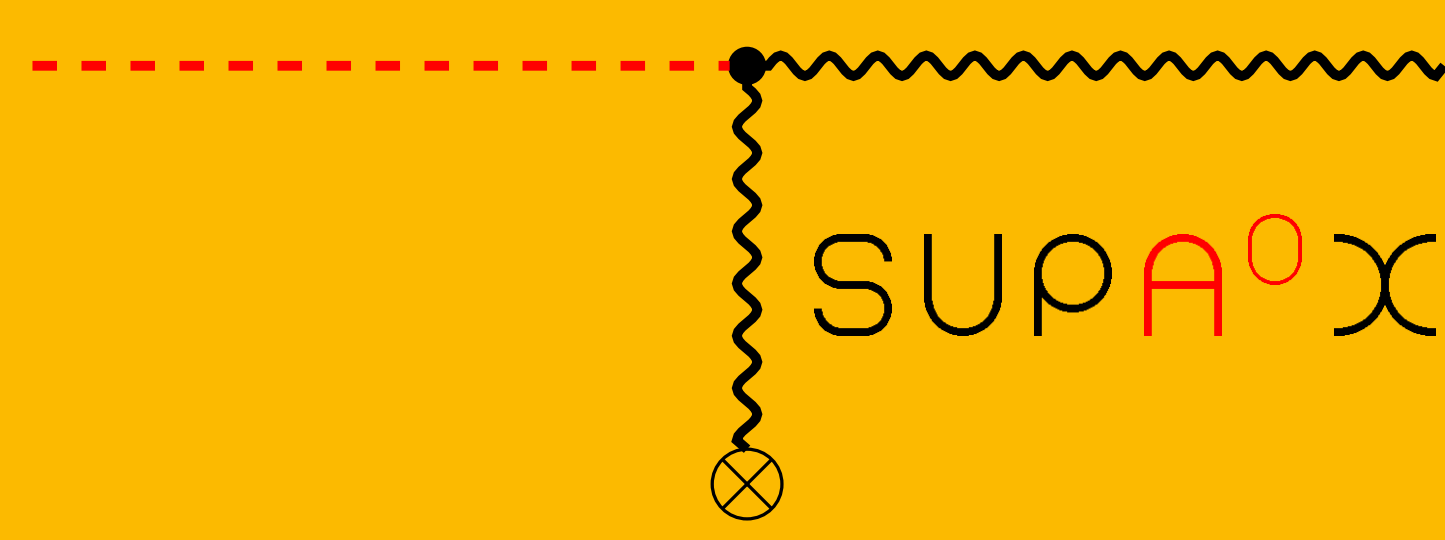


Supax — a superconducting axion search experiment

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Introduction — Supax

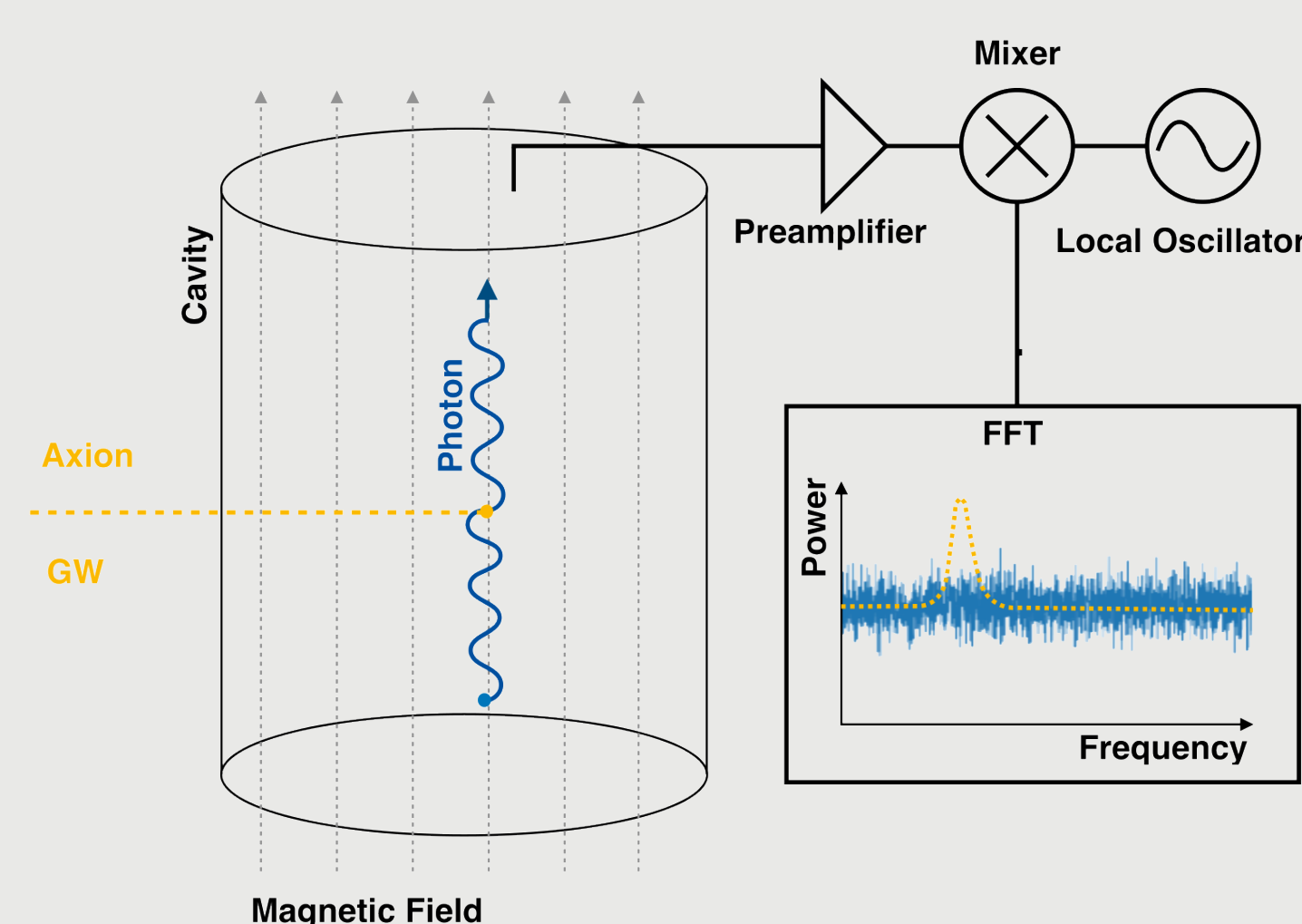
Axions are hypothetical fundamental particles that could explain *smallness of electric dipole moment* of the neutron, and is a **dark matter candidate**. Axions are expected to be low-mass particles, for which **haloscope** is a well matched approach. Supax aims at axions detection using **Primakoff effect**: $\text{axion} \rightarrow \gamma$ conversion in static magnetic field.

$$P_{\text{sig}} = |g_{a\gamma\gamma}|^2 \rho_a \frac{\beta}{\beta + 1} \frac{1}{m_A} B_0^2 V C Q_{\text{eff}} \eta \hbar^2 c^5 \varepsilon_0$$

Nearly identical experiment design can be used for **high frequency gravitational waves** (HFGW) detection via **inverse Gertsenshtein effect**.

Detection approach

- ▶ cryostat 10 mK
- ▶ magnet 12 T
- ▶ low-noise amplifier (TWPA, JPA)
- ▶ frequency $\mathcal{O}(1 \text{ GHz})$



