

The Scintillating Bubble Chamber Experiment

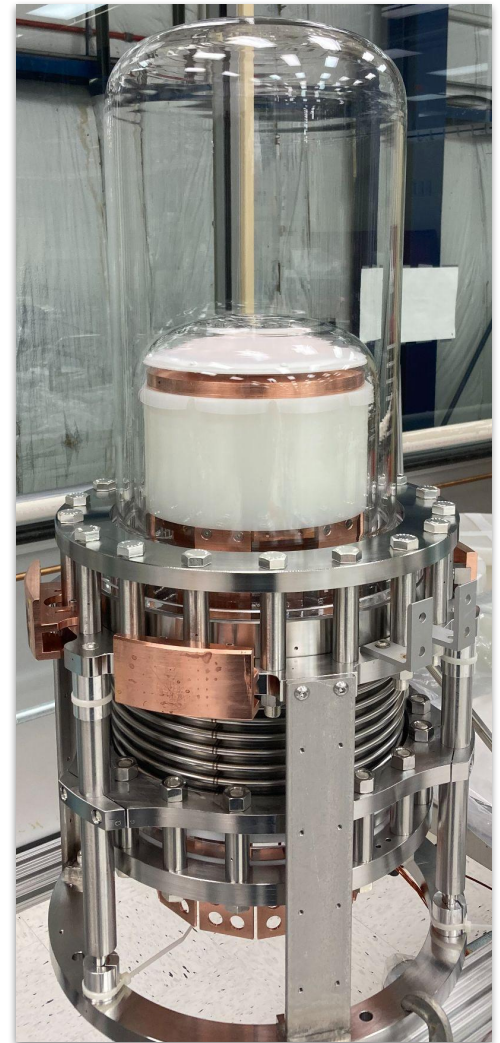


Ben Broerman
for the SBC collaboration

IDM 2026

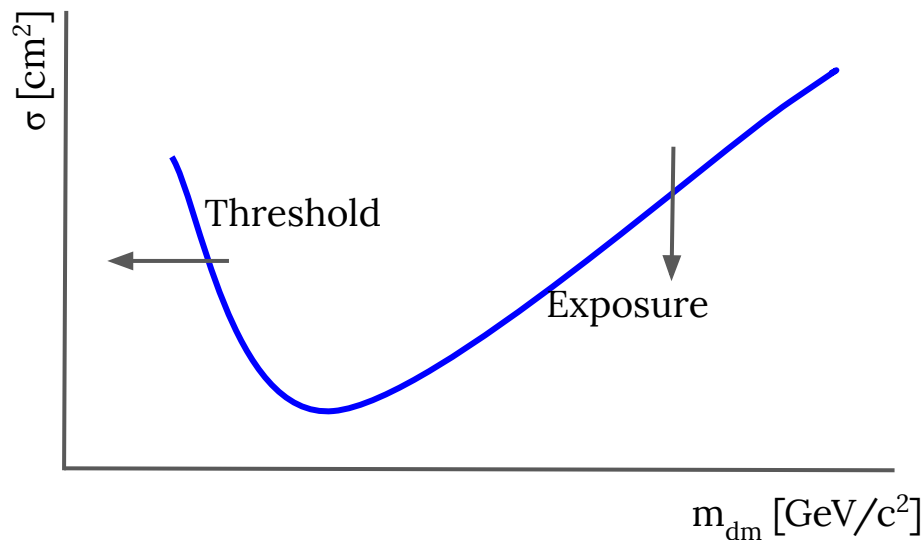


Arthur B. McDonald
Canadian Astroparticle Physics Research Institute



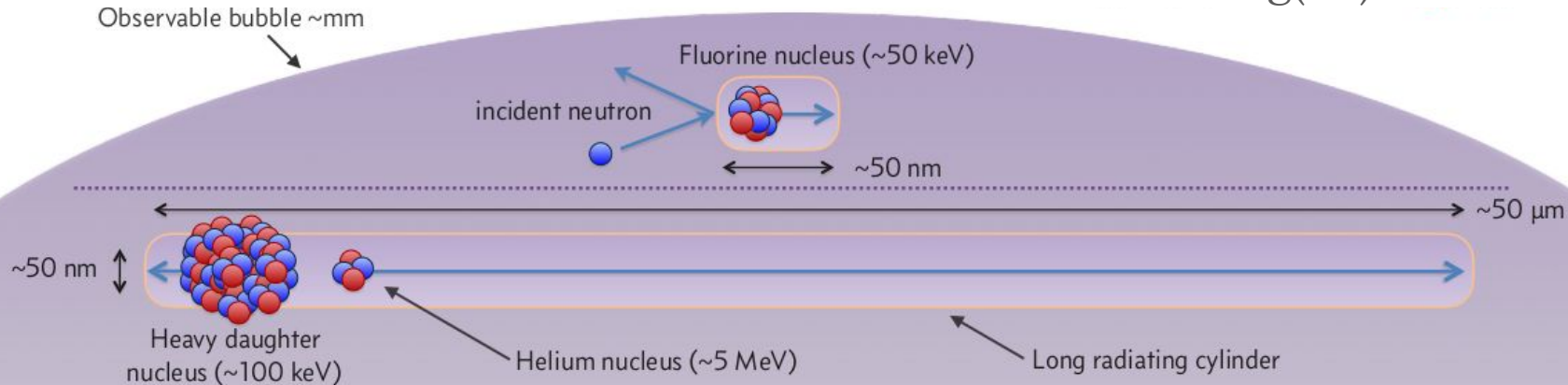
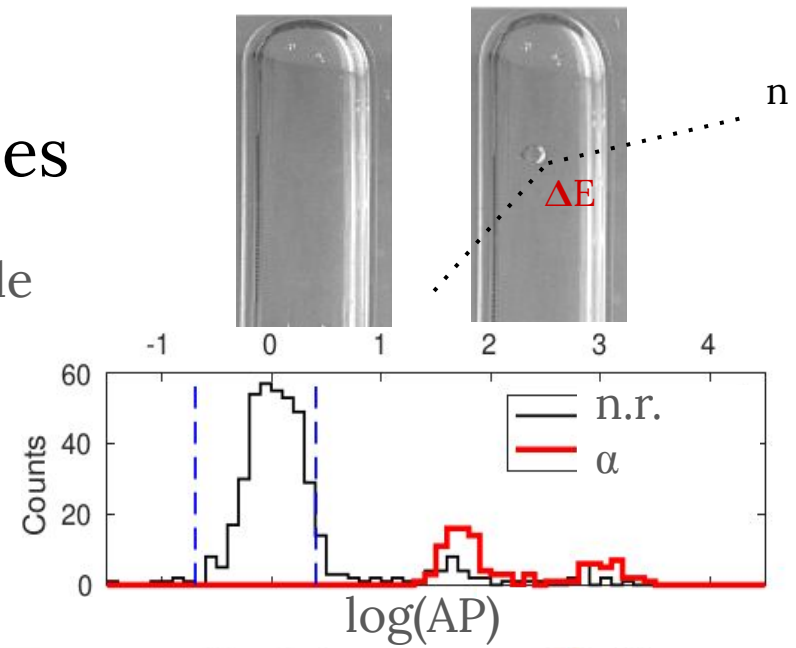
sub-keV nuclear recoils

- Difficult kinematics of GeV-scale DM-nucleon scattering require low thresholds ($\lesssim 1$ keV)
 - Future success here needs background discrimination **and** scalability
- Similar challenge measuring CE ν NS from MeV-scale ν 's



