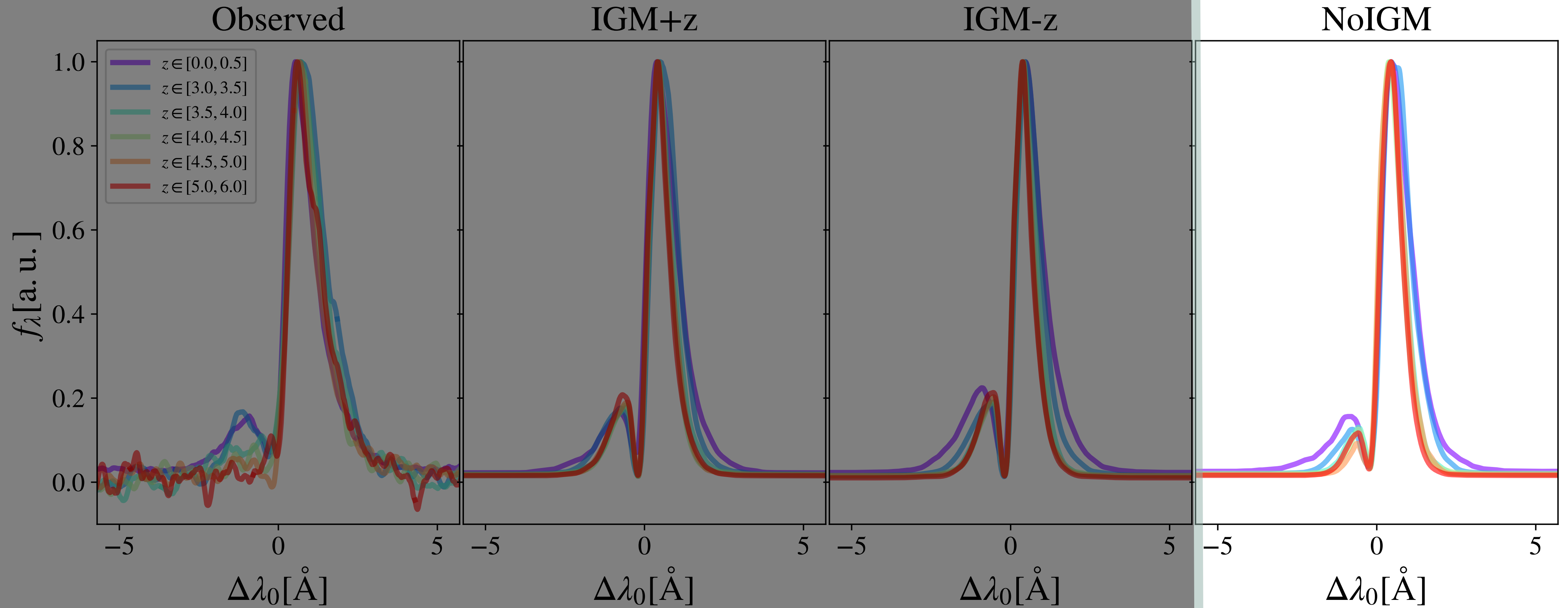


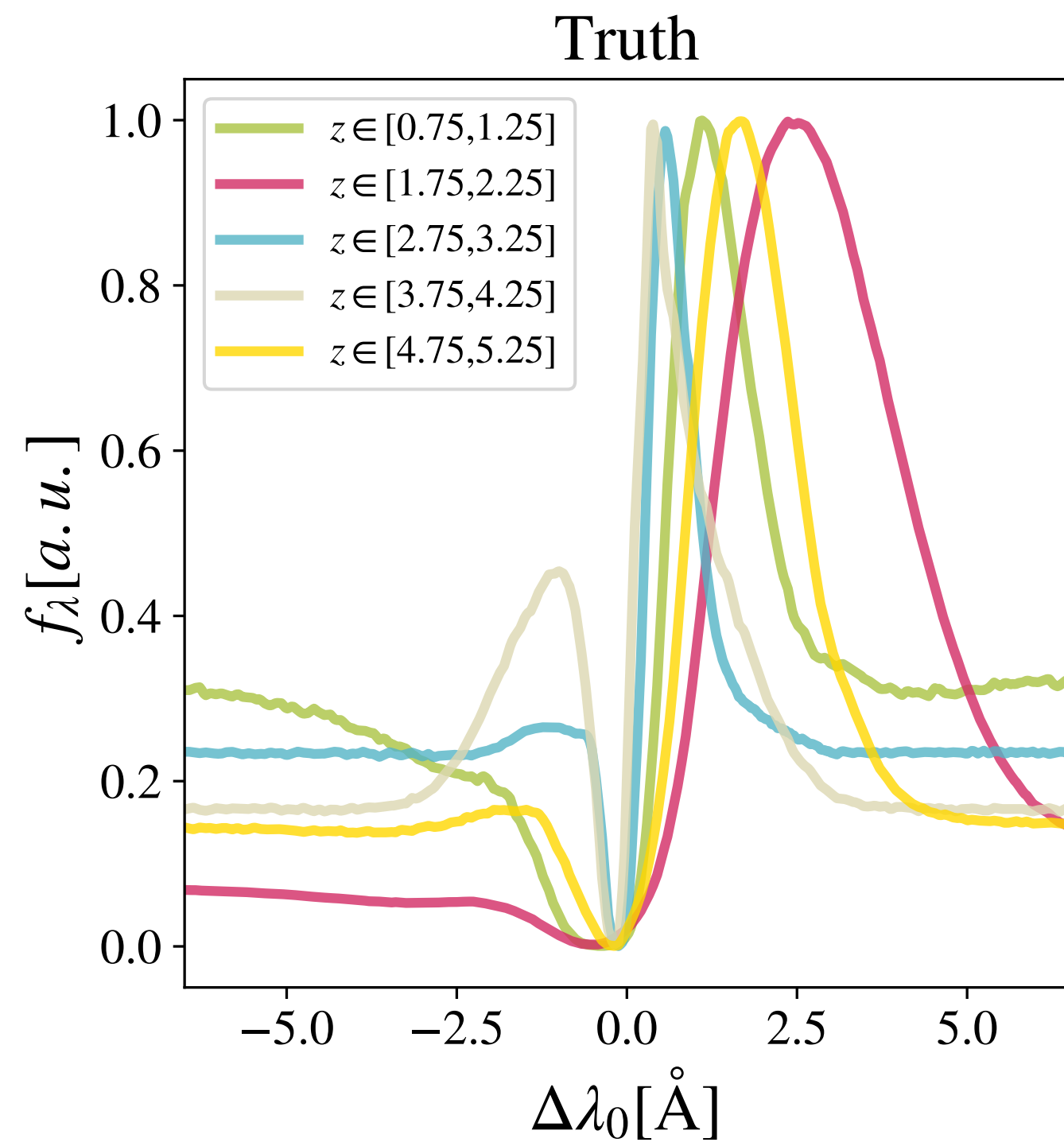


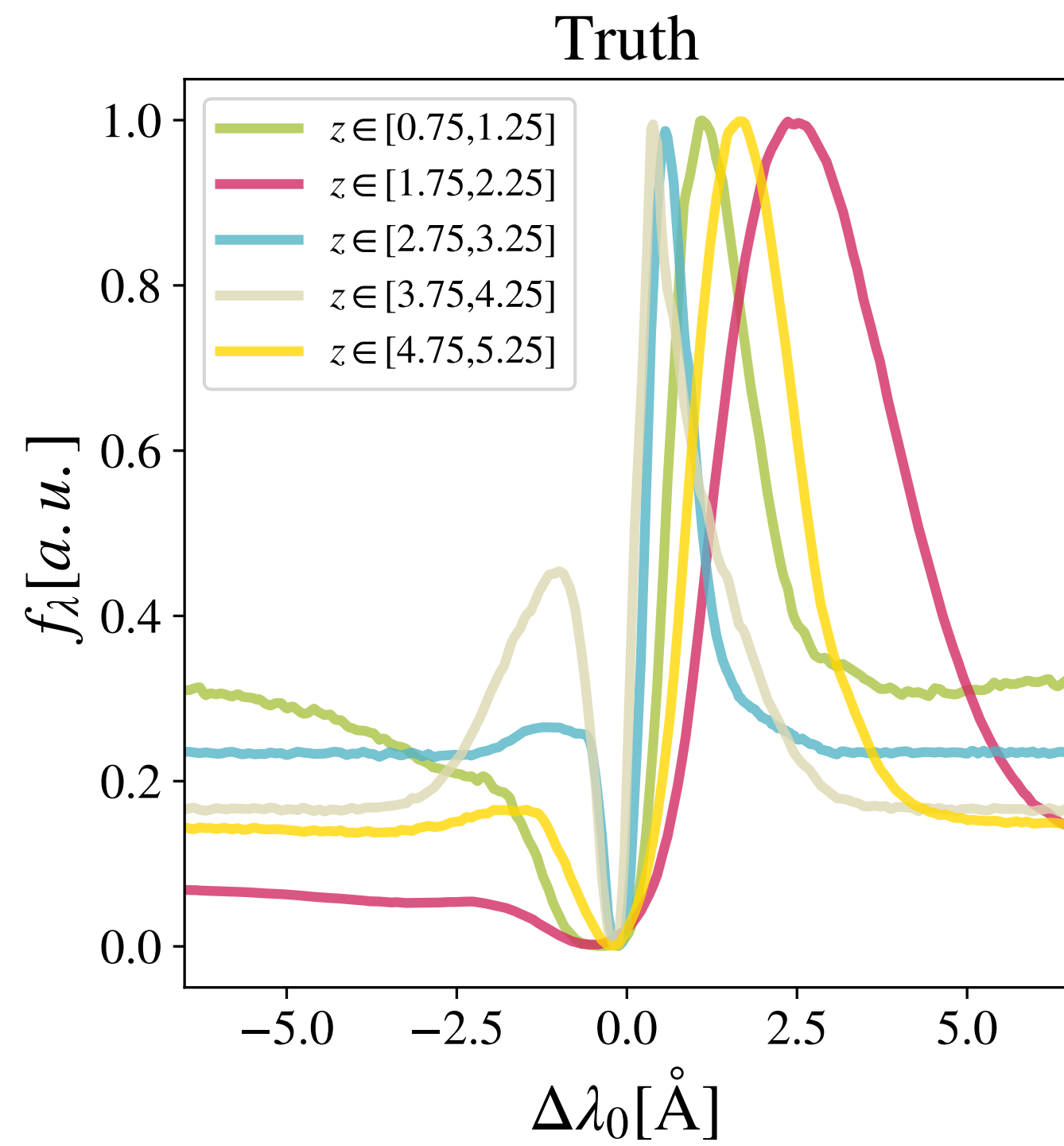






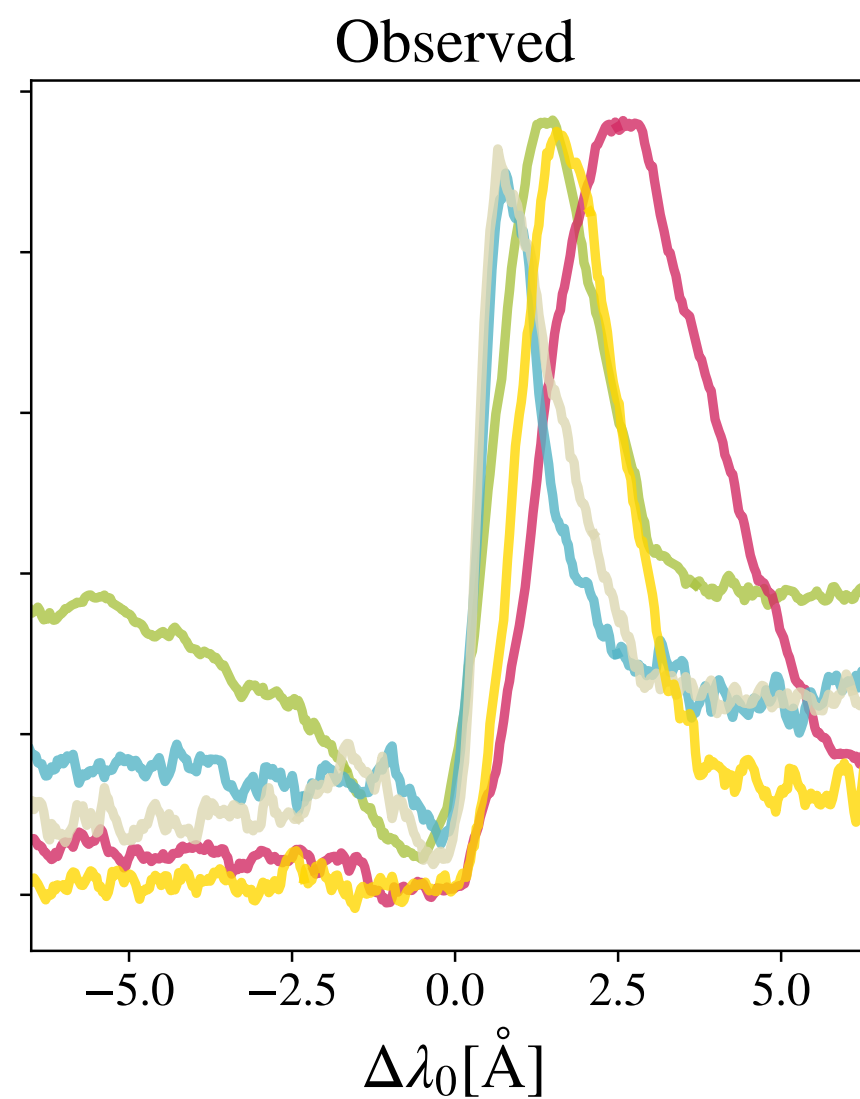
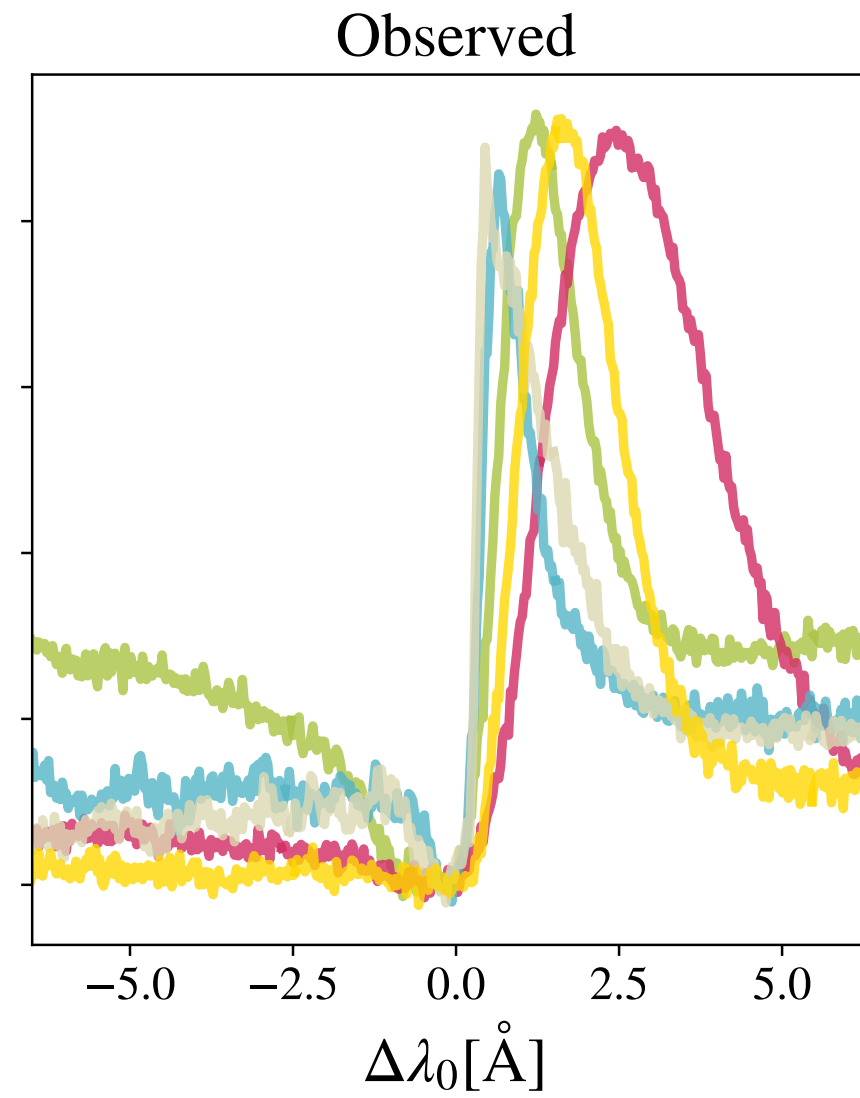