Supax — Results

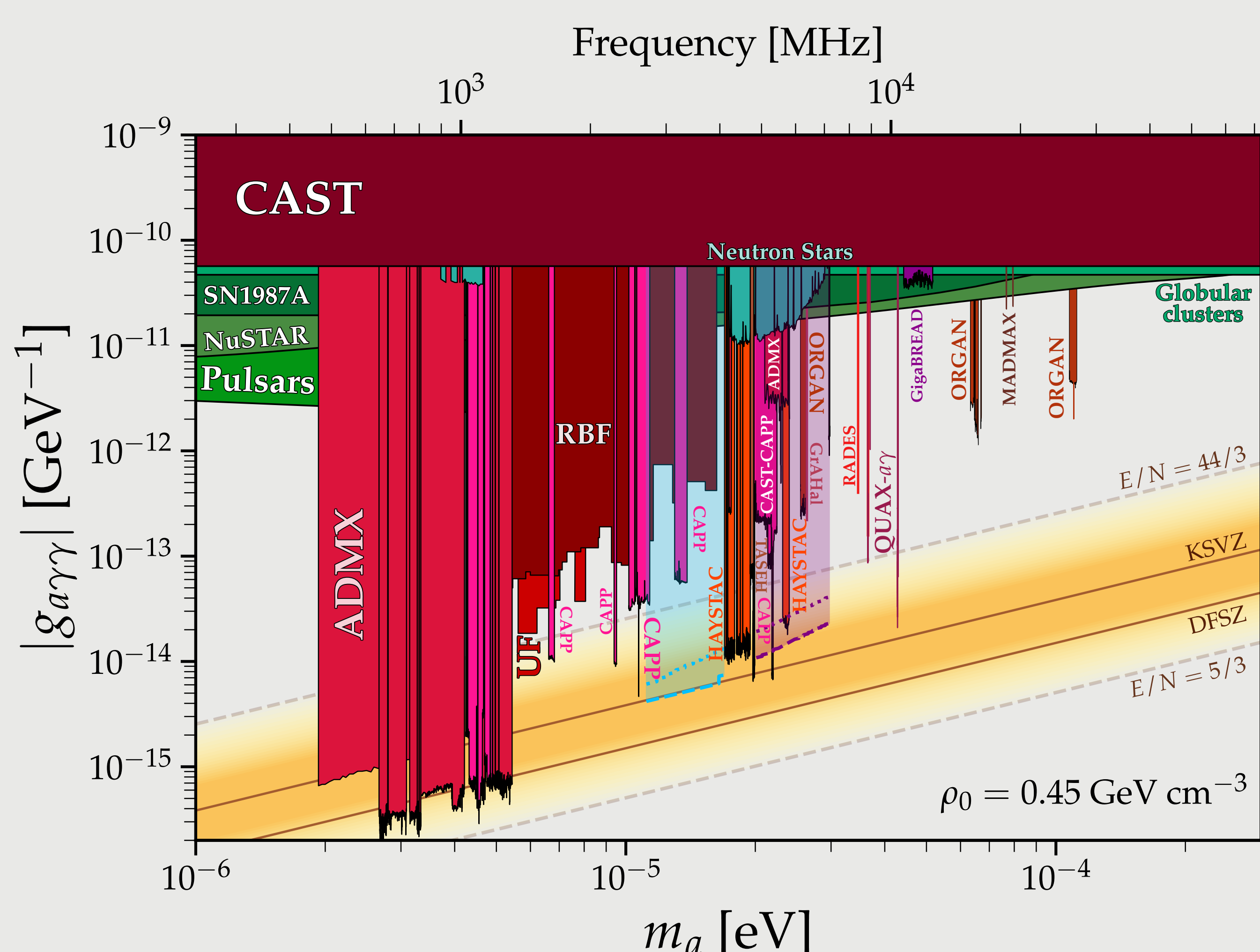
Prototype:

- ▶ resonant frequency 8.47 GHz $\rightarrow m_A \approx 34 \mu\text{eV}$
- ▶ $T_{\text{cav}} = 2 \text{ K}$, $B_0 = 12.1 \text{ T}$, $Q_0 = 50,000$
- ▶ mass range: $0.0006 \mu\text{eV}$ (1.4 MHz) around $34 \mu\text{eV}$
- ▶ **axion-photon coupling larger than $|g_{a\gamma\gamma}| > 1.6 \times 10^{-13}$ excluded at 95% confidence level**

Supax — Outlook

Assumptions:

- ▶ $2.7 \text{ GHz} < f_0 < 7.2 \text{ GHz} \rightarrow 11 \mu\text{eV} < m_A < 30 \mu\text{eV}$
- ▶ $Q_0 = 45,000$, $T_{\text{cav}} = 10 \text{ mK}$, $t_{\text{meas}} = 40 \text{ weeks}$