Bubble chambers for DM searches

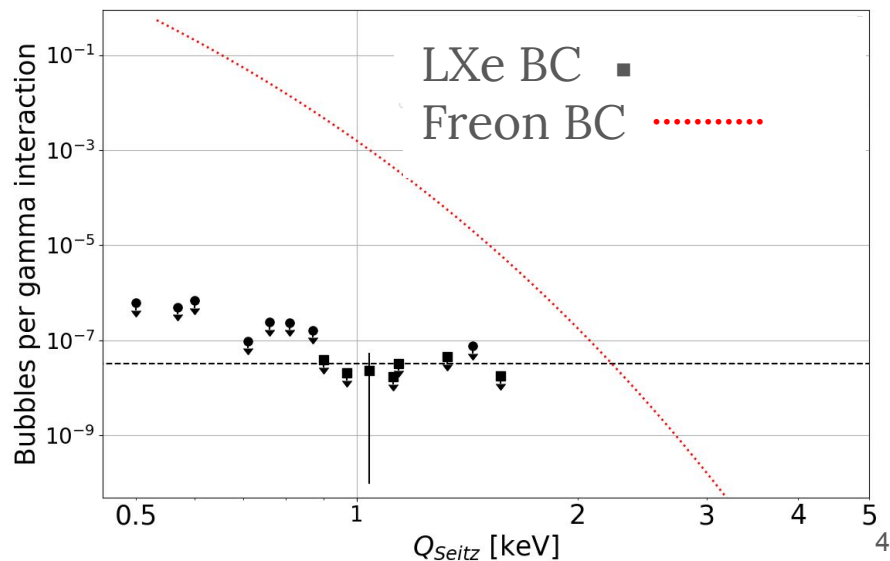
- Superheated liquids have efficient bubble formation at low n.r. thresholds
- Highly β/γ insensitive
- n.r./ α discrimination:



A liquid-noble bubble chamber

Adds in:

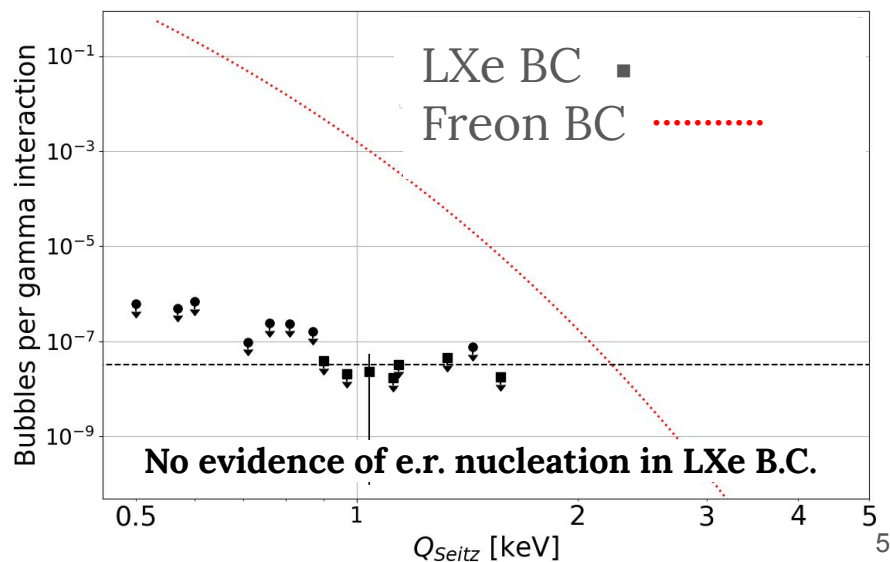
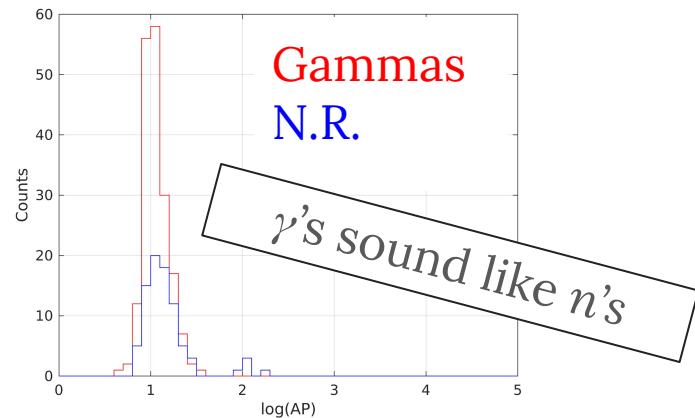
- Energy reconstruction
- Higher β/γ rejection than Freons
 - Lower threshold w/o e.r.
backgrounds increases
sensitivity to lower DM masses



A liquid-noble bubble chamber

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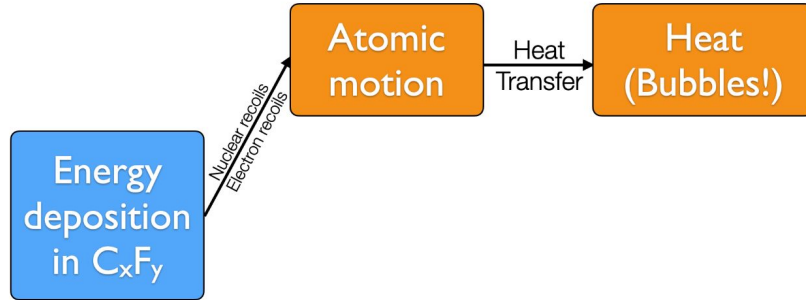
- Energy reconstruction
- Higher β/γ rejection than Freons*
 - Lower threshold w/o e.r. backgrounds increases sensitivity to lower DM masses



*Studied in liquid xenon: [PRD.111.032002](#)