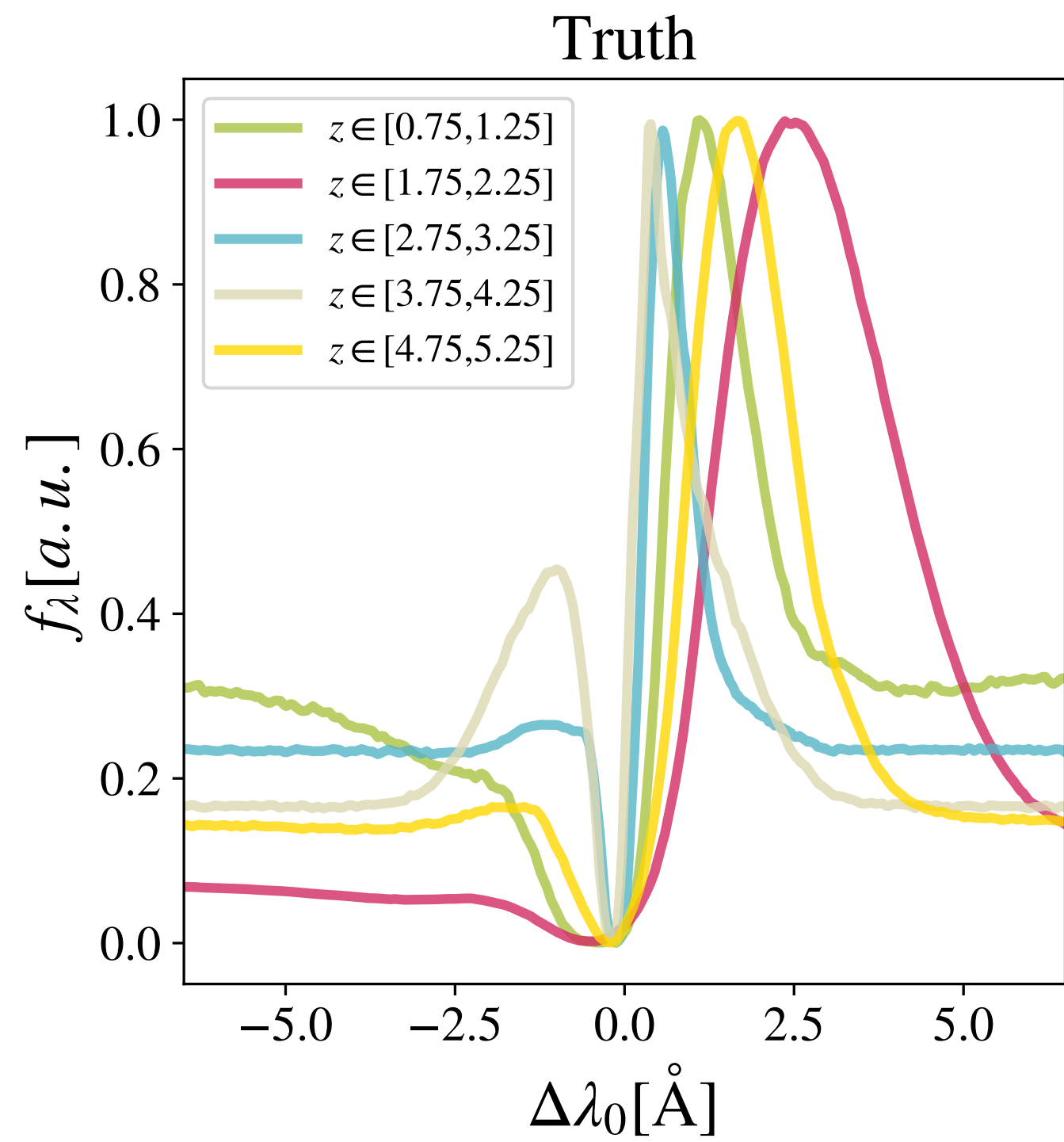




$W_g = 0.1 \text{ \AA}, S/N_p = 15.0$

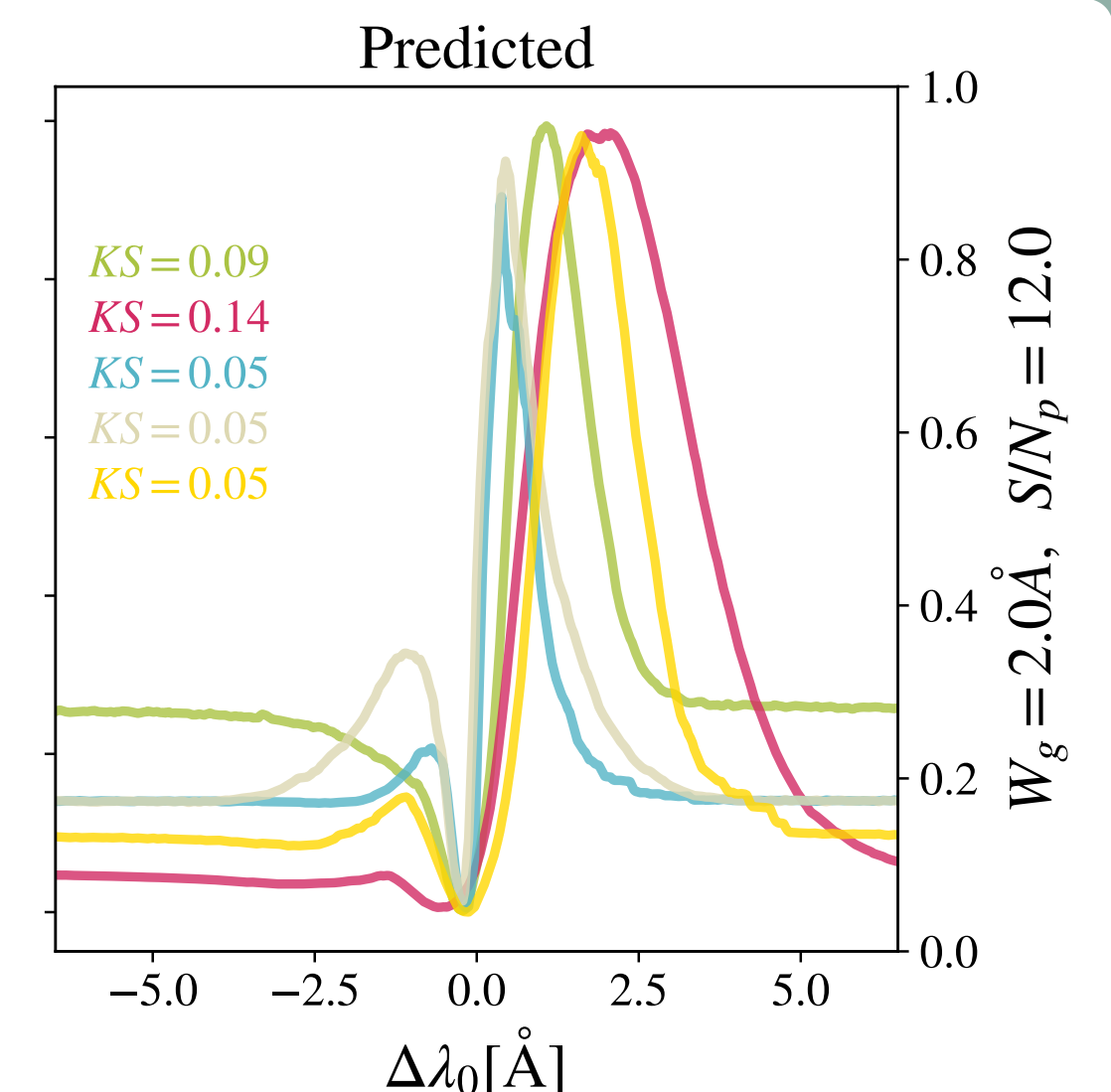