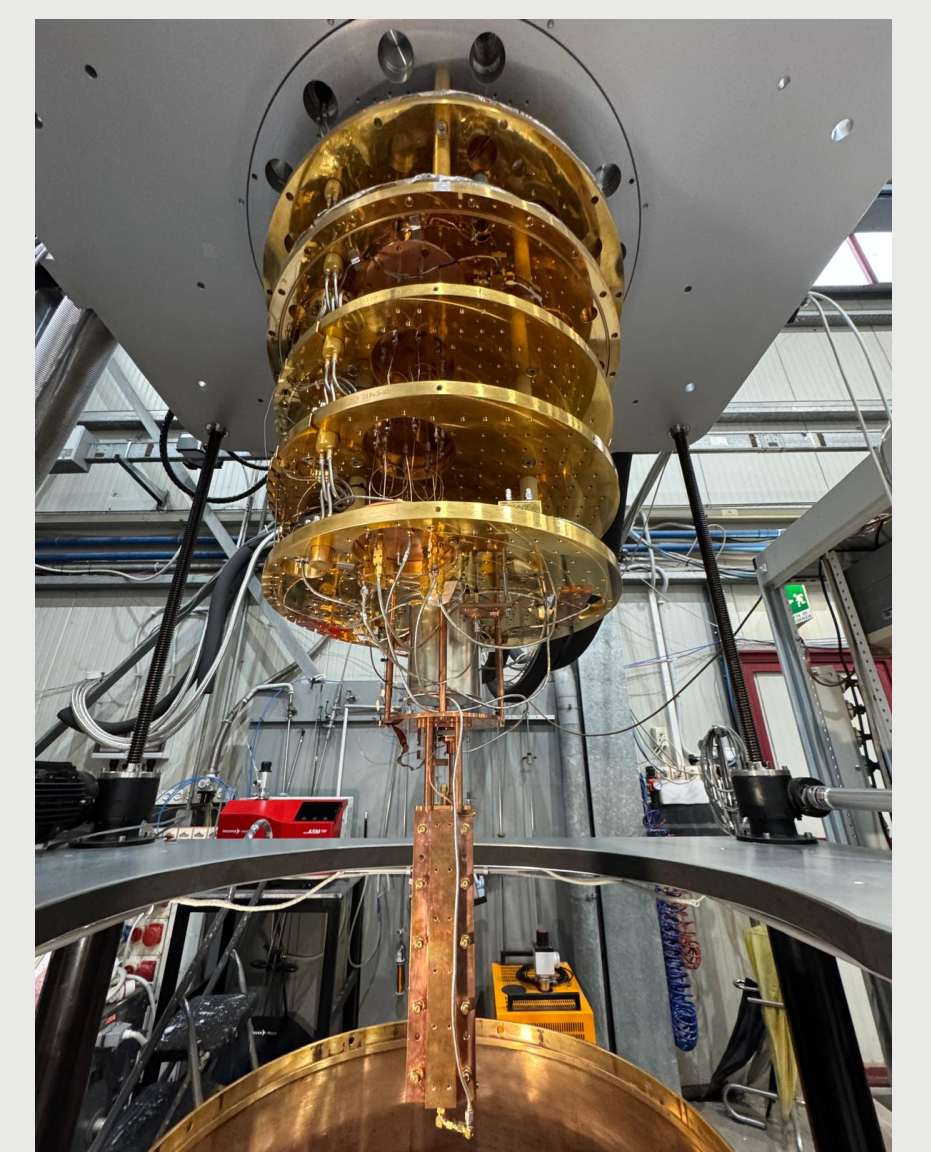
GW search — GravNet

- ▶ extension of axion search experiment
- ▶ focus on inspiral phase of **primordial black holes** (PBH) merger
- ▶ no constraints on the fraction of DM composed of PBHs in mass range:
 $10^{-16} - 10^{-10} M_{\odot}$, masses within sensitivity range
- ▶ **network** of $\mathcal{O}(1 \text{ GHz})$ cavities + 250 MHz cavity
- ▶ signal detection in network — suppressed background event count
- ▶ EM signal **power**:

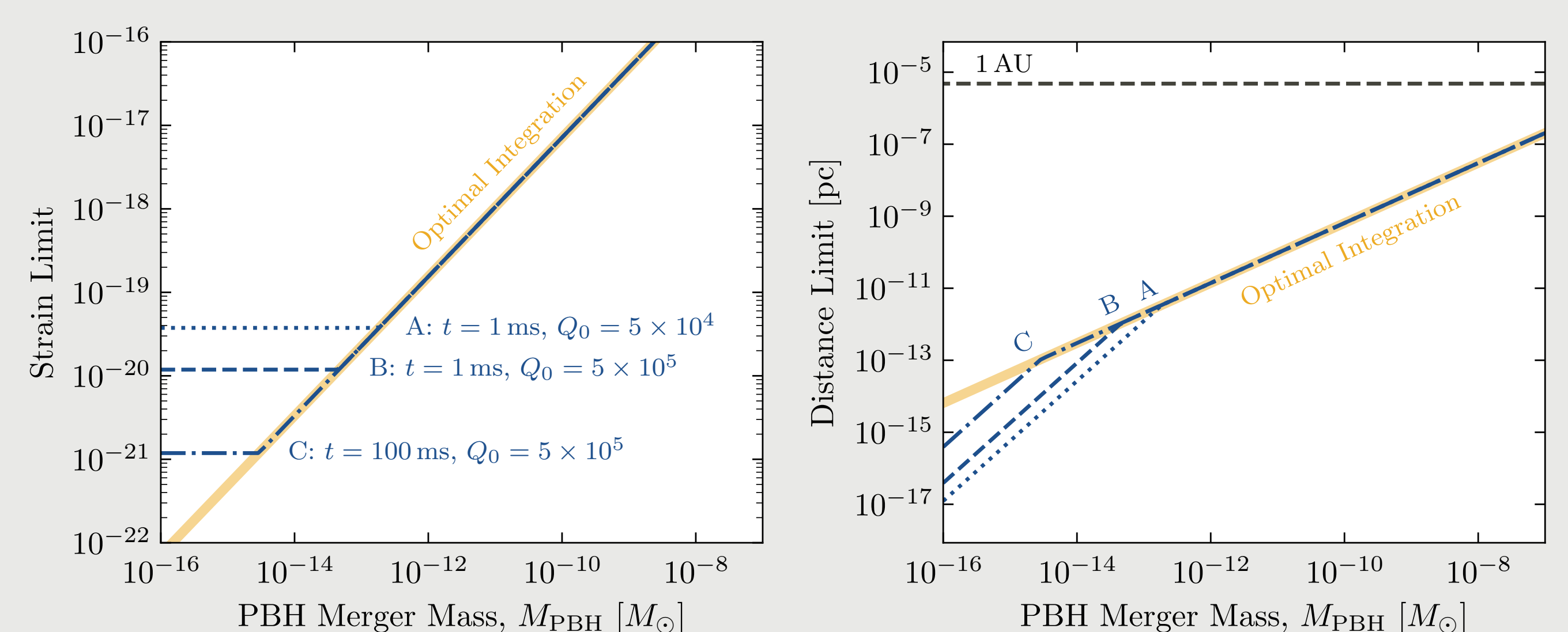
$$P_{\text{sig}} \simeq \omega_g h_{\times,+}^2 \langle B_0^2 \rangle Q_l C_{\text{GW}} V$$