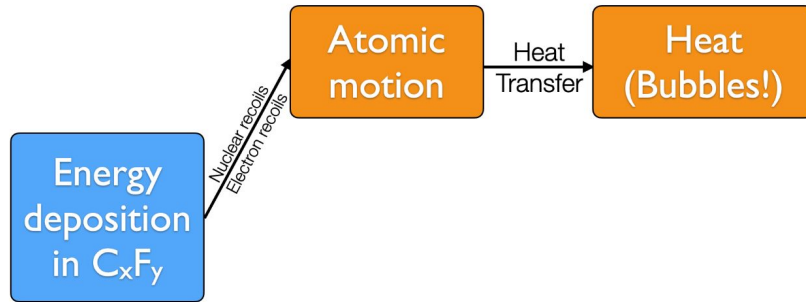
Energy deposition channels

Freon-based target fluid

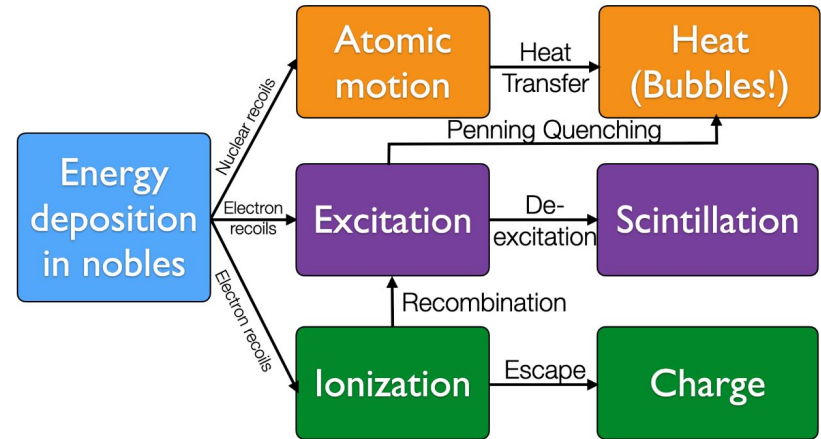


Energy deposition channels

Freon-based target fluid

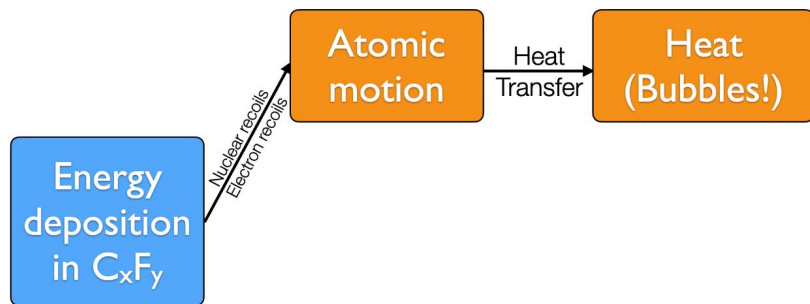


Liquid-noble target fluid

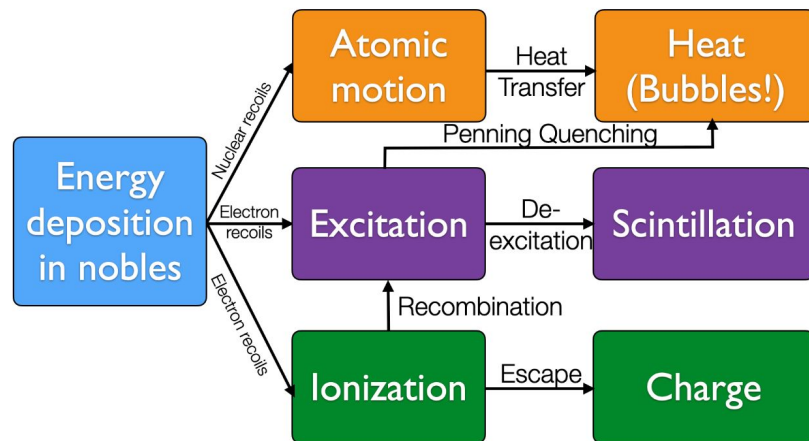


Energy deposition channels

Freon-based target fluid

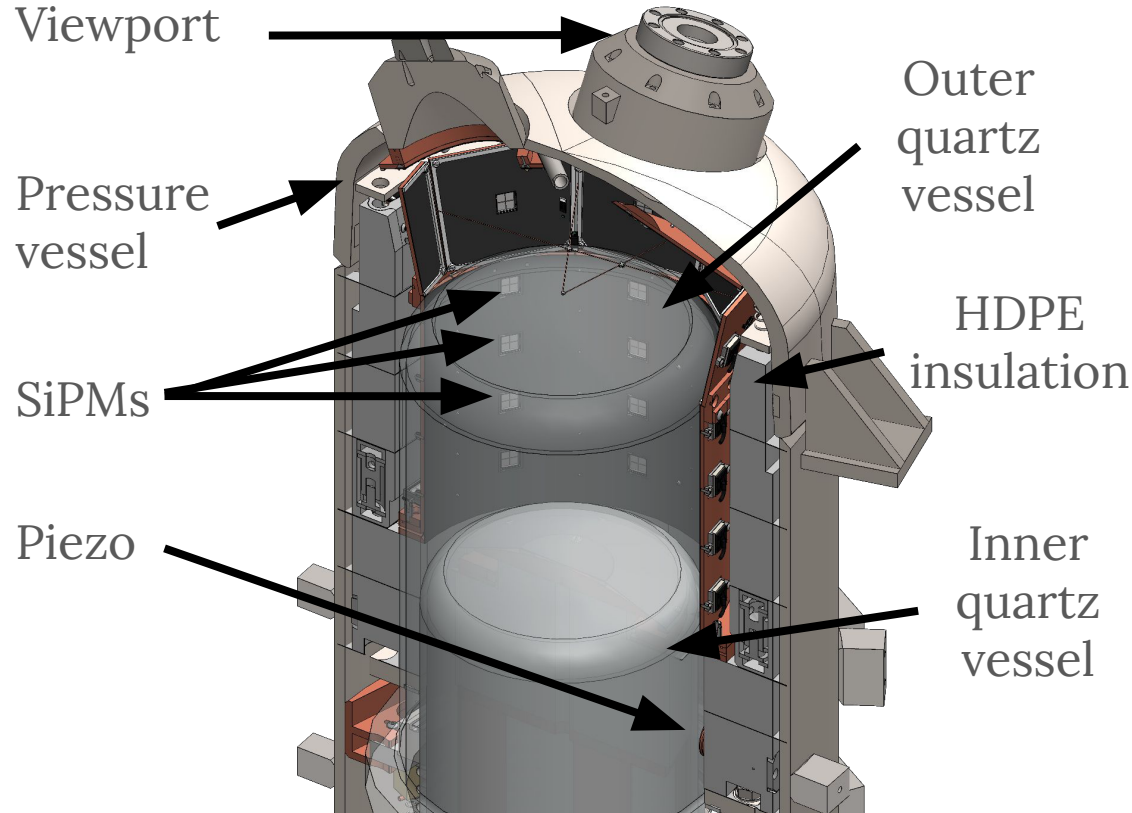


Liquid-noble target fluid



Main point: Liquid nobles remain β/γ insensitive in GeV-scale ROI
→ **sensitive to only nuclear recoils**

10 kg detector design

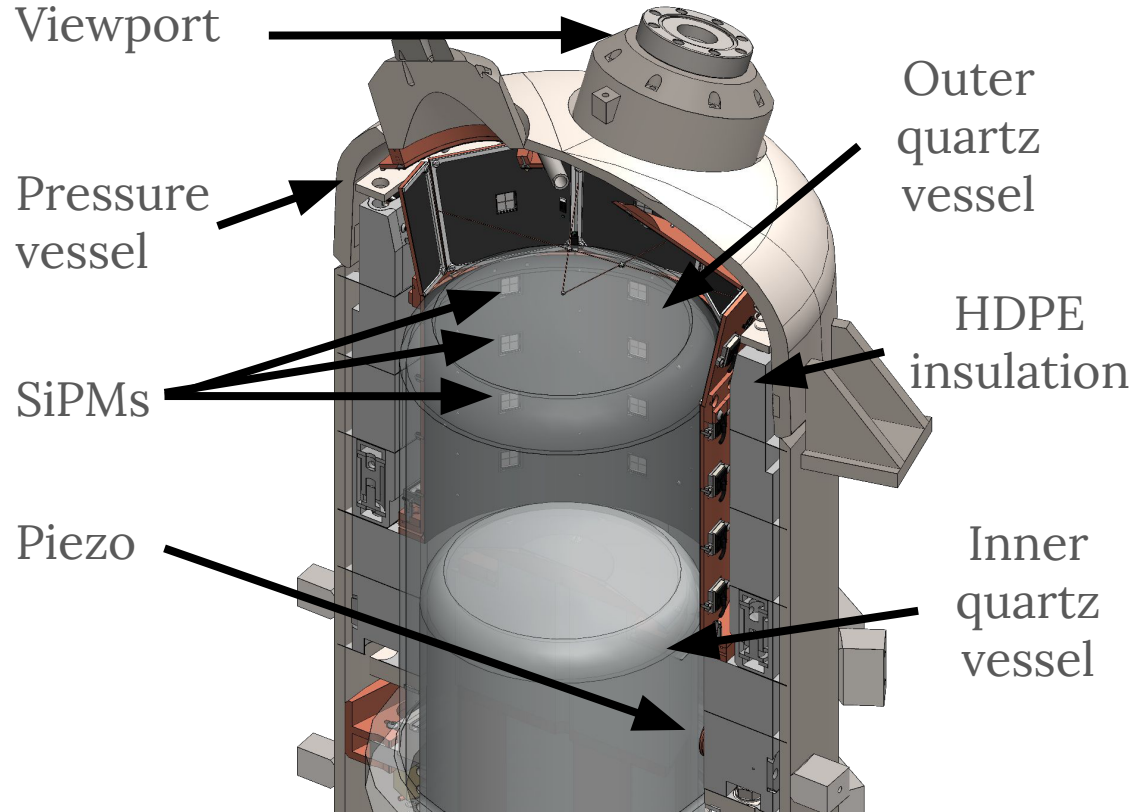


Two functionally identical detectors:

SBC-LAr10 @ Fermilab
(engineering, calibration,
future CE ν NS program)

SBC-SNOLAB @ SNOLAB
(low background
dark matter search)

10 kg detector design

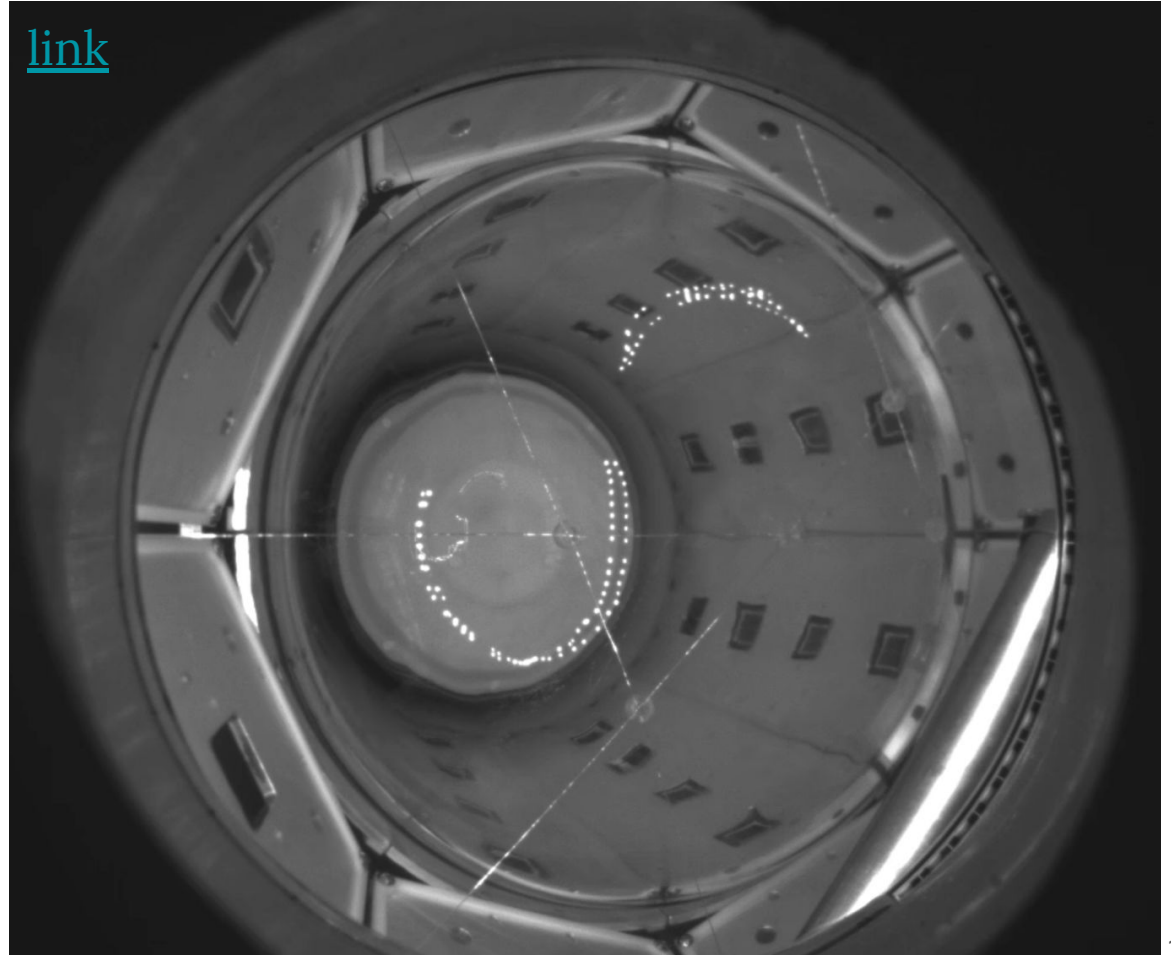


Design Goals	
Target Volume	10 L (10 kg LAr @ 130 K)
Nucleation Threshold	100 eV (30 psi , 130 K)
Thermodynamic Regulation	± 0.5 K, ± 0.1 bar, (± 5 eV Seitz threshold)
Scintillation Detection	$\sim 2\%$ collection, 1 photon/ 5 keV n.r.
Bubble Imaging	100 fps, mm resolution
Acoustic Reconstruction	Time-of-nucleation to ± 25 μ s

Operation at FNAL

[link](#)

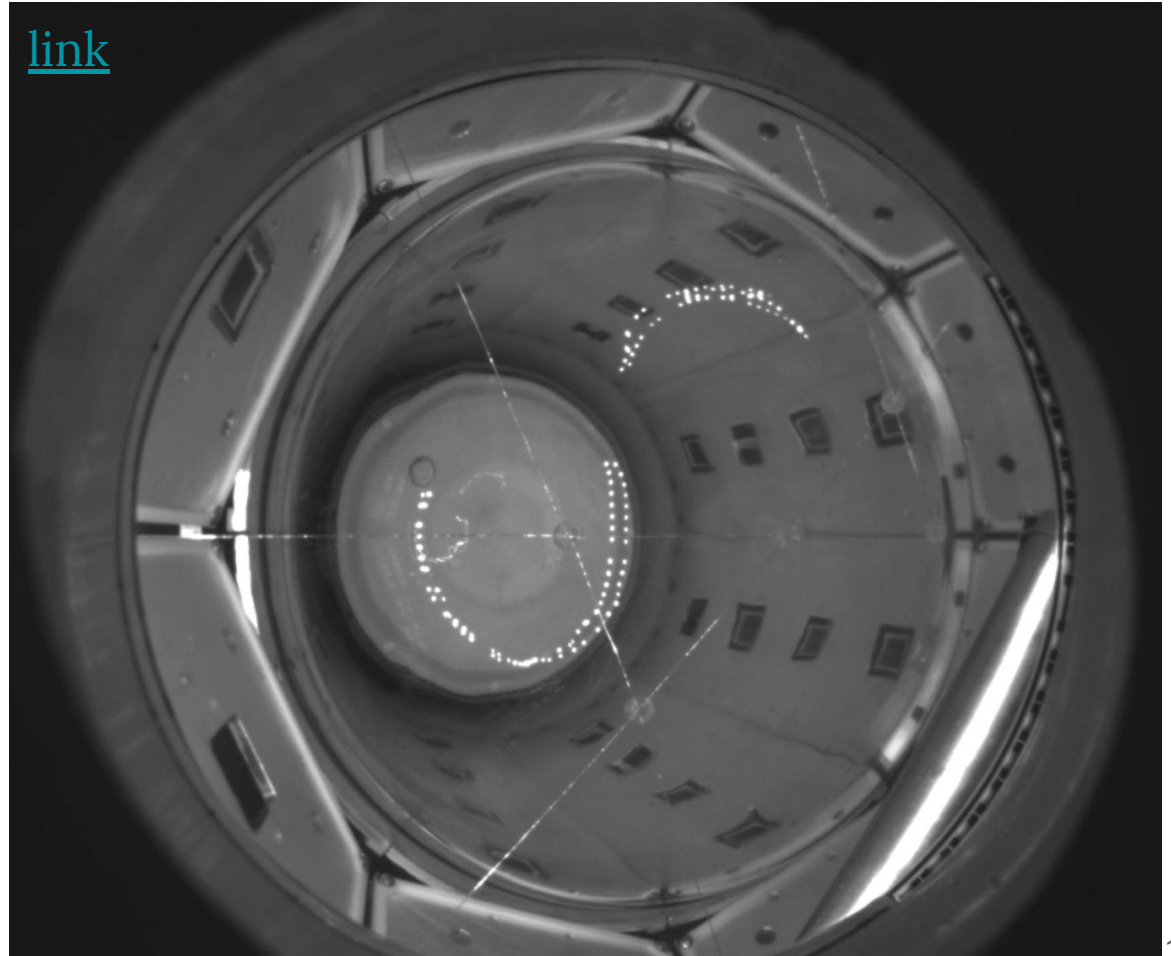
- 6 months of stable operation with nucleation thresholds from 0.46 – 3 keV
- Simultaneous measurement of scintillation and bubble nucleation observed