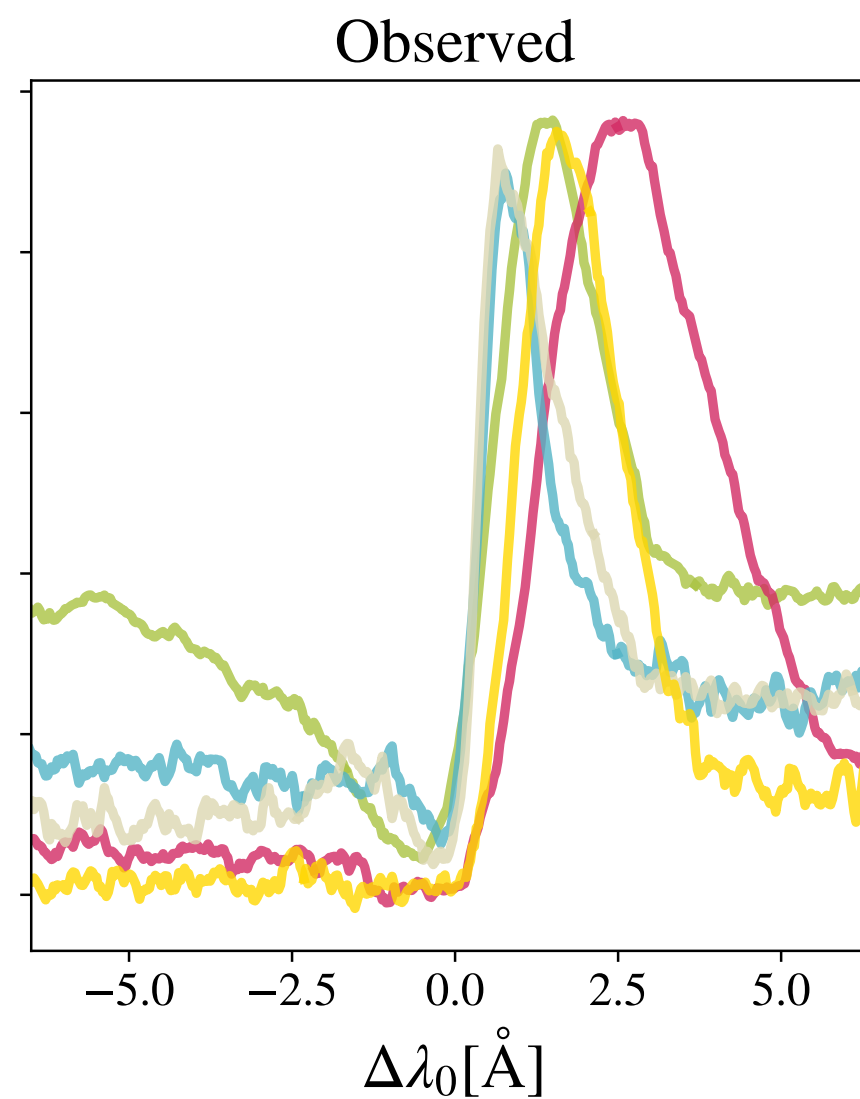
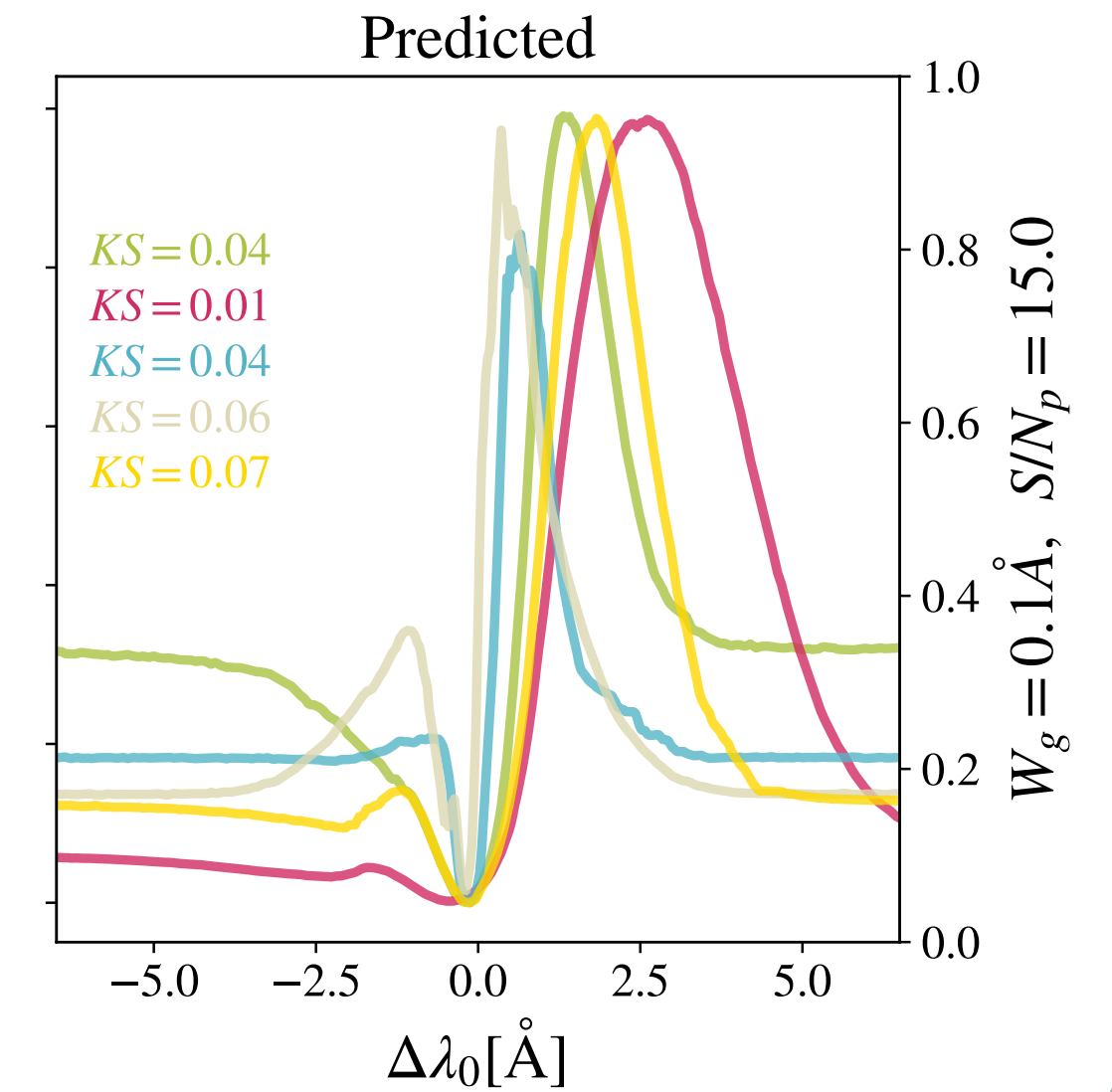
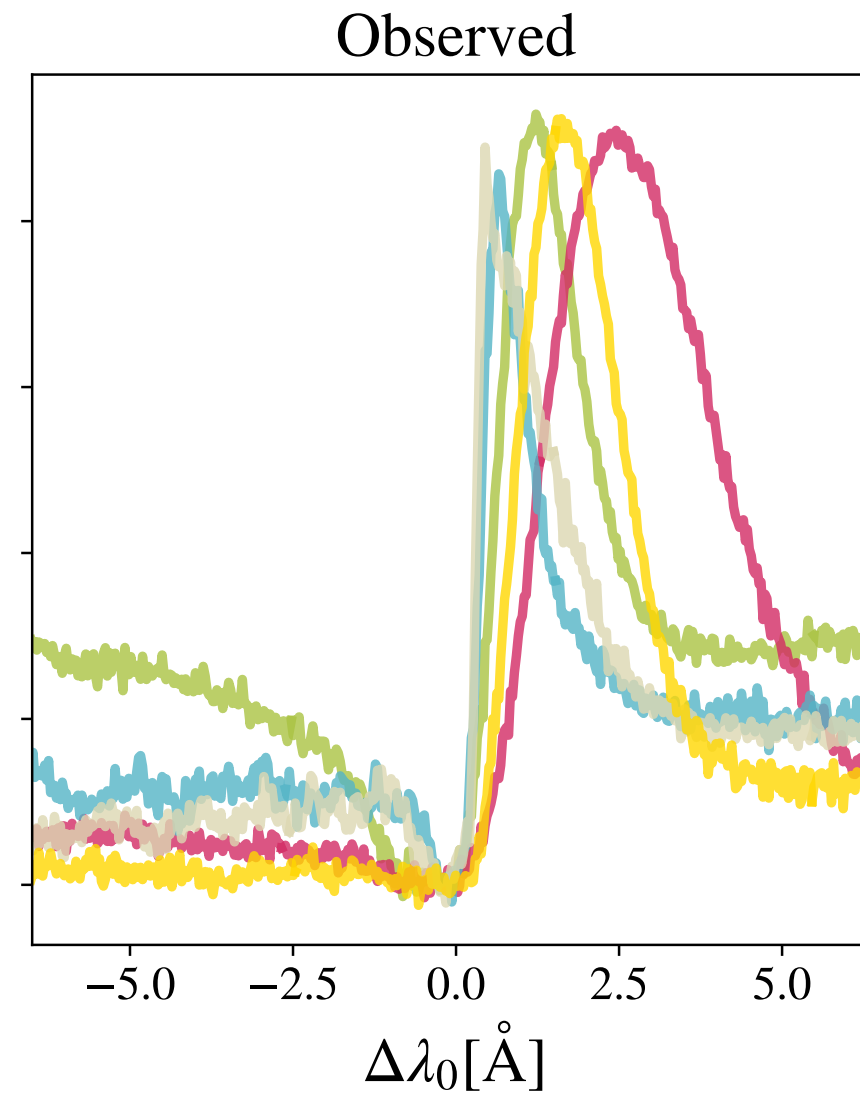
$W_g = 2.0 \text{ \AA}, S/N_p = 12.0$