Experimental setup

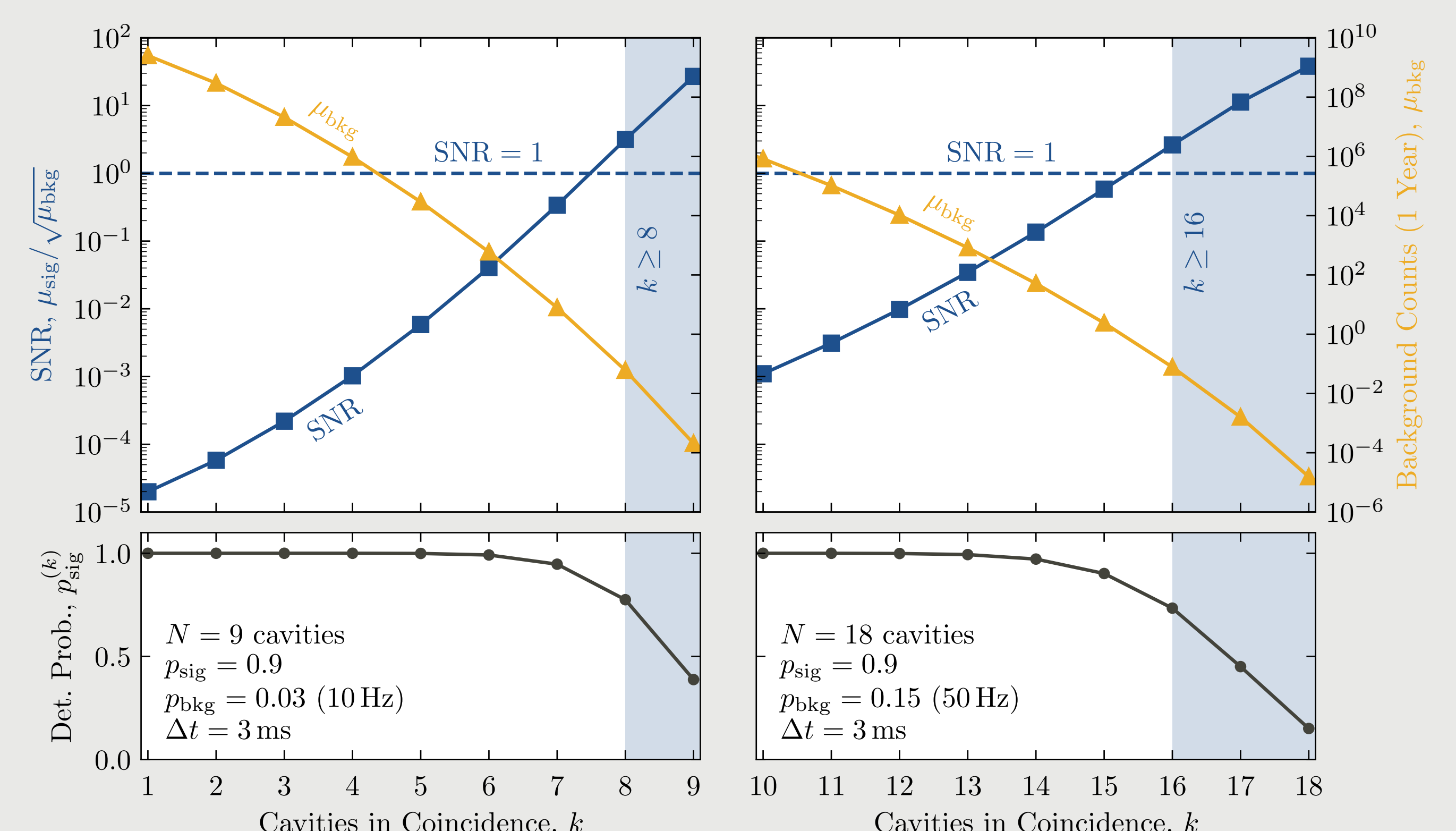
- ▶ **commercial cryostat**, same as for axion search
- ▶ **tunable cavity**, $\Delta f_0 \approx 60 \text{ MHz}$ (sapphire tuning rod)
- ▶ **quantum limited noise readout line** (TWPA amplifier)
- ▶ dipole antenna for readout



Sensitivity + one photon distance limit



Network performance



References

- ▶ Supax CDR: [arXiv:2505.07541](https://arxiv.org/abs/2505.07541)
- ▶ GravNet CDR: [arXiv:2603.24645](https://arxiv.org/abs/2603.24645)
- ▶ GravNet Collaboration website: www.pi.uni-bonn.de/gravnet



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