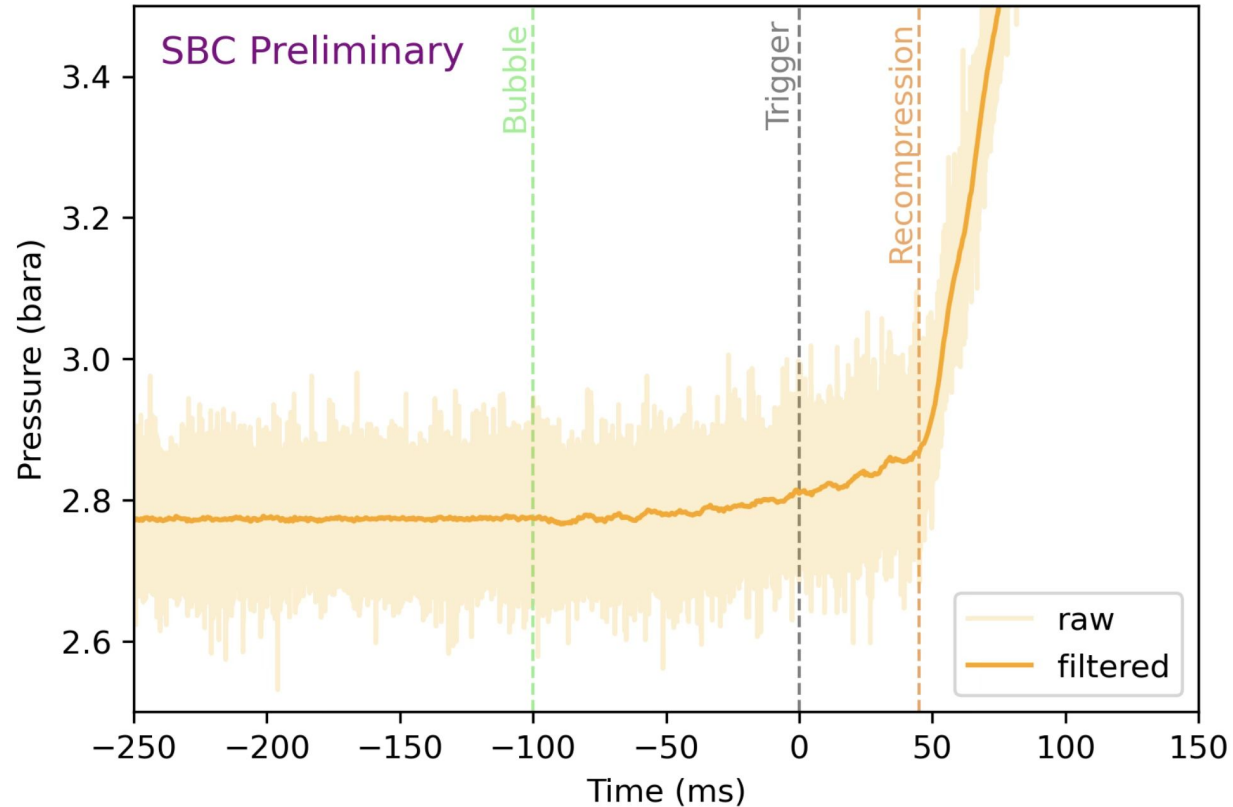
Operation at FNAL

[link](#)

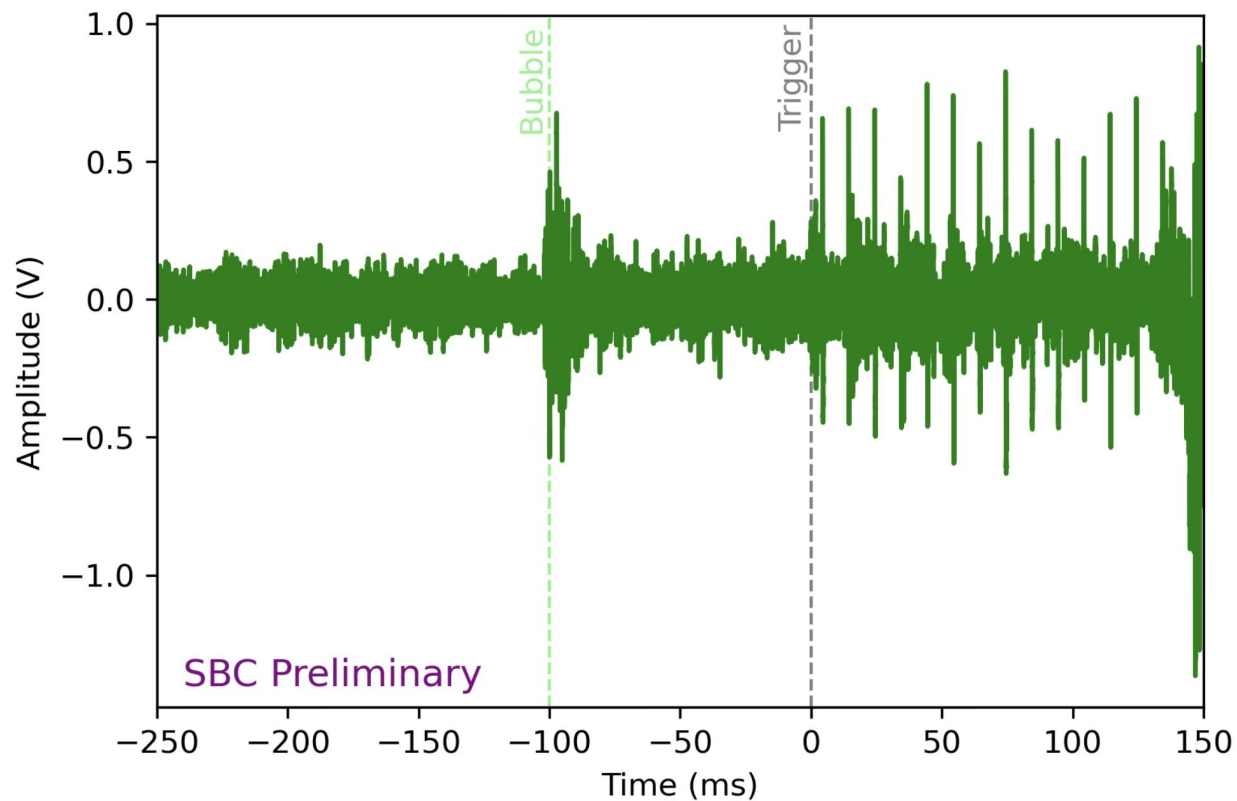
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Example pressure signal