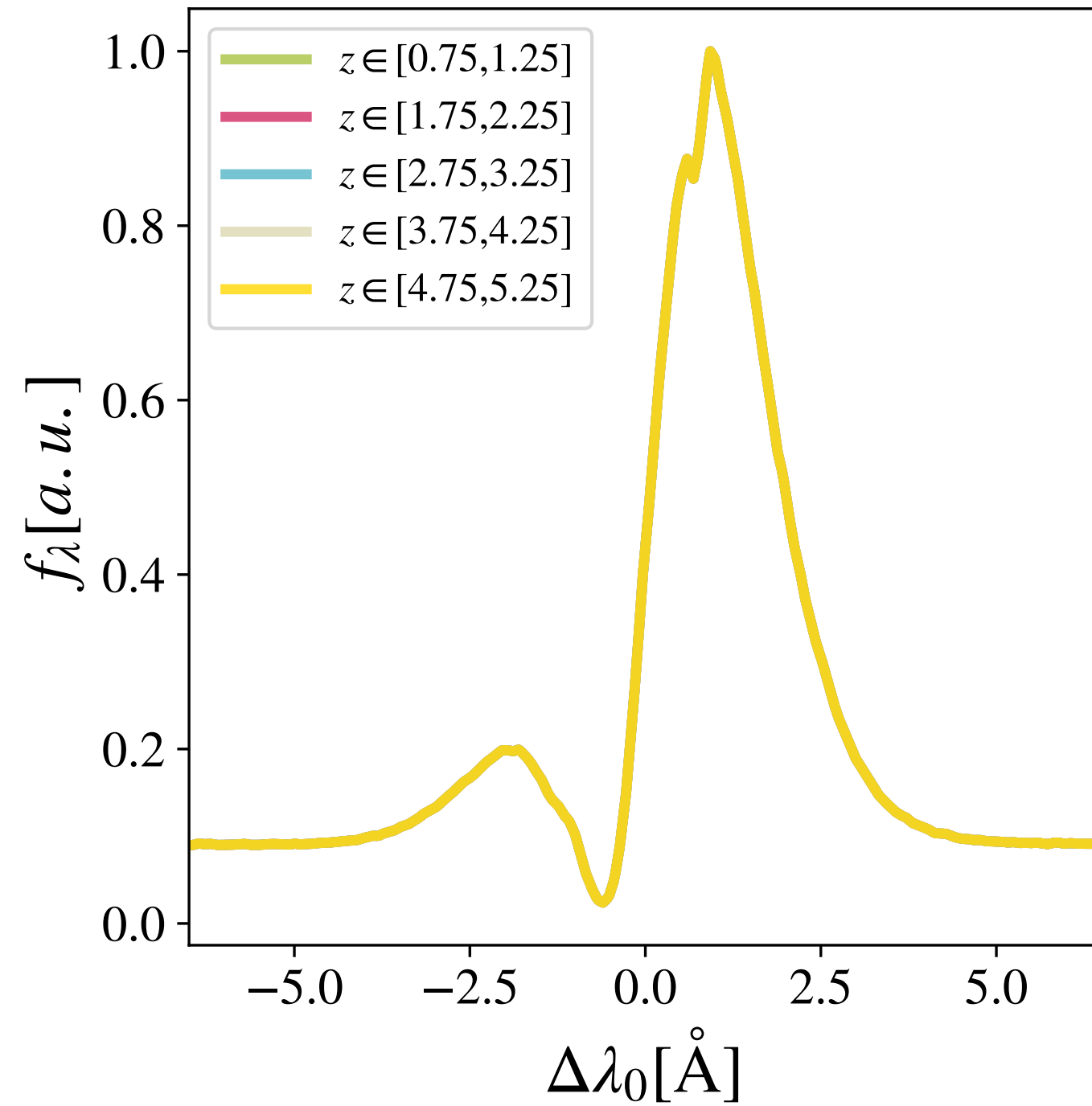


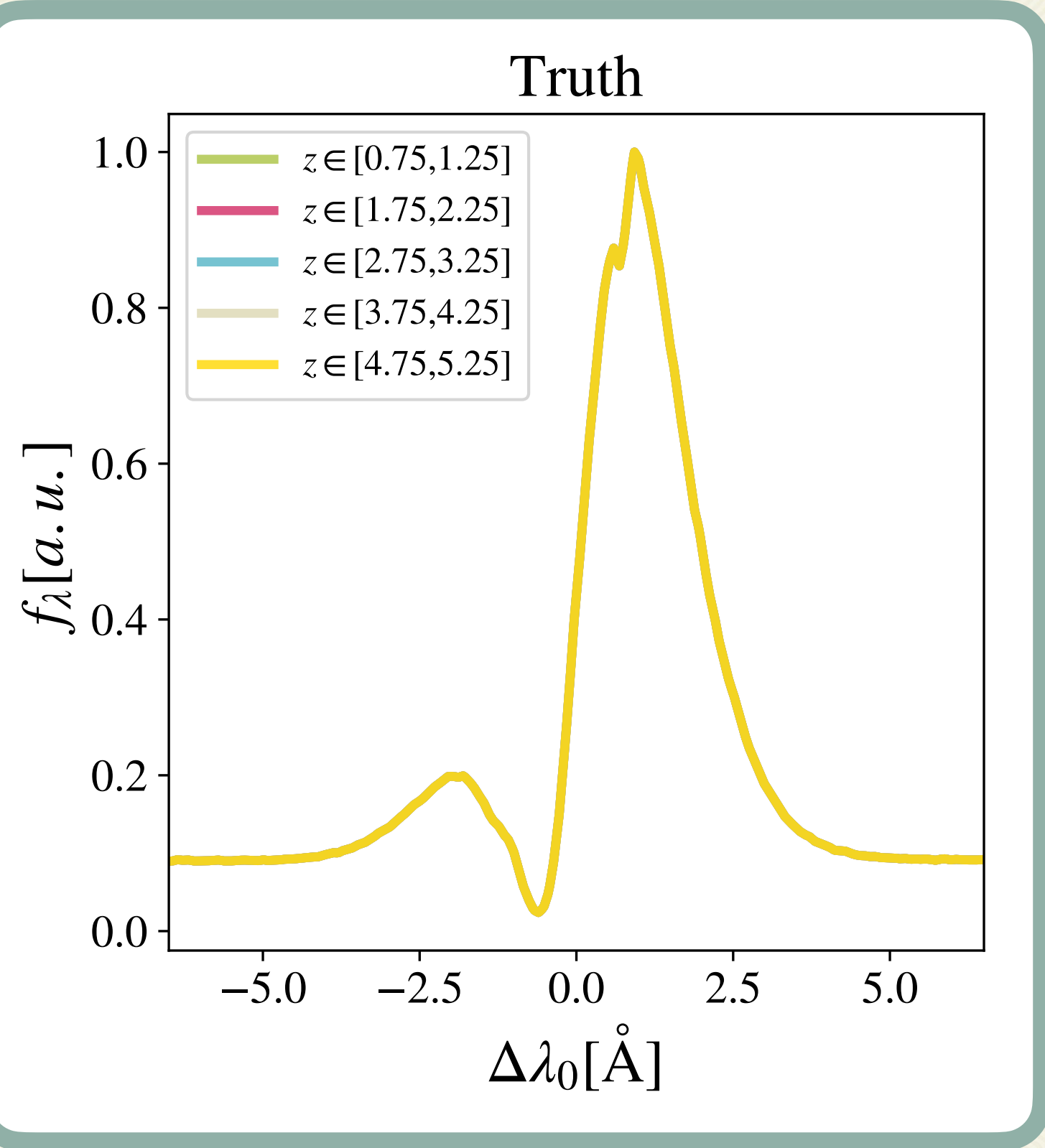
$W_g = 0.1 \text{ \AA}, S/N_p = 15.0$

$W_g = 2.0 \text{ \AA}, S/N_p = 12.0$



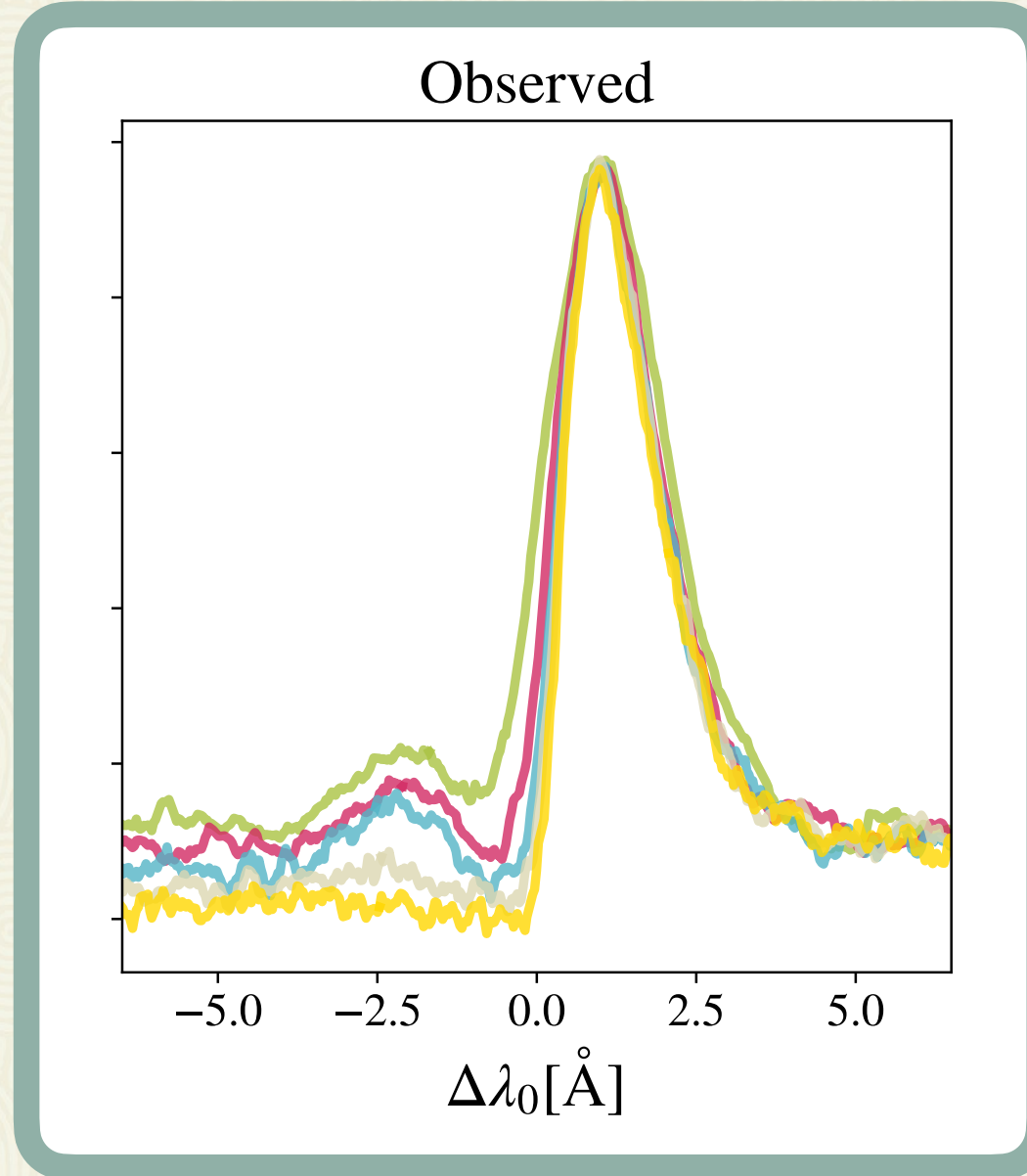
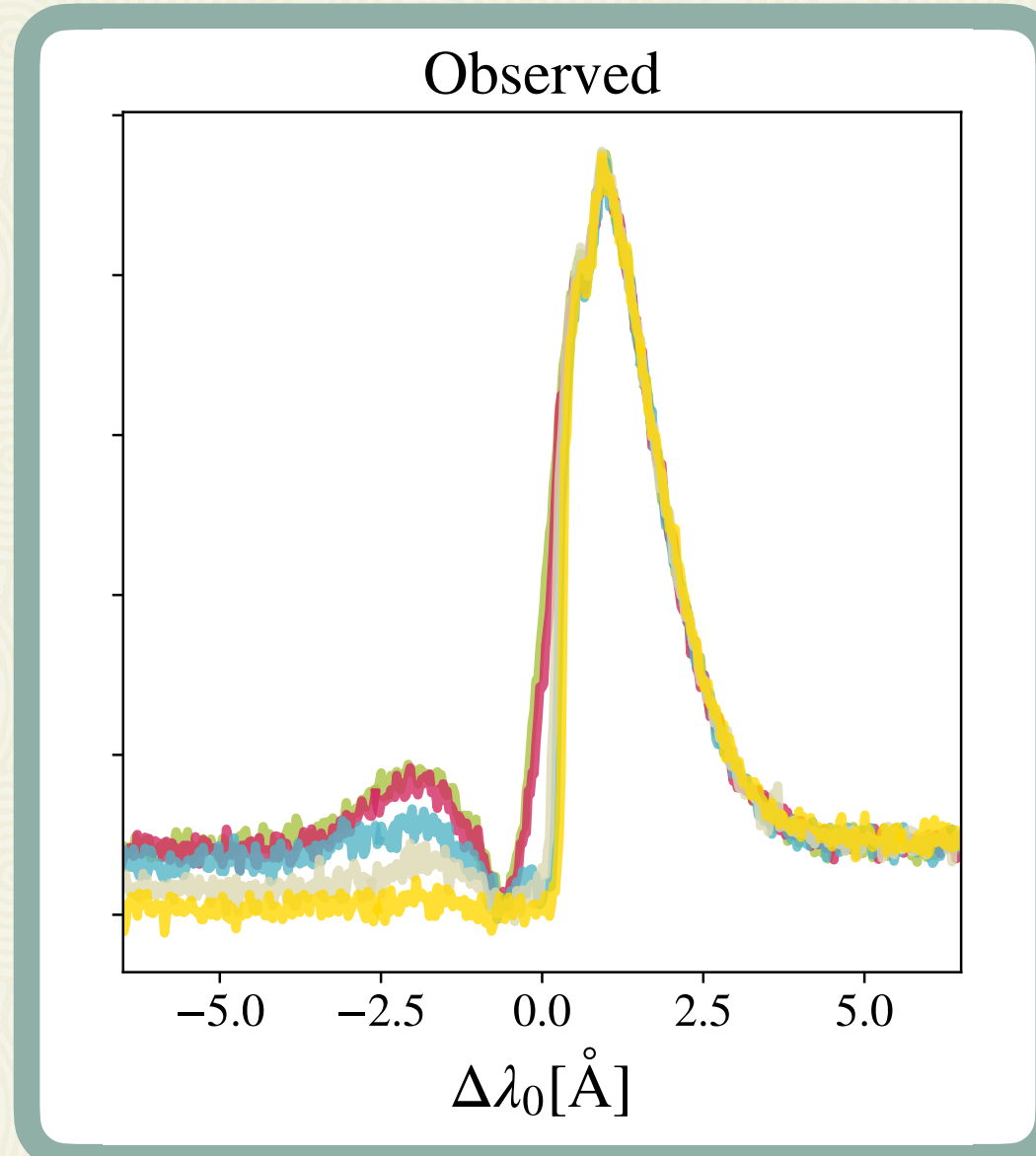
Truth

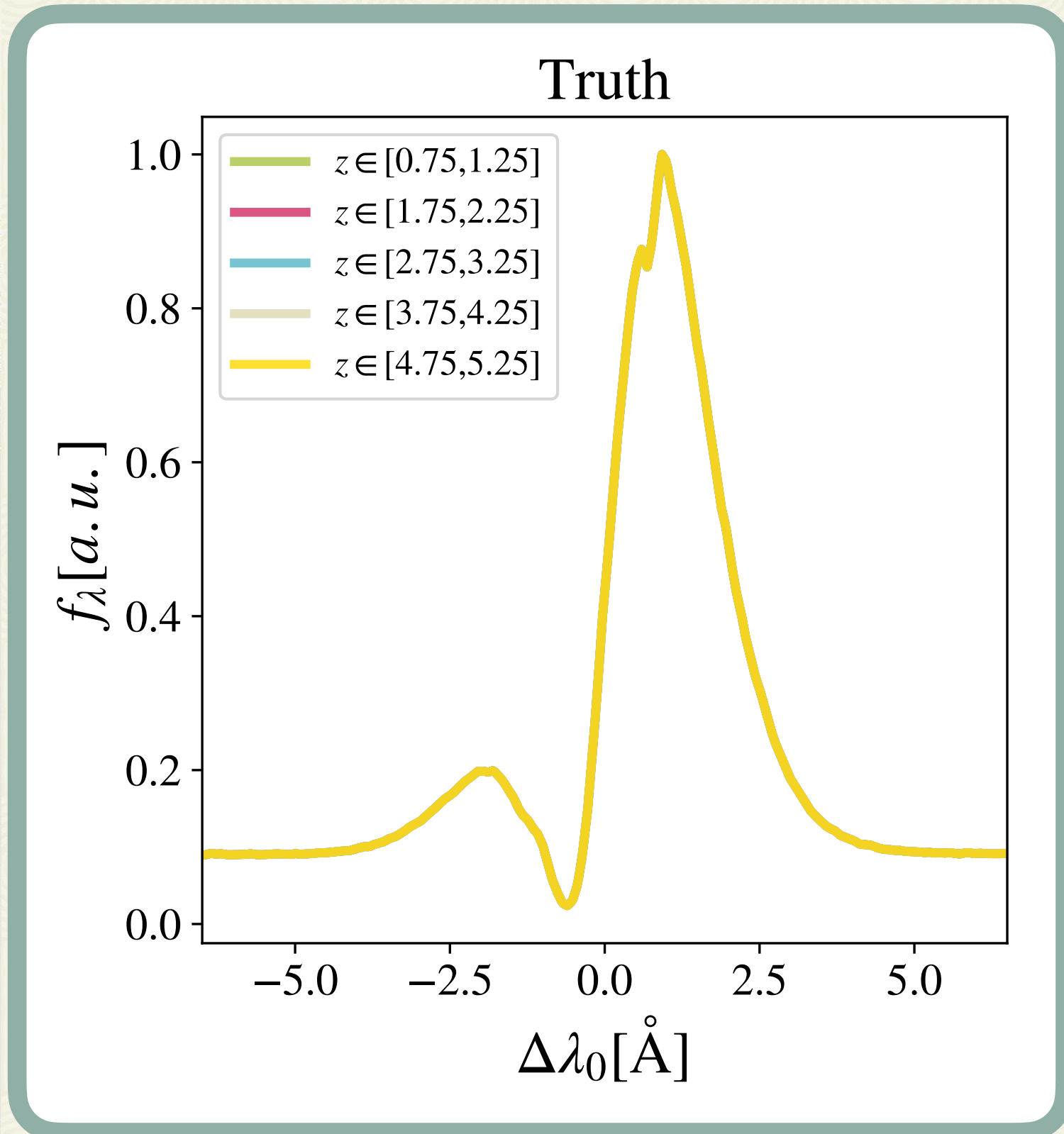




$W_g = 0.1 \text{ \AA}, S/N_p = 15.0$

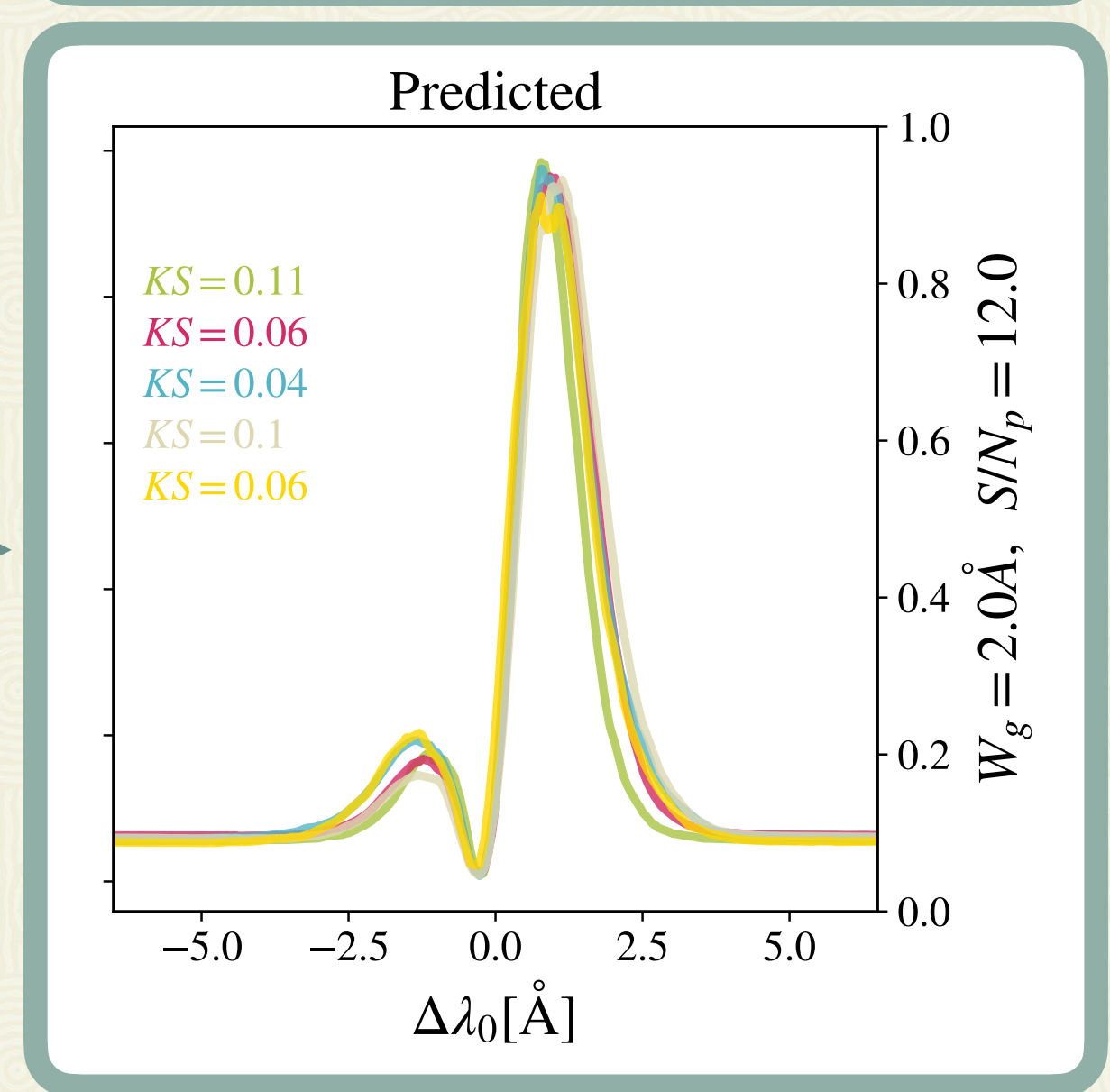