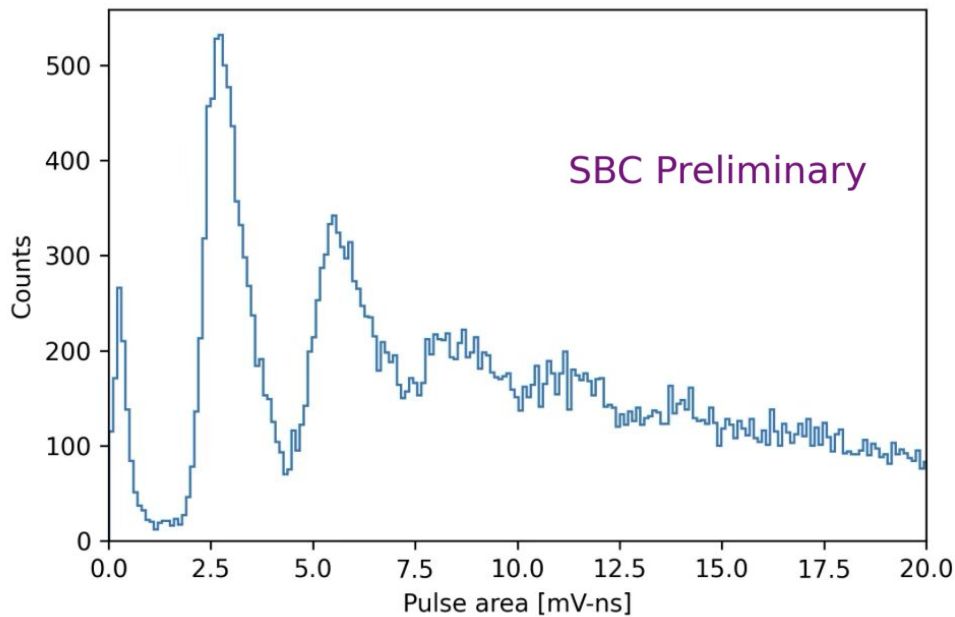
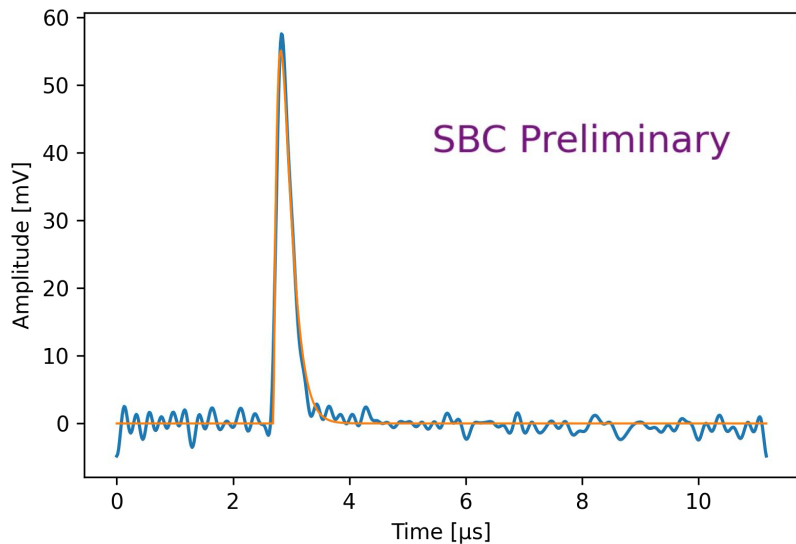