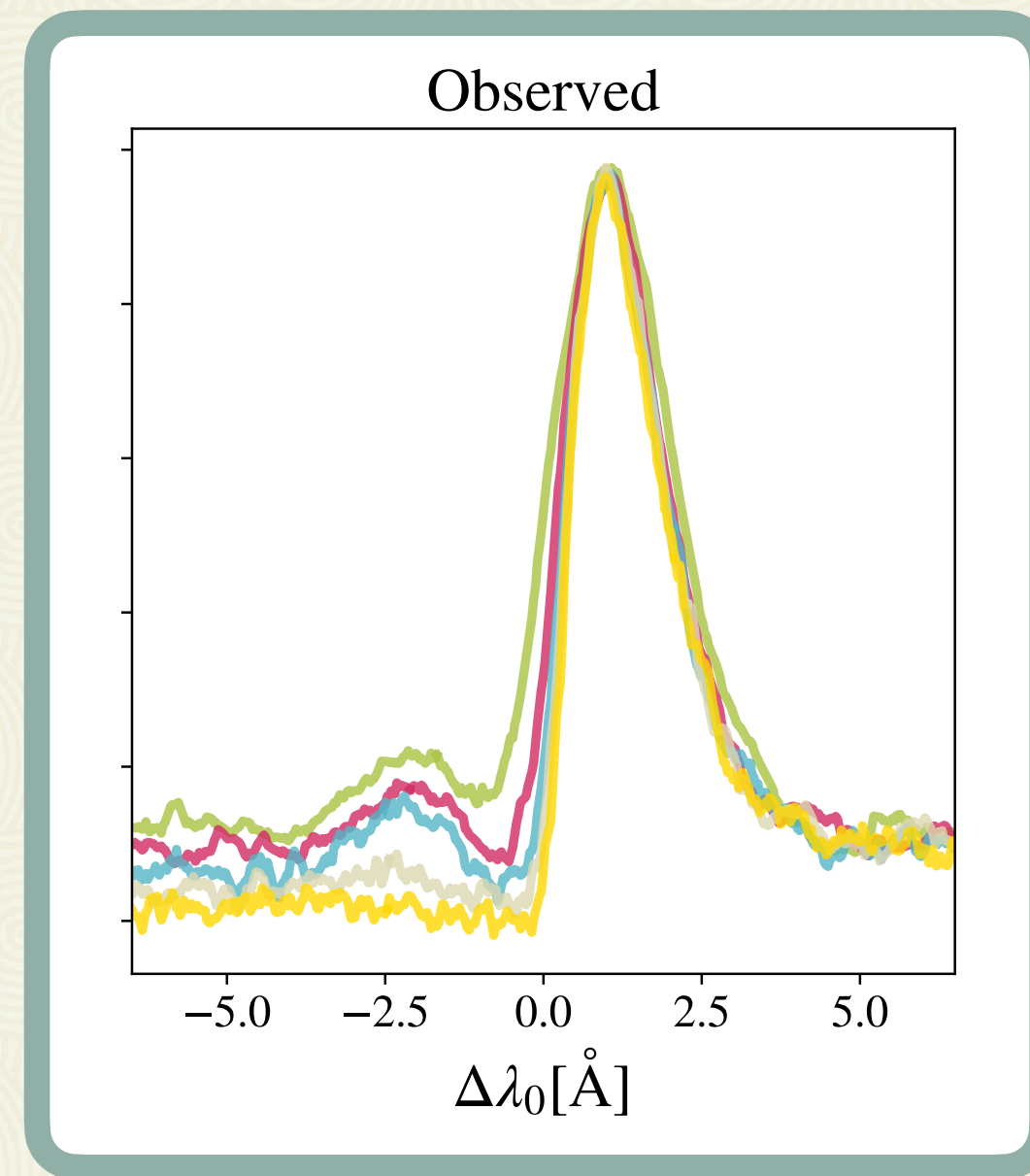
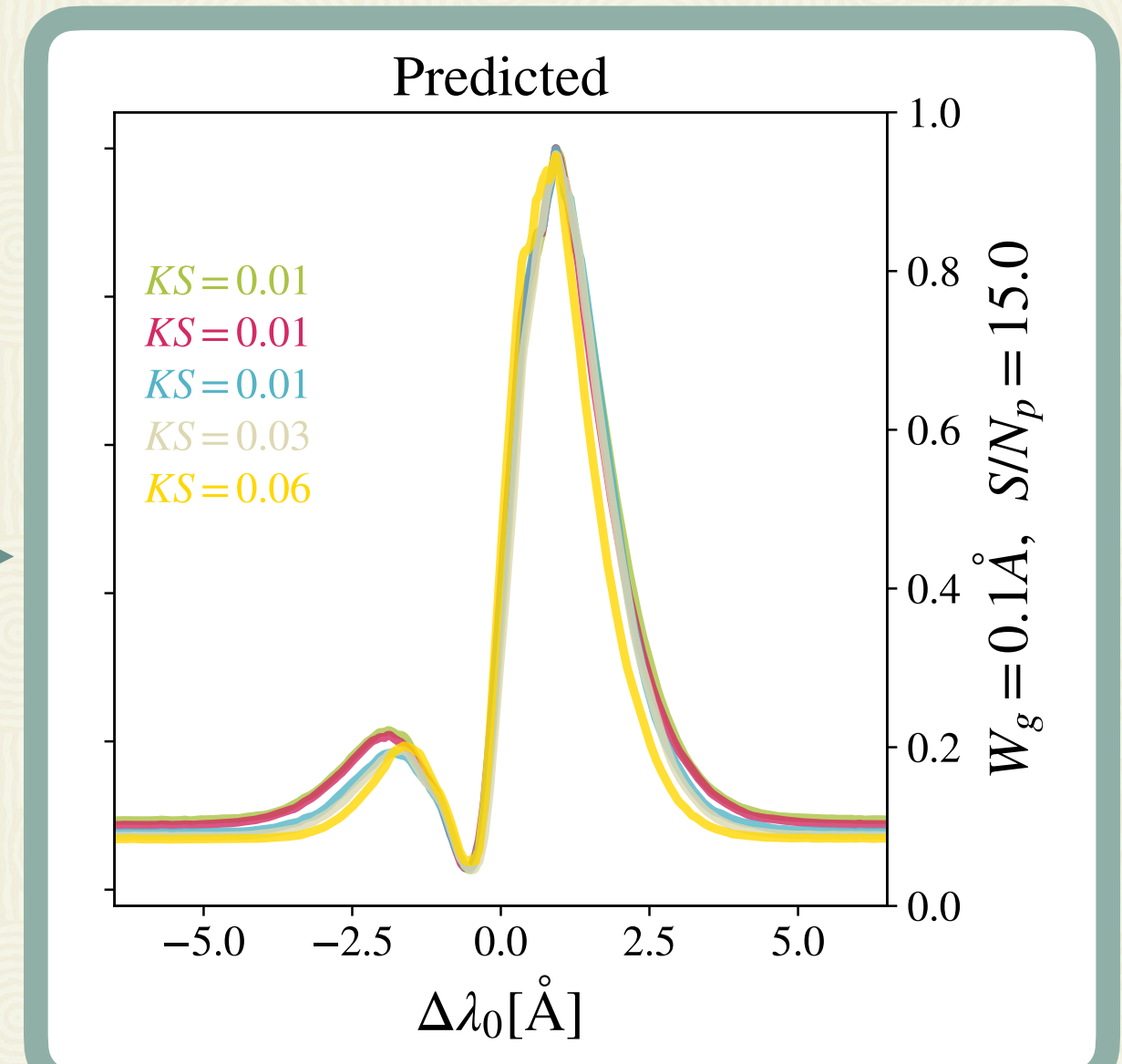
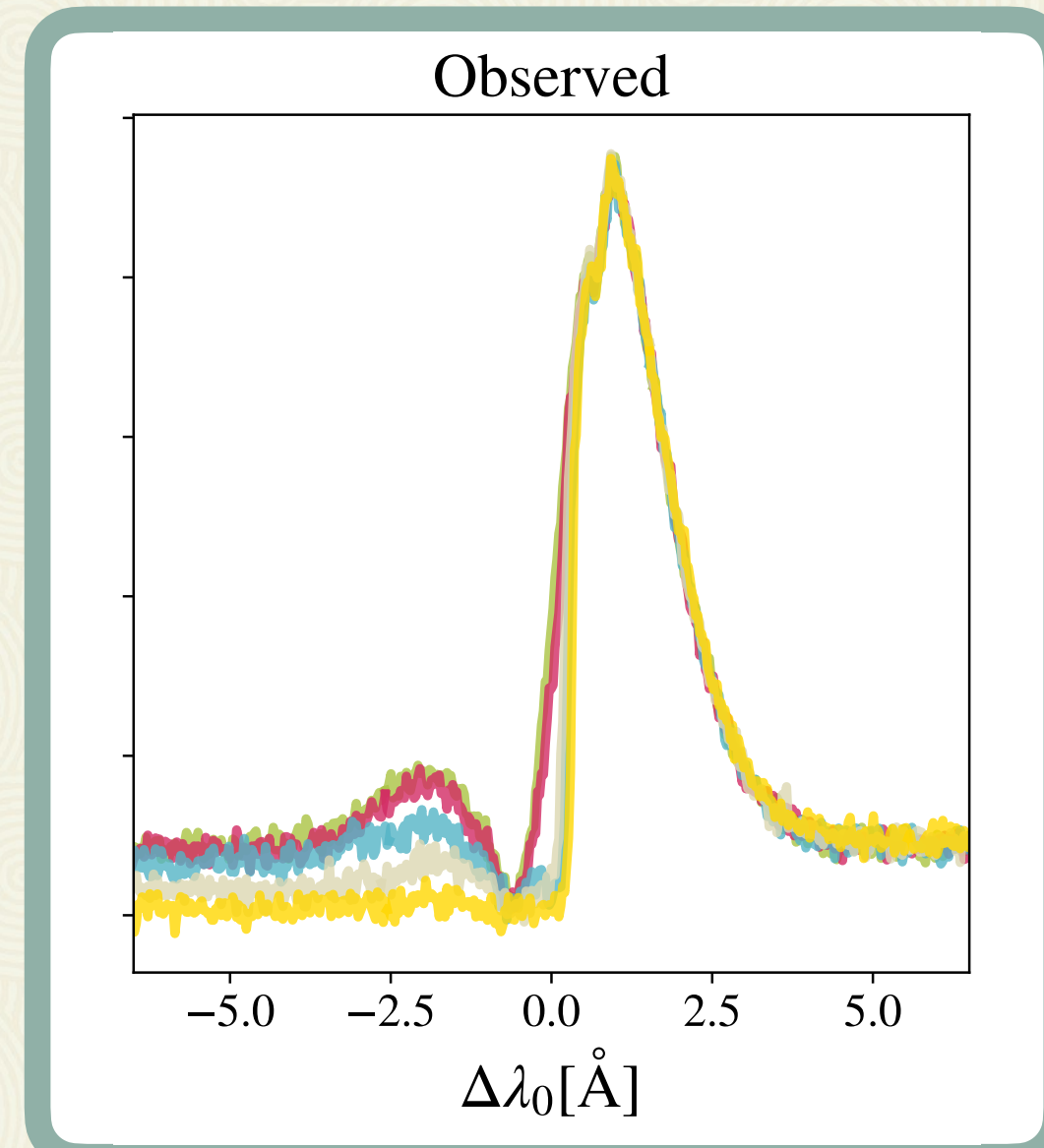
$W_g = 2.0 \text{ \AA}, S/N_p = 12.0$



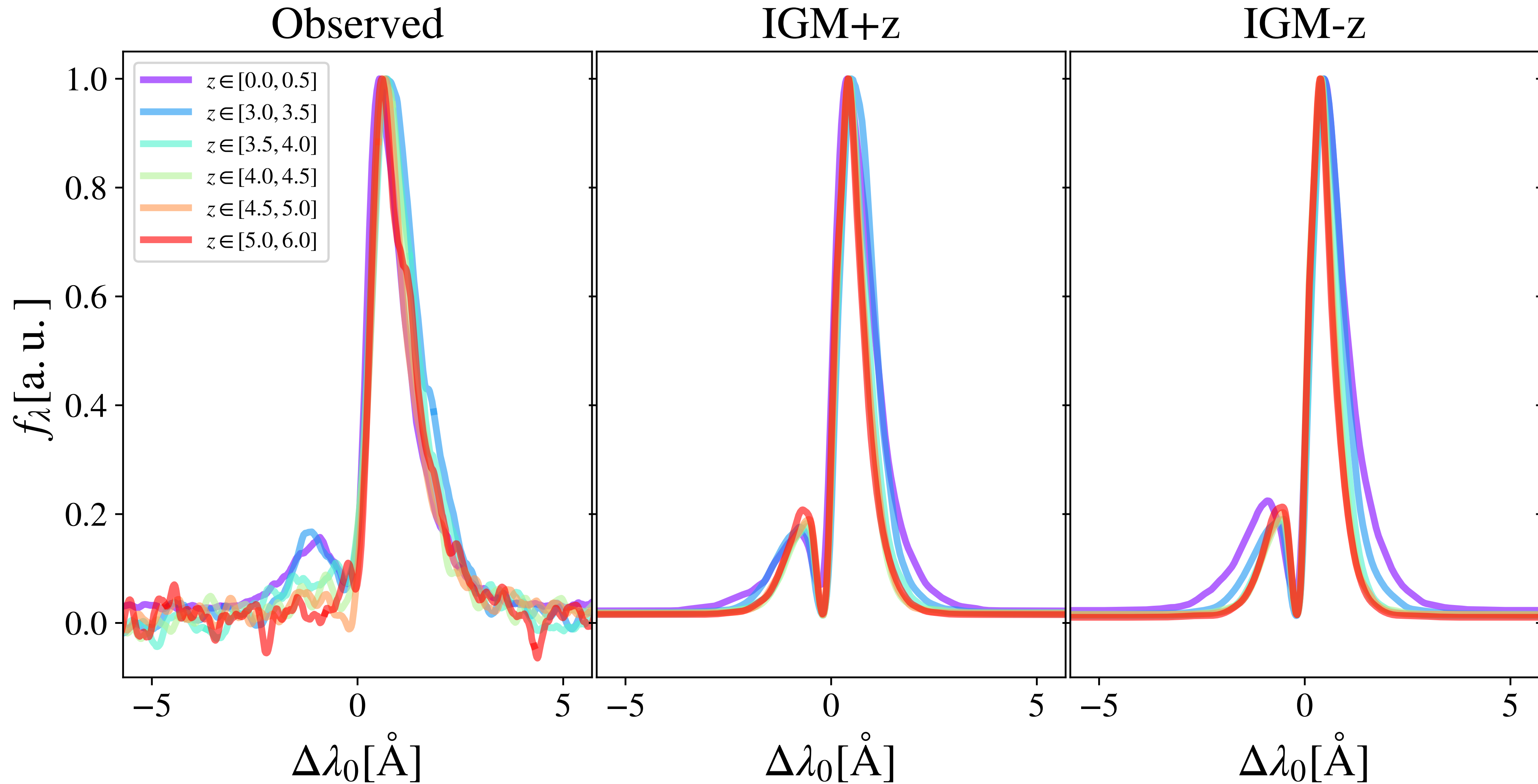


$W_g = 0.1 \text{ \AA}, S/N_p = 15.0$

$W_g = 2.0 \text{ \AA}, S/N_p = 12.0$







Lyman-alpha spectrum exhibits a large variety of features.

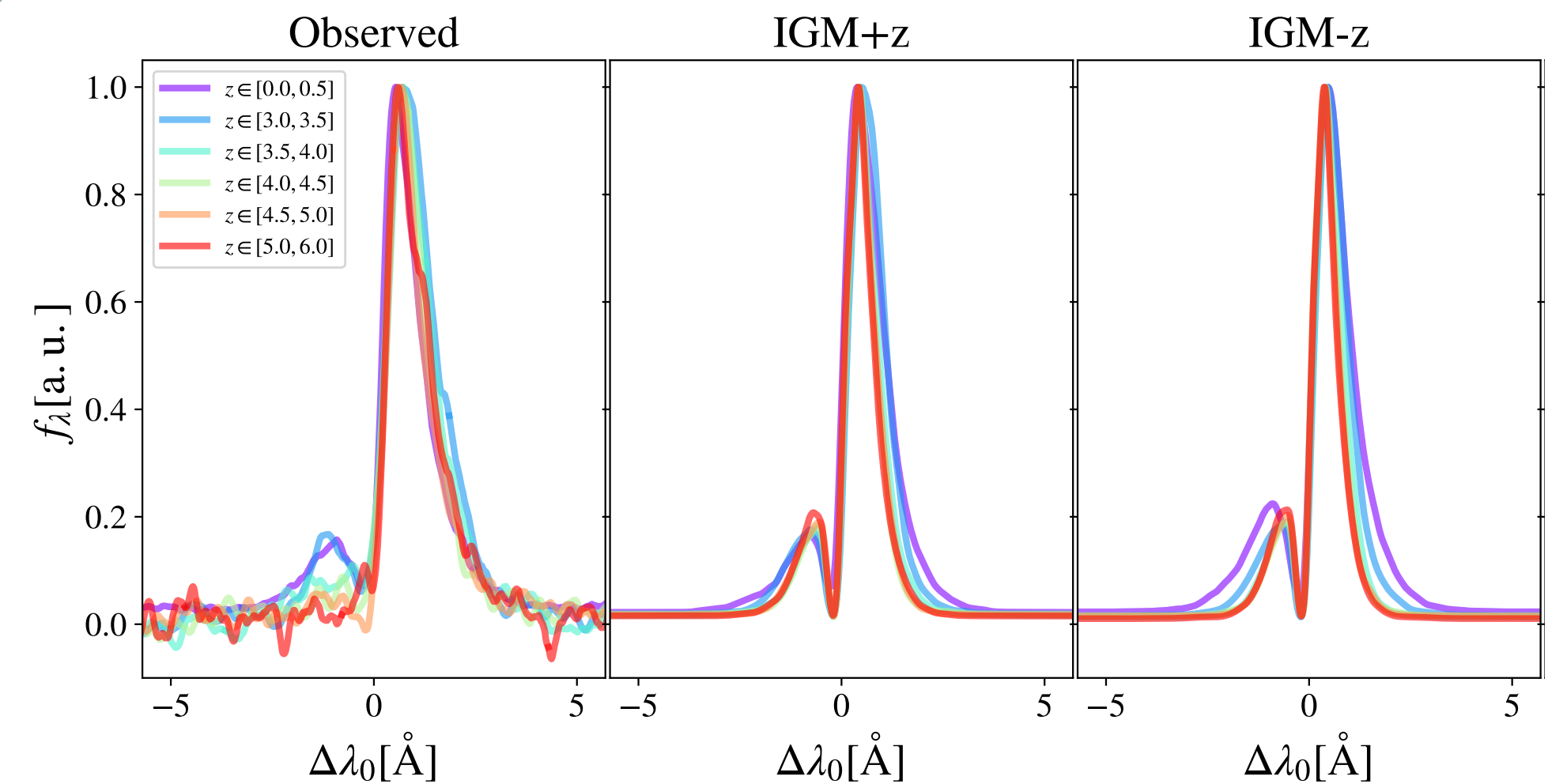
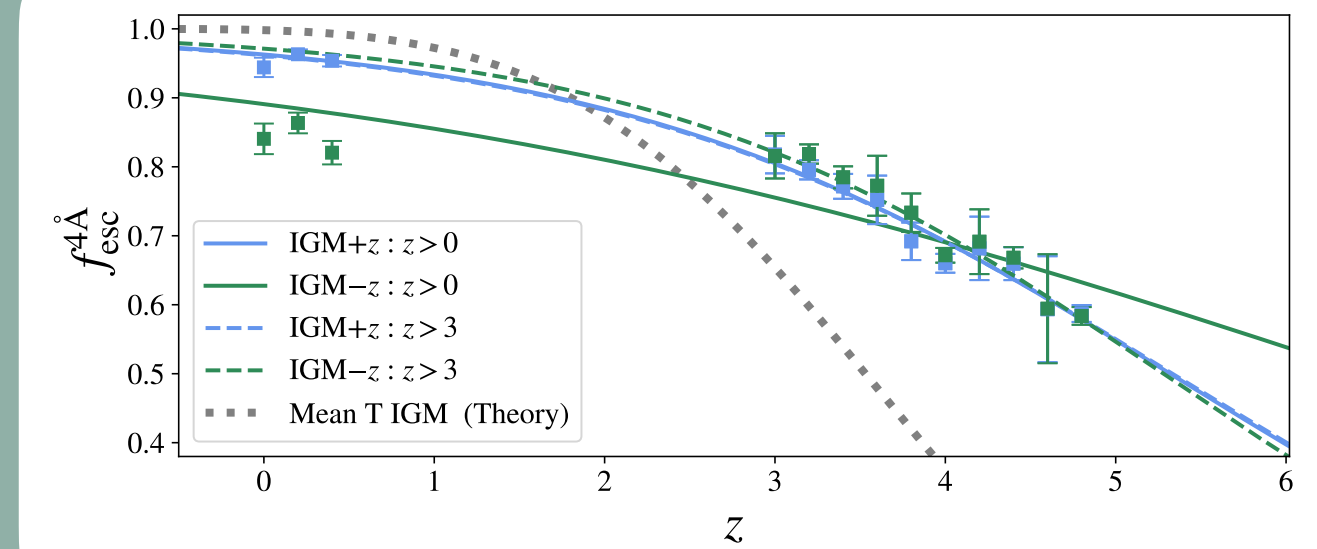
Lyman-alpha lines are shaped by the interestelar, the circumgalactic and intergalactic medium.

zELDA is an open source code to mock and fit Lyman-alpha spectrum.

zELDA is able to disentangle between the IGM / CGM and ISM contributions to the spectral features.