Example acoustic signal

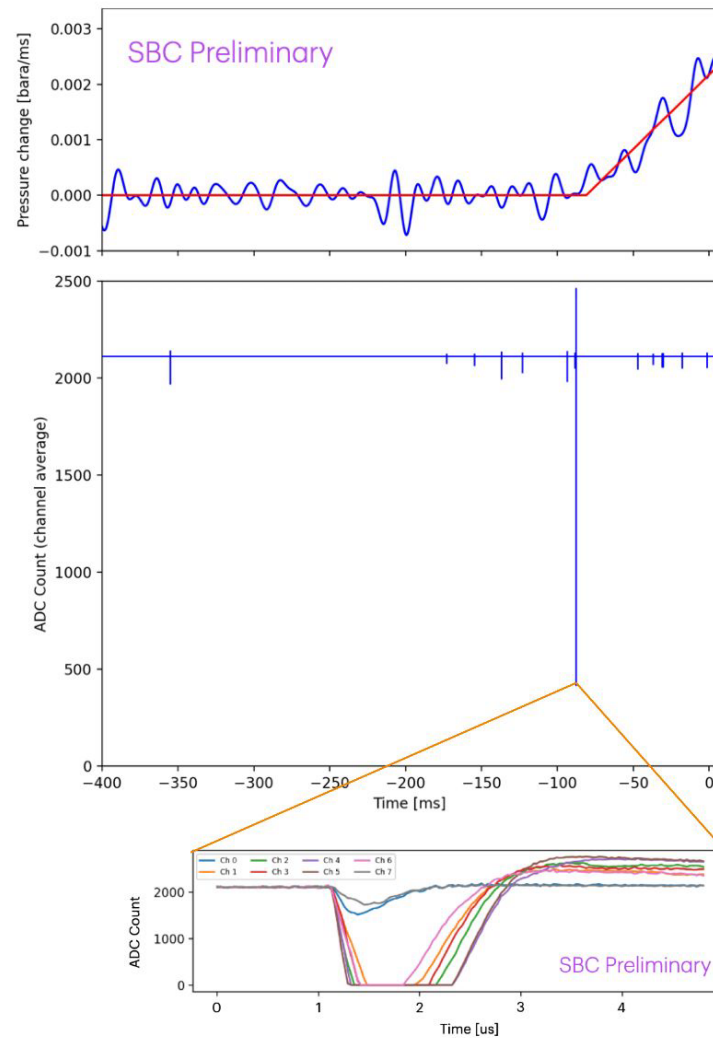


Example SiPM signal

- Pulse fitting to extract charge of the 32 SiPMs

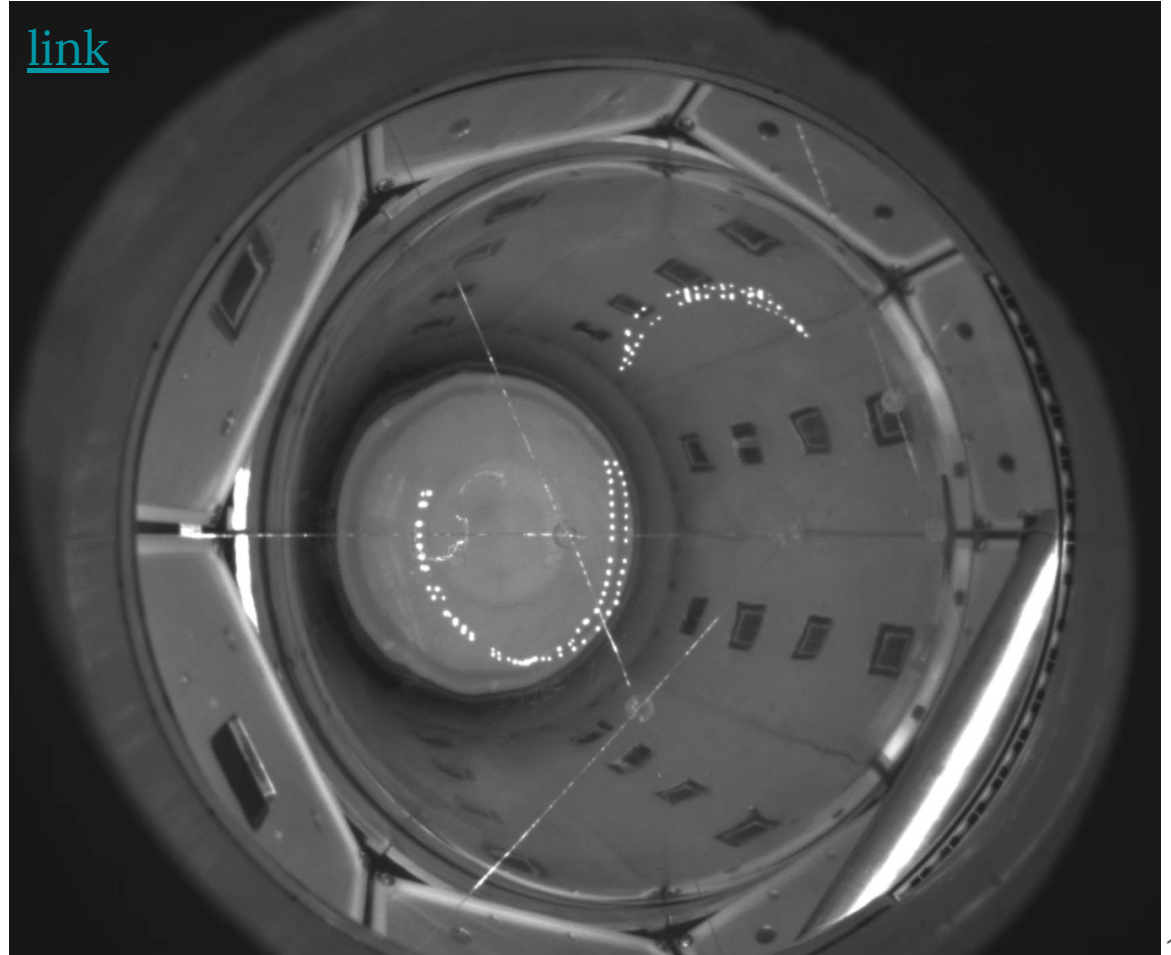


Signal coincidence: Pressure + bubble + scintillation



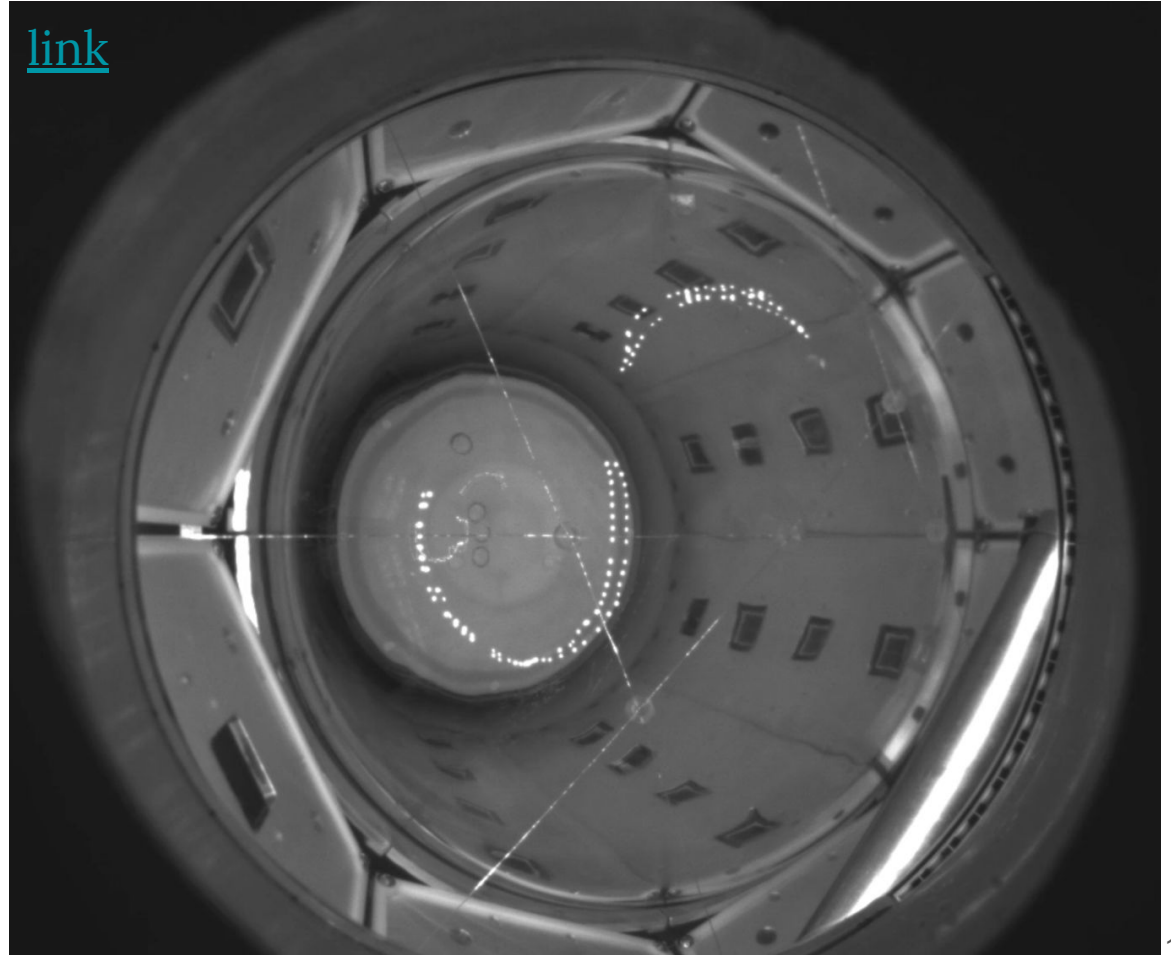
Multiscatter neutron event

[link](#)



Multiscatter neutron event

[link](#)