zELDA is able to compute IGM escape fractions for individual line profiles

The new zELDA version will be available soon



## Thanks!

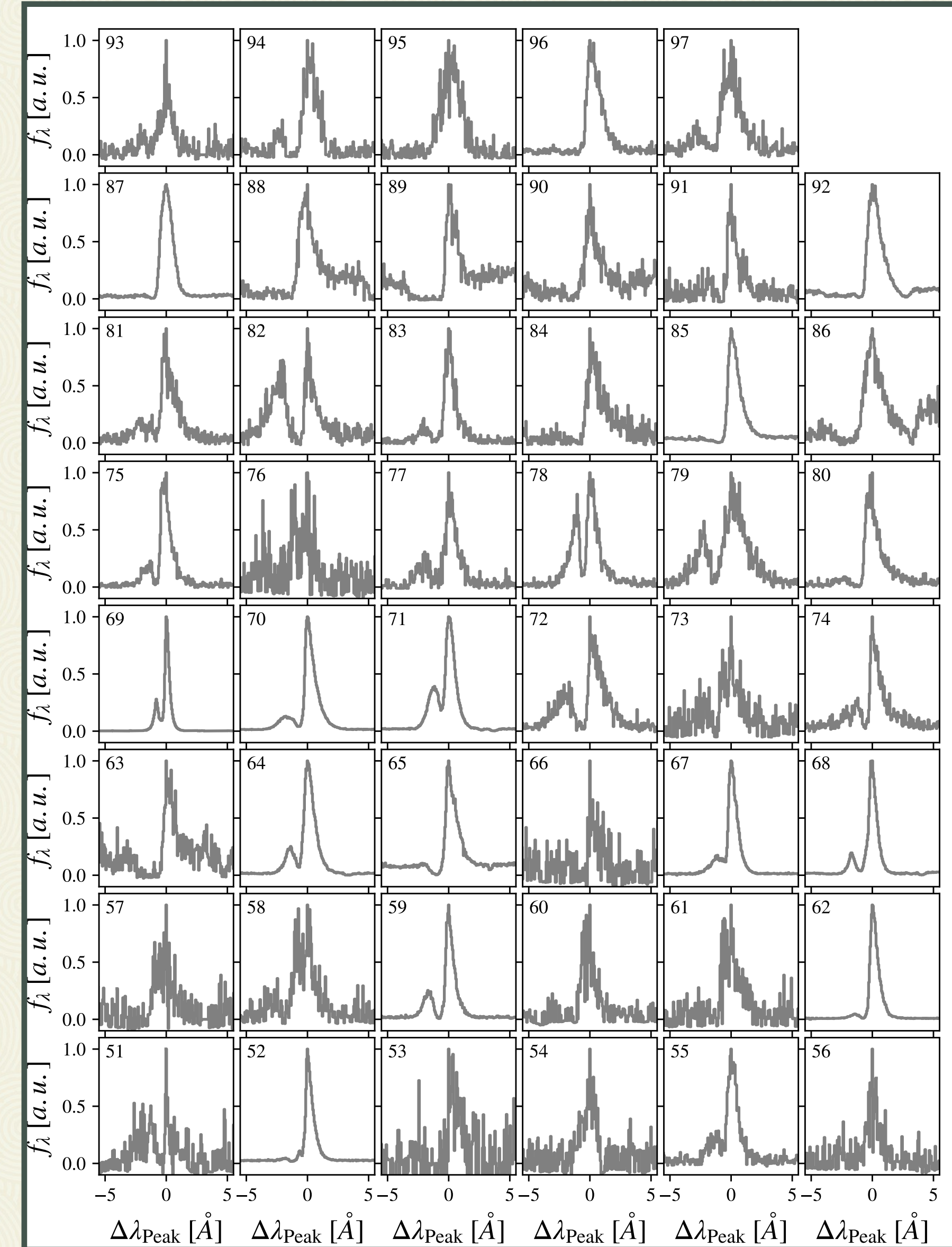
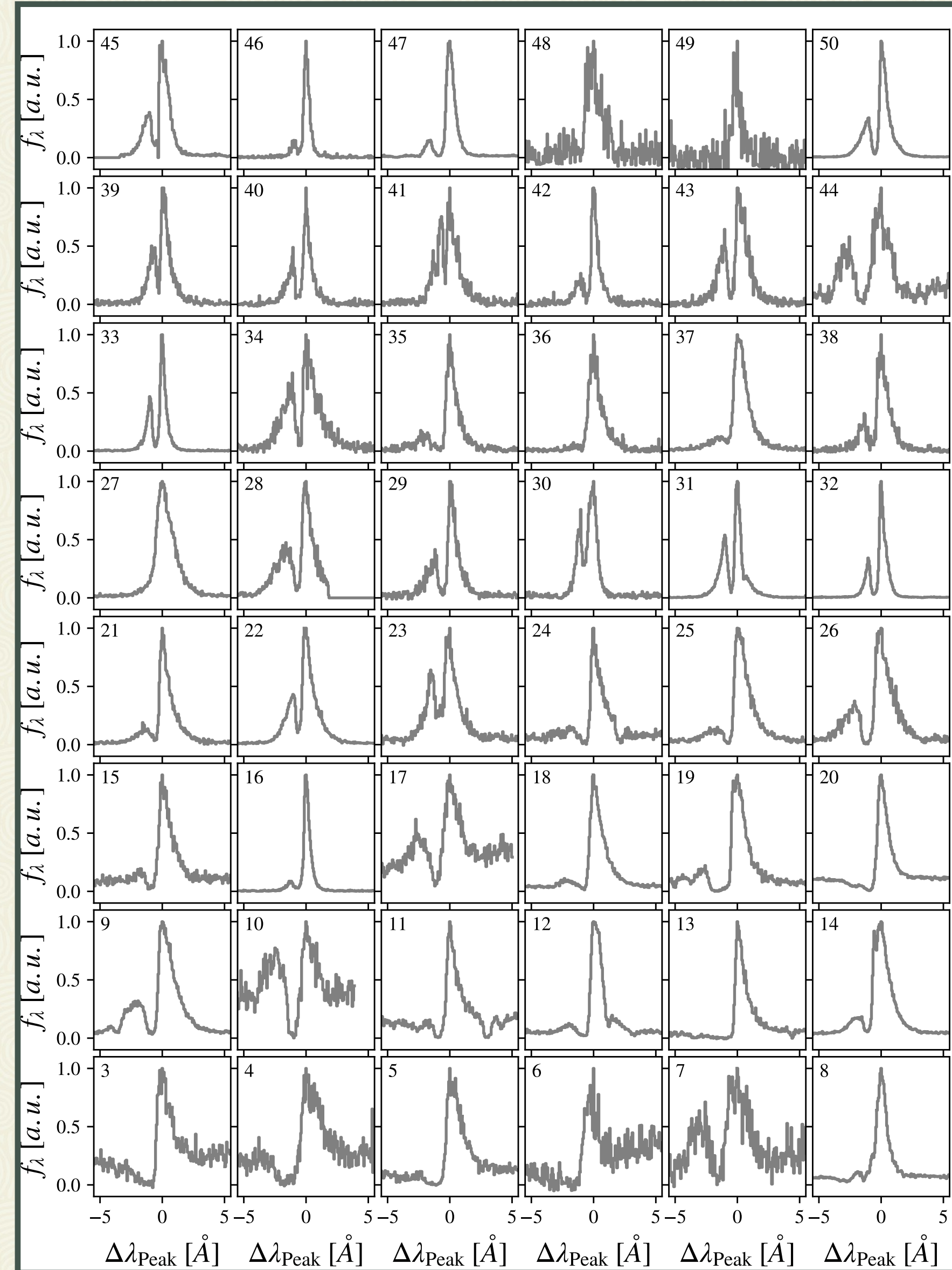




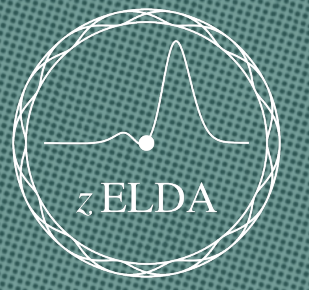
<http://lasd.lyman-alpha.com>

Runnholm et al.+20

97 Ly $\alpha$  line profiles obtained by the Cosmic Origins Spectrograph on HST  
 GO 11522 and 12027 (PI: Green, Salzer et al. 2001)  
 GO 11727 and 13017 (PI: Heckman et al. 2011,2015)  
 GO 12269 (PI: Scarlata, Songaila et al. 2018)  
 GO 12583 (PI: Hayes, Hayes et al. 2014)  
 GO 12928 (PI: Henry, Henry et al. 2015),  
 GO 13293 and 14080 (PI: Jaskot, Jaskot & Oey 2014)  
 GO 14201 (PI: Malhotra, Yang et al. 2017)  
 GO 13744 (PI: Thuan Izotov et al. 2016, 2018)



# Ly $\alpha$ as an astrophysical tool



$Q(KS=0.1), z \in [0.0,6.0]$

$f_{esc}$	S/N <sub>p</sub>	$W_g$	$W_g$	$W_g$	$W_g$	$W_g$	$W_g$
		0.1Å	0.25Å	0.5Å	1.0Å	2.0Å	4.0Å
[0.95, 1.0]	15.0	94	94	93	93	88	78
	10.0	94	94	91	88	82	69
	7.5	94	92	89	84	75	63
	5.0	90	85	77	66	57	53
[0.8, 0.95]	15.0	93	94	92	90	86	80
	10.0	94	91	91	88	83	71
	7.5	93	91	88	84	76	64
	5.0	88	86	77	68	59	54
[0.65, 0.8]	15.0	90	89	90	86	82	77
	10.0	90	88	87	84	81	76
	7.5	88	85	84	81	76	68
	5.0	86	84	78	73	65	63
[0.5, 0.65]	15.0	88	89	87	82	78	79
	10.0	88	86	84	80	77	76
	7.5	87	85	83	81	76	72
	5.0	85	80	78	73	70	64
[0.35, 0.5]	15.0	84	83	83	78	73	73
	10.0	85	81	79	78	70	69
	7.5	81	81	79	75	69	68
	5.0	81	74	73	68	65	61
[0.2, 0.35]	15.0	71	68	67	61	65	60
	10.0	67	67	62	64	61	56
	7.5	72	63	65	58	57	58
	5.0	64	60	58	62	51	55

$Q(KS=0.1), z \in [0.0,6.0]$

$f_{esc}$	S/N <sub>p</sub>	$W_g$	$W_g$	$W_g$	$W_g$	$W_g$	$W_g$
		0.1Å	0.25Å	0.5Å	1.0Å	2.0Å	4.0Å
[0.95, 1.0]	15.0	96	95	96	95	90	78
	10.0	94	95	94	92	83	69
	7.5	95	94	92	88	77	62
	5.0	93	88	80	70	57	49
[0.8, 0.95]	15.0	97	97	96	95	90	80
	10.0	96	95	95	92	85	73
	7.5	96	94	93	89	79	65
	5.0	92	88	81	72	60	52
[0.65, 0.8]	15.0	95	95	95	90	86	80
	10.0	94	94	93	88	84	77
	7.5	93	92	90	86	81	70
	5.0	89	87	82	77	68	61
[0.5, 0.65]	15.0	94	95	92	87	83	80
	10.0	94	93	91	85	81	78
	7.5	93	91	89	84	80	72
	5.0	90	86	83	79	73	65
[0.35, 0.5]	15.0	90	92	86	81	76	77
	10.0	90	87	82	79	72	72
	7.5	84	86	82	77	73	70
	5.0	80	78	79	75	72	67
[0.2, 0.35]	15.0	72	68	66	60	62	62
	10.0	68	65	65	61	64	59
	7.5	70	63	65	60	62	59
	5.0	61	62	60	65	57	59

$Q(KS=0.1), z \in [0.0,6.0]$

$f_{esc}$	S/N <sub>p</sub>	$W_g$	$W_g$	$W_g$	$W_g$	$W_g$	$W_g$
		0.1Å	0.25Å	0.5Å	1.0Å	2.0Å	4.0Å
[0.95, 1.0]	15.0	98	98	98	97	95	92
	10.0	98	98	97	96	93	86
	7.5	98	96	96	93	87	81
	5.0	94	93	86	80	73	67
[0.8, 0.95]	15.0	85	86	84	83	80	79
	10.0	85	84	83	81	79	75
	7.5	83	85	82	79	76	70
	5.0	82	81	77	71	68	65
[0.65, 0.8]	15.0	62	66	62	62	58	54
	10.0	61	65	63	60	55	57
	7.5	62	63	64	61	59	55
	5.0	62	63	65	63	59	56
[0.5, 0.65]	15.0	50	51	48	41	38	38
	10.0	49	48	48	43	39	39
	7.5	50	49	47	46	39	42
	5.0	50	52	51	50	50	48
[0.35, 0.5]	15.0	32	36	32	26	24	23
	10.0	34	34	33	28	23	26
	7.5	35	38	36	33	30	33
	5.0	41	41	45	39	41	41
[0.2, 0.35]	15.0	29	26	24	14	20	18
	10.0	28	23	24	23	20	19
	7.5	32	30	25	26	22	25
	5.0	33	32	30	35	31	38