Component arrival at SNOLAB

Main detector components are arriving on surface at SNOLAB



Pressure vessel



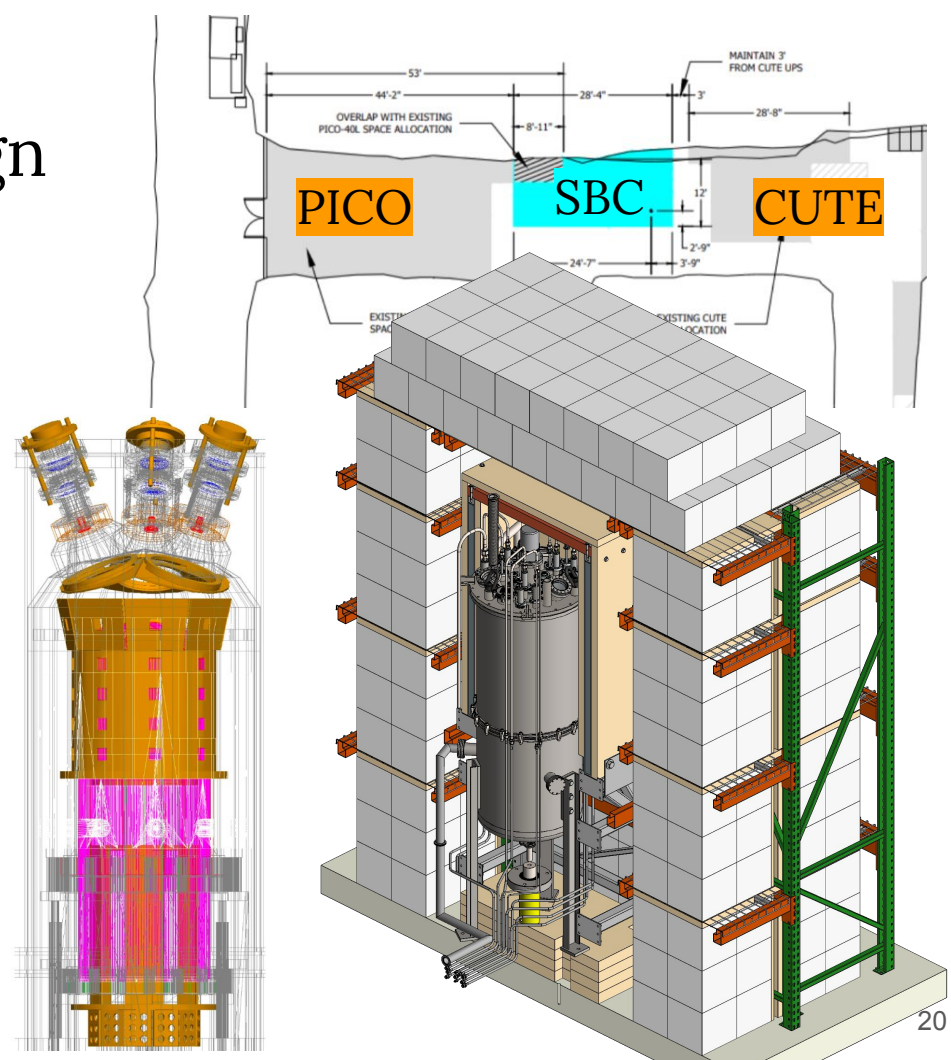
Vacuum jacket

Simulation and shield design

Space allocated in the ladder labs at SNOLAB

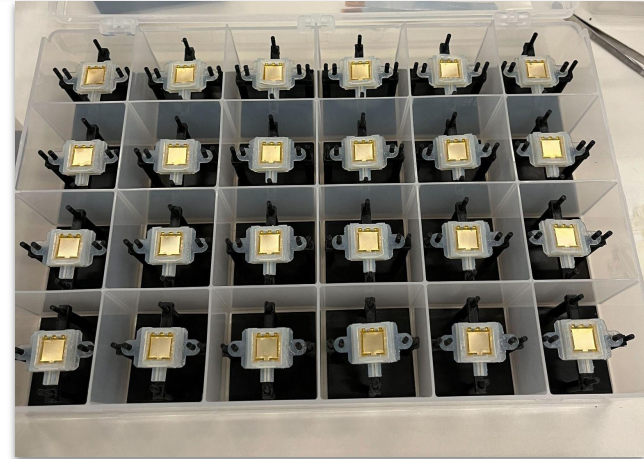
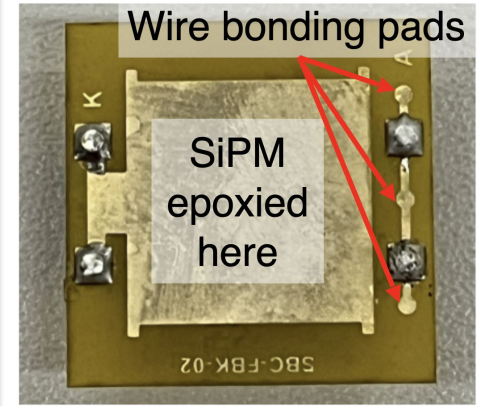
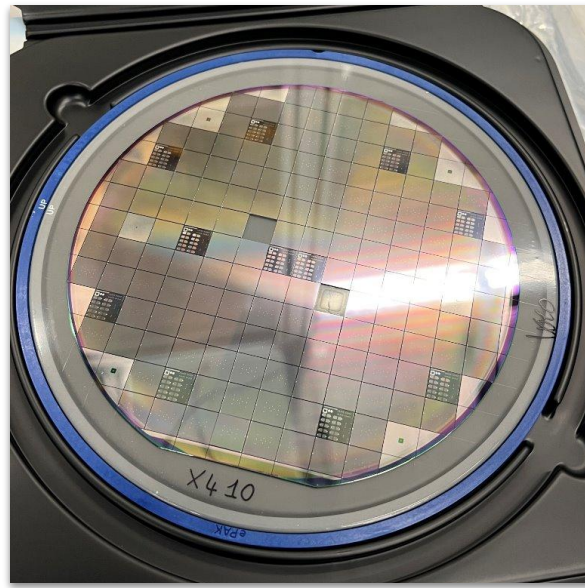
Full assay campaign and Geant4 simulation of internal and external backgrounds

- Water shield for fast neutrons
- Lead shield for gammas ($E_\gamma \gtrsim 1.4\text{MeV}$)

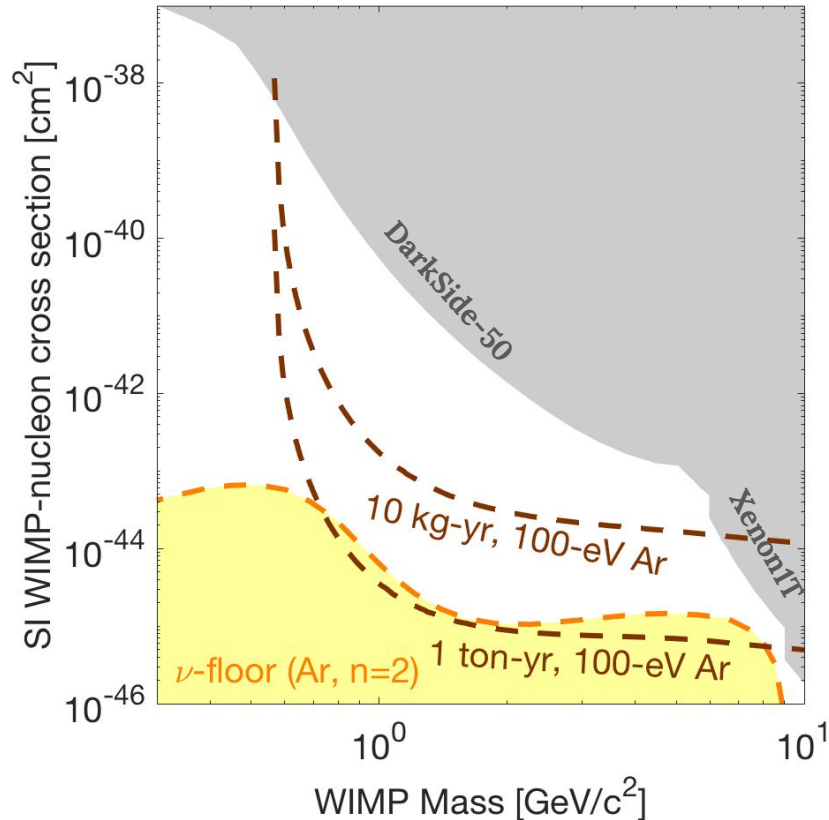


SiPM construction

- Simulations indicate high background rate from Hamamatsu VUV4 SiPMs
- Switch to FBK VUV HD3 (courtesy of nEXO)
- Wire bonding completed at TRIUMF, now characterizing at Queen's



Physics potential



- ROI: 0.1 keV - 10 keV n.r.
 - just bubble, no scintillation
- SNOLAB chamber, 10 kg-year exposure can reach 10^{-43} cm^2 @ $1 \text{ GeV}/c^2$
- Tonne-year can reach boundary of the Ar ν -fog
- Target fluid flexibility (Xe, Ar, N_2 , etc.)

Conclusion

- Exciting time for SBC-SNOLAB
 - Gateway 2 just passed and parts arriving on surface for assembly this fall

- Publications:

[arXiv: 2207.12400](https://arxiv.org/abs/2207.12400)

[Universe 9 \(2023\) 8, 346](#)

[Phys.Rev.D 111 \(2025\) 3, 032002](#)

[JINST 19 \(2024\) 08, T08003](#)

[JINST 21 \(2026\) 03, P03020](#)



UC Santa Barbara



